Some Theoretical Aspects of the Schwinger Mechanism

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Outline

$$\mathcal{L} = \overline{\psi} (i \partial \!\!\!/ - \mathfrak{m} + \!\!\!/ \mathcal{K}_{ext}) \psi$$

- A^{μ} : dynamical photon field (responsible for radiation, arises in loops)
- + ${\cal A}^{\,\mu}_{\rm ext}$: external photon field (controlled by experimental setup, not quantized)

- The Fourier components of a time independent external field all have zero frequency
- By momentum conservation, it seems that this field cannot produce anything that has an energy > 0 (e.g., an e^-e^+ pair has a minimal energy 2m)
- This is definitely correct in perturbation theory, i.e., for contributions that one may obtain from an expansion in powers of the external field
- There are also contributions in $\exp(-m^2/eE_{\rm ext})$, non-analytic in the external field (at least for a constant $E_{\rm ext}$), for which perturbation theory has nothing to say \rightarrow Schwinger effect [Schwinger, 1953]

Schwinger effect \sim tunneling



$$\mathfrak{P}_{\rm pair \ prod} \sim \exp(-\pi m^2/eE)$$

- All Taylor coefficients about e = 0 are zero (when E is not constant but has only Fourier modes of very low frequency, the dominant Taylor coefficients are at a very large order)
- Critical field: $E_c \equiv m^2/e$ (extremely large even for the lightest known charged particle, the electron)
- Numerically more important in QCD, since the strong coupling is much larger

OUTLINE

- Part I:
 - Quantum Fields coupled to (strong) external sources
 - Correlations in the Schwinger effect
 - Bogoliubov transformation
- Part II:
 - Numerical evaluation on the lattice
 - Worldline formalism
 - Dynamically assisted Schwinger effect

Quantum Fields coupled to external sources

[FG, Lappi, Venugopalan, 2006-2008]

$$\mathcal{L} \equiv \frac{1}{2} (\vartheta_{\mu} \varphi) (\vartheta^{\mu} \varphi) - \frac{1}{2} m^2 \varphi^2 - \frac{g^2}{4!} \varphi^4 + j \varphi$$

g is dimensionless in 4 dimensions, like the QCD coupling

$$\mathrm{Order}(graph) = g^{-2} g^{N_{\mathrm{final particles}}} g^{2N_{\mathrm{loops}}}(gj)^{N_{\mathrm{sources}}}$$

- Independent of $N_{\rm sources}$ if $gj \sim 1 ~~ \rightarrow ~$ strong source regime
- Still perturbative in g², but not in j
- Vacuum graphs ($N_{\rm final\ particles}=0)\sim g^{-2}$ with strong sources

VACUUM DIAGRAMS

- When $j \equiv 0$: $|0_{out}\rangle = U(+\infty, -\infty) |0_{in}\rangle$ with U unitary, The vacuum evolves into the vacuum with probability one, $\langle 0_{out} | 0_{in} \rangle = e^{iV}$ with $V \in \mathbb{R}$, Vacuum graphs are purely real; their sum is a pure phase.
- When $j \neq$ 0, unitarity tells this is no longer true,

$$1 = \sum_{\alpha} P(\alpha) = P(\text{vacuum}) + \sum_{\alpha \neq \text{vacuum}} P(\alpha),$$

If $P(\alpha) \neq 0$ for some non-empty final state, then P(vacuum) < 1, $P(vacuum) = e^{-2 \operatorname{Im} V}$, therefore vacuum graphs are complex, Their sum is not a phase, and one cannot disregard them.

EXAMPLE: CONTRIBUTION TO P(11 **PARTICLES**)



- Right: amplitude; Left: conjugate amplitude (dots: source j)
- Vacuum graph $\sim g^{-2}$ in the strong source regime $~\rightarrow~$ hopeless?

EXCLUSIVE VS INCLUSIVE

• Probability to reach a specific final state:

$$\frac{\mathrm{d} \mathsf{P}_{\mathsf{n}}}{\mathrm{d}^{3} \mathsf{p}_{1} \cdots \mathrm{d}^{3} \mathsf{p}_{\mathsf{n}}} = \frac{1}{\mathsf{n}!} \frac{1}{(2\pi)^{3} 2\mathsf{E}_{\mathsf{p}_{1}}} \cdots \frac{1}{(2\pi)^{3} 2\mathsf{E}_{\mathsf{p}_{\mathsf{n}}}} \left| \left\langle \mathsf{p}_{1} \cdots \mathsf{p}_{\mathsf{n}^{\mathrm{out}}} \middle| \mathsf{0}_{\mathrm{in}} \right\rangle \right|^{2}$$

- Vacuum graphs do not cancel
- Each of them is exponentially suppressed (tiny probability to reach a given final state)
- Inclusive particle spectrum:

$$\frac{\mathrm{dN}_1}{\mathrm{d}^3\mathbf{p}} \equiv \sum_{n=0}^{\infty} (n+1) \int \mathrm{d}^3\mathbf{p}_1 \cdots \mathrm{d}^3\mathbf{p}_n \ \frac{\mathrm{dP}_{n+1}}{\mathrm{d}^3\mathbf{p}\mathrm{d}^3\mathbf{p}_1 \cdots \mathrm{d}^3\mathbf{p}_n}$$

- Vacuum graphs DO cancel
- The moments have a well defined series expansion in g
- At LO in g, can be expressed in terms of retarded classical fields

COMBINATORICS OF MULTIPARTICLE PRODUCTION

The probability of producing n particles can always be parameterized as

$$P_n = e^{-\alpha} \sum_{p=1}^n \frac{1}{p!} \sum_{r_1 + \dots + r_p = n} b_{r_1} \cdots b_{r_p}$$

- p = number of clusters (sets of correlated particles)
- a = mean number of clusters
- $b_r = mean number of clusters with r particles$
- Unitarity: $a = b_1 + b_2 + \cdots$
- Vacuum persistence probability: $P_0 = e^{-\alpha}$

• Moments:
$$\langle n \rangle = \sum_{r} r b_{r}$$
, $\langle n^{2} \rangle - \langle n \rangle^{2} = \sum_{r} r^{2} b_{r}$, etc...

- In general, $P_0 \neq e^{-\langle n \rangle}$ (only equal if all particles are produced uncorrelated, i.e., if there are no clusters of size > 1)
- $a, b_r \sim g^{-2}$ in the strong source regime

$$\begin{split} \left\langle 0_{\mathrm{in}} \left| 0 \right| 0_{\mathrm{in}} \right\rangle &= \underbrace{\sum_{\text{states } \alpha} \left\langle \alpha_{\mathrm{out}} \left| 0 \right| \alpha_{\mathrm{out}} \right\rangle}_{\text{mixed } - +} \underbrace{\left\langle 0_{\mathrm{in}} \right| \alpha_{\mathrm{out}} \right\rangle}_{-\text{ sector } (\text{conjugate rules})} \underbrace{\left\langle \alpha_{\mathrm{out}} \right| 0_{\mathrm{in}} \right\rangle}_{+\text{ sector } (\text{normal rules})} \\ G_{++}^{0}(p) &= \frac{i}{p^{2} - m^{2} + i\varepsilon}, \qquad G_{--}^{0}(p) = \frac{-i}{p^{2} - m^{2} - i\varepsilon} \\ G_{+-}^{0}(p) &= 2\pi\theta(-p^{0})\delta(p^{2} - m^{2}), \qquad G_{-+}^{0}(p) = 2\pi\theta(+p^{0})\delta(p^{2} - m^{2}) \end{split}$$

CONNECTION TO FEYNMAN GRAPHS



Note: vacuum graphs are all zero in SK formalism without external sources, but non-zero if $j \neq \mathbf{0}$

CONNECTION TO FEYNMAN GRAPHS (CONT.)



$$\frac{dN_1}{d^3p} \propto \bigcap_{p} + \bigcap_{p$$

- The gray blobs are the SK 1-point ($\phi_{\pm}(x)$) and 2-point ($g_{-+}(x,y)$) connected correlation functions
- This formula is exact (to all orders in g and j)
- Strong source regime: $\phi_{\pm} \sim g^{-1}$, $\mathfrak{G}_{-+}(x,y) \sim 1$,
 - $\rightarrow~$ The first term dominates.

INCLUSIVE SPECTRUM AT LEADING ORDER

At tree level:

$$\begin{split} (\Box+m^2)\phi_\pm+\frac{g^2}{6}\phi_\pm^3=\mathfrak{j},\\ \lim_{x^0\to-\infty}\phi_\pm(x)=0 \end{split}$$

- Classical EOM with retarded boundary conditions
 - \rightarrow numerically straightforward
- + ϕ_\pm are equal, and real valued
- Given the Fourier decomposition of these classical fields

$$\varphi(\mathbf{y}) \equiv \int \frac{\mathrm{d}^{3}\mathbf{p}}{(2\pi)^{3}2\mathsf{E}_{\mathbf{p}}} \left[\mathsf{f}(\mathbf{y}^{0},\mathbf{p}) \, e^{-\mathrm{i}\mathbf{p}\cdot\mathbf{y}} + \mathsf{c.c.} \right],$$

the LO spectrum reads

$$\left. \frac{\mathrm{d}N_1}{\mathrm{d}^3 \mathbf{p}} \right|_{\mathrm{LO}} = \frac{1}{(2\pi)^3 2 \mathsf{E}_{\mathbf{p}}} |\mathsf{f}(+\infty,\mathbf{p})|^2$$

• Note: at this order, only particles that couple directly to the sources can be produced (e.g., gluons in QCD)

LO + NLO







- LO: can only produce particles that couple directly to the sources
- Static sources: the LO gives zero (same for the first of the NLO terms)
- NLO: a different particle may run in the loop \rightarrow can produce particles that do not couple directly to the source (e.g., electrons, quarks)
- Second NLO graph: contains a non-analytic contribution when the source is static

Simplest example:

$$\frac{\mathrm{d}\mathsf{P}_1}{\mathrm{d}^3\mathbf{p}} = \mathrm{e}^{-\alpha} \, \mathrm{b}_1(\mathbf{p})$$

- At LO: $a,b_1(p)$ given by classical fields with $\mbox{non retarded}$ boundary conditions
- Although not impossible in principle, very hard in practice

Correlations in the Schwinger effect [Fukushima, FG, Lappi, 2009]



• Consider scalar QED for simplicity:

$$\mathcal{L} \equiv (D_{\mu}\varphi)(D^{\mu}\varphi)^{*} - \mathfrak{m}^{2}\varphi\varphi^{*} ,$$

• Generating functional:

$$\begin{aligned} \mathcal{F}[z,\overline{z}] \equiv &\sum_{m,n=0}^{\infty} \frac{1}{m!n!} \int \prod_{i=1}^{m} \frac{d^{3}p_{i}}{(2\pi)^{3}2\mathsf{E}_{p_{i}}} \, z(p_{i}) \prod_{j=1}^{n} \frac{d^{3}q_{j}}{(2\pi)^{3}2\mathsf{E}_{q_{j}}} \, \overline{z}(q_{j}) \\ & \times \left| \left\langle \underbrace{p_{1}\cdots p_{m}}_{\text{particles}} \underbrace{q_{1}\cdots q_{n}}_{\text{out}} | \mathbf{0}_{\text{in}} \right\rangle \right|^{2} \end{aligned}$$

$$\frac{\mathrm{dN}_{1}^{+}}{\mathrm{d}^{3}\mathbf{p}} = \frac{\delta\mathcal{F}[z,\overline{z}]}{\delta z(\mathbf{p})}\Big|_{z=\overline{z}=1}, \qquad \frac{\mathrm{dN}_{1}^{-}}{\mathrm{d}^{3}\mathbf{q}} = \frac{\delta\mathcal{F}[z,\overline{z}]}{\delta\overline{z}(\mathbf{q})}\Big|_{z=\overline{z}=1}$$
$$\frac{\mathrm{dN}_{1}^{++}}{\mathrm{d}^{3}\mathbf{p}_{1}\mathrm{d}^{3}\mathbf{p}_{2}} = \frac{\delta^{2}\mathcal{F}[z,\overline{z}]}{\delta z(\mathbf{p}_{1})\delta z(\mathbf{p}_{2}))}\Big|_{z=\overline{z}=1}$$
$$\frac{\mathrm{dN}_{1}^{--}}{\mathrm{d}^{3}\mathbf{q}_{1}\mathrm{d}^{3}\mathbf{q}_{2}} = \frac{\delta^{2}\mathcal{F}[z,\overline{z}]}{\delta\overline{z}(\mathbf{q}_{1})\delta\overline{z}(\mathbf{q}_{2}))}\Big|_{z=\overline{z}=1}$$
$$\frac{\mathrm{dN}_{1}^{+-}}{\mathrm{d}^{3}\mathbf{p}\mathrm{d}^{3}\mathbf{q}} = \frac{\delta^{2}\mathcal{F}[z,\overline{z}]}{\delta z(\mathbf{p})\delta\overline{z}(\mathbf{q})}\Big|_{z=\overline{z}=1}$$

ONE-LOOP GENERATING FUNCTIONAL



- For a real external field, the time-ordered scattering matrices ${\mathfrak T}_\pm$ can be related to the retarded one, ${\mathfrak T}_{_R}$
- For a spatially homogeneous external field:

$$\ln \mathcal{F}[z, \overline{z}] = \text{constant} - V \int \frac{d^3 p}{(2\pi)^3} \ln \left[1 - (z(p)\overline{z}(-p) - 1) f_p \right]$$

$$i \mathcal{T}_{_{R}}(p,-k) = 2E_{p} (2\pi)^{3} \delta(p+k) \beta_{p}, \quad f_{p} \equiv \left|\beta_{p}\right|^{2}$$

• In practice: $\mathcal{T}_{R}(p,-k)$ is obtained by solving the classical EOM for φ , starting in the past with a negative frequency plane wave of momentum -k and projecting it in the future on a positive frequency plane wave of momentum p

SPECTRA

$$\frac{dN_{1}^{+}}{d^{3}p} = \frac{dN_{1}^{-}}{d^{3}p} = \underbrace{\frac{V}{(2\pi)^{3}} f_{p}}_{\equiv n_{p}}$$

$$\frac{dN_{2}^{++}}{d^{3}pd^{3}p'} - \frac{dN_{1}^{+}}{d^{3}p} \frac{dN_{1}^{+}}{d^{3}p'} = \delta(p-p') n_{p} f_{p}$$

$$\frac{dN_{2}^{+-}}{d^{3}pd^{3}q} - \frac{dN_{1}^{+}}{d^{3}p} \frac{dN_{1}^{-}}{d^{3}q} = \delta(p+q) n_{p} (1+f_{p})$$

- No correlations at different momenta for particles
- · Particles and antiparticles of opposite momenta are correlated
- Consistent with Poisson distribution if $f_{\rm p} \ll 1$
- In general: Bose enhancement

• For m particles and n antiparticles of momentum k:

$$P_{\mathbf{k}}(m,n) = \delta_{m,n} \; \frac{1}{1+f_{\mathbf{k}}} \; \left(\frac{f_{\mathbf{k}}}{1+f_{\mathbf{k}}}\right)^n \label{eq:product}$$

Note: longer tails than Poisson (Bose enhancement)

• Vacuum persistence probability:

$$P_0 = \exp\Big\{-V\underbrace{\int \frac{d^3k}{(2\pi)^3} \ln(1+f_k)}_{\neq \int_k f_k}\Big\}$$

Bogoliubov Transformation

Mode functions ($A^0 = 0$ gauge)

$$\begin{split} \varphi(\mathbf{x}) &= \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \left[\frac{a_{\mathbf{k},\mathrm{in}}}{\sqrt{2E_{\mathbf{k}}^{\mathrm{in}}}} \varphi_{\mathbf{k},\mathrm{in}}^+(\mathbf{x}) + \frac{b_{\mathbf{k},\mathrm{in}}^\dagger}{\sqrt{2E_{-\mathbf{k}}^{\mathrm{in}}}} \varphi_{-\mathbf{k},\mathrm{in}}^-(\mathbf{x}) \right] \\ &= \int \frac{d^3 \mathbf{k}}{(2\pi)^3} \left[\frac{a_{\mathbf{k},\mathrm{out}}}{\sqrt{2E_{\mathbf{k}}^{\mathrm{out}}}} \varphi_{\mathbf{k},\mathrm{out}}^+(\mathbf{x}) + \frac{b_{\mathbf{k},\mathrm{out}}^\dagger}{\sqrt{2E_{-\mathbf{k}}^{\mathrm{out}}}} \varphi_{-\mathbf{k},\mathrm{out}}^-(\mathbf{x}) \right] \\ \mathsf{E}_{\mathbf{k}}^{\mathrm{in,out}} &= \sqrt{m^2 + (\mathbf{k} + e\mathbf{A}^{\mathrm{in,out}})^2} \end{split}$$

$$\begin{split} \varphi^+_{k,\mathrm{in}}(x) &= e^{-iE_k^{\text{in}}x^0 + i\mathbf{k}\cdot \mathbf{x}} & \text{for } x^0 \to -\infty \\ \varphi^-_{k,\mathrm{in}}(x) &= e^{iE_k^{\text{in}}x^0 + i\mathbf{k}\cdot \mathbf{x}} & \text{for } x^0 \to -\infty \\ \varphi^+_{k,\mathrm{out}}(x) &= e^{-iE_k^{\text{out}}x^0 + i\mathbf{k}\cdot \mathbf{x}} & \text{for } x^0 \to +\infty \\ \varphi^-_{k,\mathrm{out}}(x) &= e^{iE_k^{\text{out}}x^0 + i\mathbf{k}\cdot \mathbf{x}} & \text{for } x^0 \to +\infty \end{split}$$

- EOM linear: there is a linear mapping between the coefficients in the in and out representations
- Spatially homogeneous background: no mixing between the modes k

$$a_{k,\mathrm{out}} = \alpha_k \, a_{k,\mathrm{in}} + \beta_k \, b^\dagger_{-k,\mathrm{in}}, \quad b^\dagger_{k,\mathrm{out}} = \alpha^*_{-k} \, b^\dagger_{k,\mathrm{in}} + \beta^*_{-k} \, a_{-k,\mathrm{in}}$$

- Consistency with canonical commutation relations: $|\alpha_{\mathbf{k}}|^2 |\beta_{\mathbf{k}}|^2 = 1$
- Inverse mapping:

$$a_{k,in} = \alpha_k^* a_{k,out} - \beta_k b_{-k,out}^{\dagger}, \quad b_{k,in} = \alpha_{-k}^* b_{k,out} - \beta_{-k} a_{-k,out}^{\dagger}$$

• In and Out vacua are related by:

$$\left| \theta_{\rm in} \right\rangle = e^{-\frac{V}{2} \int \frac{d^3 p}{(2\pi)^3} \ln(1+|\beta_{\rm p}|^2)} e^{\int \frac{d^3 p}{(2\pi)^3} \frac{\beta_{\rm p}}{\alpha_{\rm p}^*} \alpha_{\rm p,out}^\dagger b_{-\rm p,out}^\dagger} \left| \theta_{\rm out} \right\rangle$$

- This formula contains all the information about final state particle content when the initial state is empty
- Example:

$$\frac{dN_{1}^{+}}{d^{3}p} = \frac{1}{(2\pi)^{3}} \left\langle \boldsymbol{\theta}_{\mathrm{in}} \right| \boldsymbol{\alpha}_{p,\mathrm{out}}^{\dagger} \boldsymbol{\alpha}_{p,\mathrm{out}} \left| \boldsymbol{\theta}_{\mathrm{in}} \right\rangle = \frac{V}{(2\pi)^{3}} \left| \boldsymbol{\beta}_{p} \right|^{2}$$

EXAMPLE: SAUTER POTENTIAL



$$\frac{dN_1}{d^3p} = \frac{V}{(2\pi)^3} \left(\frac{\sinh\left[\pi(\lambda+\mu-\nu)\right]\sinh\left[\pi(\lambda-\mu+\nu)\right]}{\sinh\left(2\pi\mu\right)\sinh\left(2\pi\nu\right)} \right)$$

$$\mu \equiv \frac{\tau}{2} \sqrt{m^2 + p_\perp^2 + (p_z - 2eE\tau)^2}$$
$$\nu \equiv \frac{\tau}{2} \sqrt{m^2 + p_\perp^2 + p_z^2}$$
$$\lambda \equiv eE\tau^2$$

Note: analytic in <code>eE</code> as long as $\tau < \infty$

SAUTER POTENTIAL: p_z SPECTRUM AT FIXED p_\perp



Thin horizontal dotted line: $\exp(-\pi(p_{\perp}^2 + m^2)/(eE))$
SAUTER POTENTIAL: p_z SPECTRUM AT VARIOUS p_\perp (fixed $\sqrt{eE} au=4$)



Thin horizontal dotted line: $\exp(-\pi(p_{\perp}^2 + m^2)/(eE))$

Numerical evaluation on the lattice

$$\frac{dN_1}{d^3p} = \frac{1}{(2\pi)^3 2\mathsf{E}_p} \int \frac{d^3k}{(2\pi)^3 2\mathsf{E}_k} \left| \int d^3x \; \varphi_{p,\mathrm{out}}^{+\dagger}(x) \left(\vartheta_t - i\mathsf{E}_p \right) \varphi_{k,\mathrm{in}}^{-}(x) \right|_{x_0 = +\infty}^2$$

- Note: in general, time evolution non-diagonal in p
- This observable belongs to a generic class of objects that can be written as

$$\left\langle \Phi^{\dagger} \mathbf{M} \Phi \right\rangle \equiv \int \frac{\mathrm{d}^{3} \mathbf{k}}{(2\pi)^{3} 2 \mathsf{E}_{\mathbf{k}}} \left[\Phi_{\mathbf{k}, \mathrm{in}}^{-\dagger} \mathbf{M} \Phi_{\mathbf{k}, \mathrm{in}}^{-} \right]_{x^{0} = y^{0} = +\infty}$$

 $(M_{xy} = Hermitean operator)$

"BRUTE FORCE" LATTICE APPROACH

- Discretize space as a $N\times N\times N$ lattice
- Use link variables instead of Aⁱ to ensure gauge invariance
- N³ conjugate momenta
- Solve the EOM for N^3 incoming mode functions
- Numerical cost $\sim N_{t} \times N^{3} \times N^{3}~~$ (N $_{t} =$ number of time steps)
 - \rightarrow quite unfavorable scaling
- Note: if the field is independent of some of the coordinates, this cost can be reduced since the evolution conserves the corresponding momentum

STATISTICAL SAMPLING [FG, TANJI, 2013]

- · Goal: avoid summing over all the mode functions to reduce cost
- Strategy: exploit the linearity of the EOM and solve it for a random linear superposition of all the mode functions

$$\varphi^-_c(x) \equiv \frac{1}{\sqrt{V}} \sum_{\mathbf{k}} \frac{1}{\sqrt{2E_{\mathbf{k}}}} c_{\mathbf{k}} \varphi^-_{\mathbf{k},\mathrm{in}}(x)$$

 $c_{\bf k}=$ Gaussian random numbers with $\langle c_{\bf k}\rangle=0$ and $\langle c_{\bf k}c^*_{{\bf k}'}\rangle_c=\delta_{{\bf k}{\bf k}'}$

$$\left\langle \varphi^{\dagger} M \varphi \right\rangle = \left\langle \left[\varphi_{c}^{-\dagger} M \varphi_{c}^{-} \right]_{x^{0} = y^{0} = +\infty} \right\rangle_{c}$$

- Numerical cost: $N_t \times N^3 \times N_c$ (plus $N_c \times N^6$ for preparing the initial conditions) \rightarrow favorable if $N_c \ll N^3, N_t$
- Statistical error $\sim N_c^{-1/2}$
- Related to *low cost fermions* [Borsanyi, Hindmarsh; Saffin, Tranberg; Berges, Gelfand, Sexty, Kasper, Hebenstreit, 2009-2014]

EXAMPLE: SAUTER FIELD



•
$$N_x = N_y = 48$$
, $N_z = 128$
• $\sqrt{eE} a_x = \sqrt{eE} a_y = 0.42$, $\sqrt{eE} a_z = 0.16$,
• $N_c = 256 (\ll 48^2 \times 128 = 294912)$

WORKS ALSO FOR WEAK FIELDS



•
$$eE = 0.25 \text{ m}^2$$
, $\sqrt{eE} \tau = 25.5$

•
$$N_x = N_y = 48$$
, $N_z = 256$, $ma_z = 0.048$

•
$$N_c = 48$$

- So far, assume that the external field is unmodified by produced charged particles
- Energy is not conserved in this approximation (roughly ok if the field energy dominates)
- The produced charges screen the external field, and weaken it
- Feedback can be included by simultaneously solving Maxwell's equation:

$$\partial_{\mu}F^{\mu\nu}(\mathbf{x}) = \langle \hat{\mathbf{j}}^{\nu}(\mathbf{x}) \rangle$$

 $\langle \hat{J}^\nu(x) \rangle =$ quantum expectation value of the current operator

EXAMPLE: INITIALLY CONSTANT E₀



ENERGY CONSERVATION



• Energy carried by the field and particles, normalized by $\mathcal{E}_0 \equiv \frac{1}{2}E_0^2$

p_z SPECTRUM (EARLY TIMES)



• Very similar to the Sauter potential (charges produced with $p_z \approx 0$ and accelerated in the +z direction)

p_z SPECTRUM (LATER TIMES)



- The field direction oscillates, and the acceleration changes sign
- Existing particles encounter newly created ones, and Pauli blocking leads to interferences

Worldline Formalism

[Bern, Kosower, 1988; Strassler, 1992] [Schubert, 1996, 2001] [Schmidt, Schubert, 1993] [Dunne, Schubert, 2005]

SETUP

• Total particle production probability (at one loop):

$$\begin{split} \left< \vartheta_{\rm out} \left| \vartheta_{\rm in} \right> &= e^{i\,V}, \quad \sum_{n=1}^{\infty} P_n = 1 - P_0 = 1 - e^{-2\,{\rm Im}\,V} \\ i\,V &= \sum \left(\text{connected vacuum diagrams} \right) \\ \text{Scalar QED}: \quad V_{1\,\,{\rm loop}} &= \ln\,\det\,\left(g_{\mu\nu}D^{\mu}D^{\nu} + m^2\right) \end{split}$$

• Worldline formalism is Euclidean, so consider instead:

$$V_{{}_{\text{E}},1\;\text{loop}} \equiv \ln\,\det\,\left(-D^{\,\text{i}}D^{\,\text{i}}+\mathfrak{m}^2\right) = \mathrm{tr}\,\ln\left(-D^{\,\text{i}}D^{\,\text{i}}+\mathfrak{m}^2\right)$$

• Schwinger proper time representation:

$$\begin{split} \left(-D^{i}D^{i}+\mathfrak{m}^{2}\right)^{-1} &= \int_{0}^{\infty} dT \; \exp\left(-T\left(-D^{i}D^{i}+\mathfrak{m}^{2}\right)\right) \\ \ln\left(-D^{i}D^{i}+\mathfrak{m}^{2}\right) &= -\int_{0}^{\infty} \frac{dT}{T} \; \exp\left(-T\left(-D^{i}D^{i}+\mathfrak{m}^{2}\right)\right) \end{split}$$

$$V_{E,1 \text{ loop}} = -\int_0^\infty \frac{dT}{T} e^{-m^2 T} \\ \times \int_{x^i(0)=x^i(T)} [Dx^i(\tau)] \exp\left(-\int_0^T d\tau \left(\frac{\dot{x}^i \dot{x}^i}{4} + ie \, \dot{x}^i A^i(x)\right)\right)$$

- + $\kappa^i(\tau) =$ trajectory of length T in Euclidean spacetime of a fictitious point-like particle
- Closed paths because of the trace: $x^i(0) = x^i(T)$
- The mass suppresses the long paths (longer than the Compton wavelength). $T\approx$ 0 controls the UV
- Euclidean metric ensures convergence
- In vacuum, one has

$$\int_{x^{i}(0)=x^{i}(T)} \left[Dx^{i}(\tau) \right] \exp\left(-\int_{0}^{T} d\tau \; \frac{\dot{x}^{i} \dot{x}^{i}}{4} \right) = \frac{1}{(4\pi T)^{d/2}} \underset{d=4}{=} \frac{1}{(4\pi T)^{2}}$$

SCALES



- Path length = T
- + Size of explored region $\sim \sqrt{T}$
- Area $\sim T$

• Split x^i into barycenter of the loop and deviation:

$$x^i(\tau)\equiv X^i+r^i(\tau),\quad \int_0^T d\tau\;r^i(\tau)=0$$

- Background field \rightarrow Wilson loop centered at X^i , averaged over all paths of length T:

$$W_{\rm x}[\mathbf{r}] \equiv \exp\left(-ie\int_0^{\mathsf{T}} d\tau \,\dot{\mathbf{r}}^i(\tau) A^i(\mathbf{X}+\mathbf{r}(\tau))\right)$$
$$\langle W_{\rm x} \rangle_{\mathsf{T}} \equiv (4\pi\mathsf{T})^2 \int_{\mathbf{r}^i(0)=\mathbf{r}^i(\mathsf{T})} [\mathsf{D}\mathbf{r}^i(\tau)] \, W_{\rm x}[\mathbf{r}] \, \exp\left(-\int_0^{\mathsf{T}} d\tau \, \frac{\dot{\mathbf{r}}^i \dot{\mathbf{r}}^i}{4}\right)$$

- Average is dominated by an ensemble of loops localized around the barycenter X^i (up to a distance of order $T^{1/2}$)
- + $\langle W_{\rm X} \rangle_{\tau}$ encapsulates the local properties of the quantum field theory in the vicinity of X^i

• One-loop Euclidean vacuum diagrams:

$$\mathcal{V}_{\rm E,1\ loop} = -\frac{1}{(4\pi)^2} \int d^4 X \int_0^\infty \frac{d\mathsf{T}}{\mathsf{T}^3} \ e^{-\mathfrak{m}^2 \mathsf{T}} \ \langle W_x \rangle_\mathsf{T}$$

• For a constant E, choose a gauge where A^i is linear in coordinates $\rightarrow \langle W_x \rangle_T$ given by a Gaussian integral:

$$\langle W_{\rm x} \rangle_{\rm T} = \frac{e{\rm ET}}{\sin(e{\rm ET})}$$

- The imaginary part of $\mathcal{V}_{_{E},1\rm\ loop}$ comes from poles located at $T_{n}=n\pi/(eE)$:

Im
$$(\mathcal{V}_{E,1 \text{ loop}}) = \frac{V_4}{16\pi^3} (eE)^2 \sum_{n=1}^{\infty} \frac{(-1)^{n-1}}{n^2} e^{-n\pi m^2/(eE)}$$

• Note: the terms n > 1 encode Bose-Einstein correlations

•
$$\mathsf{E} \parallel z \rightarrow \mathsf{A}^{\mathsf{i}} = (0, 0, 0, -\mathsf{i}\mathsf{E}x^3)$$

$$W_{\rm x}[{\rm r}] = {\rm e}^{-e\,{\rm E},{\cal A}}, \quad {\rm with} \; {\cal A} \equiv \int_0^{\rm T} {\rm d}\tau \; \dot{\rm r}_4(\tau) {\rm r}_3(\tau)$$

(A =projected area of the loop on the plane 34)

• Note: probability distribution for A:

$$\mathcal{P}_{\mathsf{T}}(\mathcal{A}) = \frac{\pi}{4\mathsf{T}} \, \frac{1}{\cosh^2\left(\frac{\pi \mathcal{A}}{2\mathsf{T}}\right)} \quad \rightarrow \text{ typical worldlines: } \mathcal{A} \lesssim \mathsf{T}$$

• After a rescaling $\mathbb{I}\equiv\tau\mathcal{A}\text{, }s\equiv-i\tau/e\text{E}\text{:}$

$$\mathcal{V}_{\rm E,1\;loop} = \left(\frac{e\mathsf{E}}{4\pi}\right)^2 \int d^4X \int_0^\infty \frac{ds}{s^3}\; e^{-\mathfrak{i}(\mathfrak{m}^2/(e\mathsf{E}))s}\; \left\langle e^{-\mathfrak{i}s\mathfrak{I}} \right\rangle$$

- Evaluate $\left\langle e^{-\mathrm{i} \mathrm{s} \Im}
 ight
 angle$ once for all (from an ensemble of loops)
- Difficulty: small $eE/m^2 \rightarrow$ small $s \rightarrow$ large areas needed

CONSTANT FIELD GIES, KLINGMULLER, 2005



ALGORITHM FOR NON-CONSTANT FIELDS

- External field and ensemble of loops are entangled
- Other difficulty: integral over areas converges only for $|T| \le \pi/(eE)$ (location of the first pole, T₁)
- Analytical ansatz for the distribution of areas:

$$W_{X}[r] = e^{-eE(X)TJ}, \quad J \equiv \frac{i}{\tau E(X)} \int_{0}^{T} d\tau \ \dot{r}^{i}(\tau) A^{i}(X + r(\tau))$$

Ansatz: $\mathcal{P}_{X}(J) = N \ \frac{1}{\cosh^{2\nu}\left(\frac{\pi \, \alpha J}{2}\right)}$

Fit α and ν from an ensemble of loops generated by Monte-Carlo. Then, do the ${\mathcal J}$ integral analytically

$$\int_{-\infty}^{+\infty} d\mathfrak{I} \, \mathfrak{P}_{X}(\mathfrak{I}) \, e^{-e \mathsf{E}(X)\tau \mathfrak{I}} = \mathsf{N} \, \frac{4^{\nu}}{\pi \alpha} \, \frac{\Gamma(\nu + \frac{e\mathsf{E}(X)\tau}{\pi \alpha})\Gamma(\nu - \frac{e\mathsf{E}(X)\tau}{\pi \alpha})}{\Gamma(2\nu)}$$

- Integrate over T and X^{i} numerically

SAUTER FIELD $E^1 = E/\cosh^2(kx^1)$ Gies, Klingmuller, 2005



• When the background field results of a lattice computation (and is given in terms of link variables on the lattice), we may write

$$\operatorname{tr}\ln\left(-\mathsf{D}^{i}\mathsf{D}^{i}+\mathfrak{m}^{2}\right)=-\sum_{n=0}^{\infty}\frac{1}{n}\frac{1}{(2\tilde{d})^{n}}\sum_{\mathbf{x}\in \,\operatorname{lattice}}\,\sum_{\gamma\in\Gamma_{n}(\mathbf{x},\mathbf{x})}\prod_{\ell\in\gamma}U_{\ell}$$

- n plays the role of the fictitious time T
- + $U_\ell = link$ variable on the edge ℓ
- + $\Gamma_n(x, x) = \text{set of loops (from } x \text{ to } x) \text{ of length } n \text{ (in lattice units)}$
- $\tilde{d} = d + \frac{1}{2}m^2 a^2$ ($\tilde{d} > d$ suppresses the long loops)



WORLDLINE INSTANTON APPROXIMATION DUNNE, SCHUBERT, 2005

- Define $\tau \equiv Tu$ and $m^2T = s.$ Then:

$$\mathcal{V}_{_{E},1\;\mathrm{loop}} = -\int\limits_{0}^{\infty} \frac{ds}{s}\; e^{-s}_{x^{\,i}(0)=x^{\,i}(1)} \left[Dx^{\,i}(u) \right] \; \exp\left(-\int_{0}^{1} du \left(\frac{m^{2}}{4s} \frac{\dot{x}^{2}}{4} + ie\dot{x}^{\,i}A^{\,i}(x) \right) \right)$$

• The integral over s gives a Bessel function:

$$\mathcal{V}_{E,1 \text{ loop}} = -2 \int_{x^{i}(0)=x^{i}(1)} \left[Dx^{i}(u) \right] K_{0} \left(\left(m \int_{0}^{1} du \, \dot{x}^{2} \right)^{\frac{1}{2}} \right) \exp \left(-ie \int_{0}^{1} du \, \dot{x}^{i} A^{i}(x) \right)$$

• In the regime where $m^2 \int_0^1 du \dot{x}^2 \gg 1$, approximate $K_0(z) \approx \sqrt{\pi/2} e^{-z} / \sqrt{z}$ and perform a stationary phase approximation. We need extrema of

$$\label{eq:states} \begin{split} & \mathcal{S} \equiv \mathfrak{m} \Bigl(\int_0^1 du \; \dot{x}^2 \Bigr)^{1/2} + i e \int_0^1 du \; \dot{x}^i A^i(x) \end{split}$$

- They are the closed paths $\boldsymbol{x}^i(\boldsymbol{u})$ that obey

$$m\frac{\ddot{x}^{i}}{\sqrt{\int_{0}^{1}du\;\dot{x}^{2}}}=ieF^{ij}\dot{x}^{j}$$

Note: $\dot{x}^i \dot{x}^i = \text{const}$

• For each extremum

$${\rm Im}\, {\cal V}_{_E,1\,\,{\rm loop}} \sim e^{-S_{\rm extremum}}$$

• The prefactor is obtained by integrating Gaussian deviations about the extremal path

EXAMPLE: SAUTER FIELD $E(x^3) = E/\cosh^2(kx^3)$

• Use the gauge potential

$$A^4 = -i\,\frac{E}{k}\,\tanh(kx^3)$$

• Equations of motion for the stationary solutions

$$\begin{split} \dot{x}^3 &= \nu \sqrt{1-\gamma^{-2} \tanh^2(kx^3)} , \qquad \dot{x}^4 &= -\gamma^{-1} \nu \tanh(kx^3) \\ \nu &\equiv mk/(eE)) \end{split}$$

• Countable infinity of periodic solutions:

$$\begin{aligned} x^{3}(u) &= \frac{m}{eE} \frac{1}{\gamma} \operatorname{arcsinh} \left(\frac{\gamma}{\sqrt{1 - \gamma^{2}}} \sin(2\pi n \, u) \right) \\ x^{4}(u) &= \frac{m}{eE} \frac{1}{\gamma \sqrt{1 - \gamma^{2}}} \arcsin\left(\gamma \, \cos(2\pi n \, u)\right) \end{aligned}$$

(n = winding index of the solution)

- Extended field ($\gamma \rightarrow$ 0): circular solutions
- + $\gamma \rightarrow$ 1: very elongated orbits, action becomes infinite
 - \rightarrow no pair production (field coherence length too small)



Dynamically assisted Schwinger effect

[Schutzhold, Gies, Dunne, 2008] [Di Piazza, Lotstedt, Milstein, Keitel, 2009] [Orthaber, Hebenstreit, Alkofer, 2011] [Monin, Voloshin, 2012] [Taya, Fujii, Itakura, 2014] + many others

COMPARISON BETWEEN EXACT AND PERTURBATIVE RESULTS

- Consider the Sauter temporal pulse: $A^3(t)=E\tau \tanh(t/\tau)$
- Exact spectrum:

$$\frac{dN_1}{d^3p} = \frac{V}{(2\pi)^3} \left(\frac{\sinh\left[\pi(\lambda + \mu - \nu)\right] \sinh\left[\pi(\lambda - \mu + \nu)\right]}{\sinh\left(2\pi\mu\right) \sinh\left(2\pi\nu\right)} \right)$$

$$\begin{split} \mathfrak{u} &\equiv \frac{\tau}{2} \sqrt{\mathfrak{m}^2 + \mathfrak{p}_{\perp}^2 + (\mathfrak{p}_z - 2 \mathbf{e} \mathbf{E} \tau)^2} \\ \mathbf{v} &\equiv \frac{\tau \mathbf{E}_{\mathbf{p}}}{2}, \quad \lambda \equiv \mathbf{e} \mathbf{E} \tau^2 \end{split}$$

One-photon spectrum:

$$\stackrel{p+p'}{\longrightarrow} \longrightarrow \frac{dN_1^{(1\gamma)}}{d^3p} = \frac{Ve^2E^2}{(2\pi)^3} \left[1 - \left(\frac{p_z}{E_p}\right)^2\right] \frac{\pi^2\tau^4}{\sinh^2(\pi E_p\tau)}$$

NUMERICAL COMPARISON FOR SUBCRITICAL FIELDS



- · Solid lines: exact result. Dashed lines: one-photon result
- Black dotted line: constant field result ($\exp(-\pi m^2/(eE))$)
- Note: considerably enhanced spectrum in the regime $m\tau\sim 1$

• Consider the sum of two Sauter fields, with $E_1 \gg E_2$ and $\tau_1 \gg \tau_2$:

$$E_z(t) = \frac{E_1}{\cosh^2\left(\frac{t}{\tau_1}\right)} + \frac{E_2}{\cosh^2\left(\frac{t}{\tau_2}\right)}$$

- E1: strong and slow (one-photon process forbidden)
- E2: weak and fast (one-photon process possible)
- Non-trivial effects since the spectrum is non-linear in the field



•
$$eE_1 = 0.25 m_e^2$$
, $\tau_1 = 10^{-4} eV^{-1}$
• $eE_2 = 0.025 m_e^2$, $\tau_2 = 7 \times 10^{-6} eV^{-1}$

QUALITATIVE INTERPRETATION



- Slope due to the strong and slow field E₁ (no appreciable slope from the weak field)
- A single photon from the weak and fast field E₂ raises a hole excitation (the more, the better)
- Tunneling distance is reduced, which affects exponentially the resulting spectrum



•
$$eE_1 = 0.25 m_e^2$$
, $m_e \tau_1 = 510$

- $eE_2 = 0.025 m_e^2$, variable $m_e \tau_2$
- Black dashed line: E1 alone
- Maximal enhancement when $m_e\tau_2\approx 0.6~~\text{(roughly,}~\tau_2\sim (2m_e)^{-1}\text{)}$

Summary
- Very interesting playground for studying QFT in a non-perturbative regime and testing novel methods (with a few exact results to compare with)
- A number of approaches have been applied to this problem:
 - Mode functions on the lattice
 - Worldline formalism
 - Quantum kinetic equations
 - Wigner formalism
 - Holography, AdS/CFT
- For high enough frequency, perturbative (one-photon) result dominates over constant field result
- Dynamical enhancement can achieved by superimposing slow and fast fields