High energy factorization in Nucleus-Nucleus collisions

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Outline

Introduction

- Single gluon spectrum
- Initial field perturbation
- JIMWLK Hamiltonian
- Factorization
- Final remarks

- Introduction
- Single gluon spectrum at LO and NLO
- Expression as variations of the initial fields
- Leading log divergences and JIMWLK Hamiltonian
- Leading Log factorization
- Final remarks

(FG, T. Lappi and R. Venugopalan, in preparation)



- Parton saturation
- Color Glass Condensate
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Saturation domain







CGC degrees of freedom

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The fast partons (large x) are frozen by time dilation
 b described as static color sources on the light-cone :

$$J_a^{\mu} = \delta^{\mu +} \delta(x^-) \rho_a(\vec{x}_{\perp}) \qquad (x^- \equiv (t-z)/\sqrt{2})$$

Slow partons (small x) cannot be considered static over the time-scales of the collision process > they must be treated as the usual gauge fields

Since they are radiated by the fast partons, they must be coupled to the current J^{μ}_{a} by a term : $A_{\mu}J^{\mu}$

The color sources ρ_a are random, and described by a distribution functional $W_Y[\rho]$, with Y the rapidity that separates "soft" and "hard"



CGC evolution

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Evolution equation (JIMWLK) :

$$\frac{\partial W_{_{\boldsymbol{Y}}}}{\partial Y} = \mathcal{H} \ W_{_{\boldsymbol{Y}}}$$

$$\mathcal{H} = \frac{1}{2} \int_{\vec{\boldsymbol{y}}_{\perp}} \frac{\delta}{\delta \widetilde{\mathcal{A}}_{b}^{+}(\epsilon, \vec{\boldsymbol{y}}_{\perp})} \eta_{ab}(\vec{\boldsymbol{x}}_{\perp}, \vec{\boldsymbol{y}}_{\perp}) \frac{\delta}{\delta \widetilde{\mathcal{A}}_{a}^{+}(\epsilon, \vec{\boldsymbol{x}}_{\perp})}$$

where $-\partial_{\perp}^2 \widetilde{\mathcal{A}}^+(\epsilon, \vec{x}_{\perp}) =
ho(\epsilon, \vec{x}_{\perp})$

- η_{ab} is a non-linear functional of ρ
- This evolution equation resums the powers of $\alpha_s \ln(1/x)$ and of Q_s/p_{\perp} that arise in loop corrections
- This equation simplifies into the BFKL equation when the color density ρ is small (one can expand η_{ab} in ρ)



Nucleus-nucleus collisions





CGC and Nucleus-Nucleus collisions

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$$\mathcal{L} = -\frac{1}{2} \operatorname{tr} F_{\mu\nu} F^{\mu\nu} + (\underbrace{J_1^{\mu} + J_2^{\mu}}_{J^{\mu}}) A_{\mu}$$

- Given the sources $\rho_{1,2}$ in each projectile, how do we calculate observables? Is there some kind of perturbative expansion?
- Loop corrections and factorization?



Initial particle production



- Parton saturation
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Dilute regime : one parton in each projectile interact



Initial particle production



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- Dilute regime : one parton in each projectile interact
- Dense regime : multiparton processes become crucial
 - (+ pileup of many partonic scatterings in each AA collision)



What is factorization ?

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A factorization formula divides an observable into a perturbatively calculable part (involving quarks and gluons) and a non-perturbative part describing the partonic content of hadrons or nuclei :

 $\mathcal{O} = F \otimes \mathcal{O}_{\text{partonic}}$

- Factorization has no predictive power unless the distributions *F* are intrinsic properties of the incoming projectiles :
 - *F* cannot depend on the observable
 - *F* of one projectile cannot depend on the second projectile
- Factorization can accommodate certain resummations :
 - Loop corrections in QCD generate corrections of the form $[\alpha_s \log(\cdot)]^n$, that are large in some parts of the phase-space
 - When these corrections do not depend on the observable and projectiles, they can be absorbed in the definition of *F* via an universal evolution equation



Factorization in the dilute regime

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- Factorization in the dilute small-x regime is known as $k_{\rm T}$ -factorization
- It was introduced in the discussion of heavy quark production near threshold, when $s \gg 4m_q^2$, to resum large logs of $1/x_{1,2}$ Collins, Ellis (1991), Catani, Ciafaloni, Hautmann (1991) Levin, Ryskin, Shabelski, Shuvaev (1991)
- In this framework, cross-sections read :

$$\begin{split} \frac{d\sigma}{dY d^2 \vec{P}_{\perp}} \propto \int\limits_{\vec{k}_{1\perp},\vec{k}_{2\perp}} \delta(\vec{k}_{1\perp} + \vec{k}_{2\perp} - \vec{P}_{\perp}) \varphi_1(x_1, k_{1\perp}) \varphi_2(x_2, k_{2\perp}) \frac{|\mathcal{M}|^2}{k_{1\perp}^2 k_{2\perp}^2} \\ x_{1,2} = \frac{M_{\perp}}{\sqrt{s}} e^{\pm Y} \end{split}$$

The small-x leading logs are resummed into the non-integrated gluon distributions \u03c6_{1,2} by letting them evolve according to the BFKL equation



Factorization in the dense regime

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- In the dense regime, observables are sensitive to parton correlations beyond 2-point correlations. The distributions $\varphi_{1,2}$ do not provide this information, but it is present in the source distributions $W[\rho_{1,2}]$ of the CGC
- Factorization in the dense regime at small-x has been established for DIS. The leading logs can be absorbed into W[\rho] by letting it evolve according to the JIMWLK equation
- In the collision of two dense projectiles :
 - The large logs have a coefficient that depends in a complicated way on the sources of both nuclei. One must show that they can still be absorbed in one of the two W[p]'s
 - The dependence of the observable on the sources $\rho_{1,2}$ is not known analytically, already at LO
 - Even less is known about loop corrections...



Factorization in the dense regime

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For the single gluon spectrum in AA collisions, one would like to establish a formula such as :

$$\left\langle \frac{dN}{d^{3}\vec{p}} \right\rangle_{\text{LLog}} = \int \left[D\rho_{1} D\rho_{2} \right] W_{Y_{\text{beam}} - y} [\rho_{1}] W_{y + Y_{\text{beam}}} [\rho_{2}] \left. \frac{dN}{d^{3}\vec{p}} \right|_{\text{LO}}$$

with $\frac{\partial}{\partial Y} W_{Y} = \mathcal{H} W$
- $Y_{\text{beam}} \rho_{2} y \rho_{1} + Y_{\text{beam}}$

- All the leading logs of $1/x_{1,2}$ are absorbed in the W's
- The W's obey the JIMWLK evolution equation



Factorization in four easy steps

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- I: Express the single gluon spectrum at LO and NLO in terms of classical fields and small field fluctuations. Check that their boundary conditions are retarded
- II : Write the NLO terms as a perturbation of the initial value of the classical fields on the light-cone :

$$\frac{dN}{d^{3}\vec{p}}\Big|_{\rm NLO} = \left[\frac{1}{2}\int\limits_{\vec{u},\vec{v}\in\rm LC} \mathcal{G}(\vec{u},\vec{v})\,\mathbb{T}_{u}\,\mathbb{T}_{v} + \int\limits_{\vec{u}\in\rm LC} \beta(\vec{u})\,\mathbb{T}_{u}\right] \left.\frac{dN}{d^{3}\vec{p}}\right|_{\rm LO}$$

For \vec{u}, \vec{v} on the same branch of the light-cone, one has :

$$\frac{1}{2} \int_{\vec{u}, \vec{v} \in LC} \mathcal{G}(\vec{u}, \vec{v}) \mathbb{T}_{u} \mathbb{T}_{v} + \int_{\vec{u} \in LC} \beta(\vec{u}) \mathbb{T}_{u} = \log\left(\frac{\Lambda^{+}}{p^{+}}\right) \times \mathcal{H} + \text{ finite terms}$$

IV : There are no other logs. Factorization follows trivially



Single gluon spectrum

Leading Order

Next to Leading Order

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Single gluon spectrum at LO and NLO



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Next to Leading Order

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- Leading Order = tree diagrams only
- Tag one gluon of momentum \vec{p}
- Integrate out the phase-space of all the other gluons

$$\frac{dN}{d^3\vec{\boldsymbol{p}}} \sim \sum_{n=0}^{\infty} \frac{1}{n!} \int \left[d^3\vec{\boldsymbol{p}}_1 \cdots d^3\vec{\boldsymbol{p}}_n \right] \left| \left\langle \vec{\boldsymbol{p}} \ \vec{\boldsymbol{p}}_1 \cdots \vec{\boldsymbol{p}}_n \right| 0 \right\rangle \right|^2$$



Single gluon spectrum

- Leading Order
- Next to Leading Order

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Single gluon spectrum at LO

- LO results for the single gluon spectrum :
 - Disconnected graphs cancel in the inclusive spectrum
 - At LO, the single gluon spectrum can be expressed in terms of classical solutions of the field equation of motion
 - These classical fields obey retarded boundary conditions

$$\frac{dN}{d^3\vec{p}}\Big|_{\rm LO} = \lim_{t \to +\infty} \int d^3\vec{x} d^3\vec{y} \ e^{i\vec{p}\cdot(\vec{x}-\vec{y})} \ \cdots \mathcal{A}^{\mu}(t,\vec{x}) \ \mathcal{A}^{\nu}(t,\vec{y})$$
$$\left[\mathcal{D}_{\mu},\mathcal{F}^{\mu\nu}\right] = J^{\nu}$$

$$\lim_{t \to -\infty} \mathcal{A}^{\mu}(t, \vec{x}) = 0$$

Note : retarded boundary conditions play an important role in the following. They are not automatic, but seem generic for inclusive observables



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Retarded classical fields are sums of tree diagrams :





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Retarded classical fields are sums of tree diagrams :



Note : the gluon spectrum is a functional of the value of the classical field just above the backward light-cone :

$$\frac{dN}{d^3\vec{\boldsymbol{p}}} = \mathcal{F}[\mathcal{A}_{\text{initial}}]$$







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- Next to Leading Order = 1-loop diagrams
- Connected diagrams only
- Expressible in terms of classical fields, and small fluctuations about the classical field, both with retarded boundary conditions



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1-loop graphs contributing to the gluon spectrum at NLO :



$$\frac{dN}{d^3\vec{\boldsymbol{p}}}\bigg|_{_{\rm NLO}} = \lim_{t \to +\infty} \int d^3\vec{\boldsymbol{x}} d^3\vec{\boldsymbol{y}} \ e^{i\vec{\boldsymbol{p}}\cdot(\vec{\boldsymbol{x}}-\vec{\boldsymbol{y}})} \ \cdots \left[\mathcal{G}^{\mu\nu}(x,y)\right]$$

 $+\beta^{\mu}(t,\vec{\boldsymbol{x}}) \mathcal{A}^{\nu}(t,\vec{\boldsymbol{y}}) + \mathcal{A}^{\mu}(t,\vec{\boldsymbol{x}}) \beta^{\nu}(t,\vec{\boldsymbol{y}})$

• $\mathcal{G}^{\mu\nu}$ is a 2-point function on top of the classical field

• β^{μ} is a small field fluctuation driven by a 1-loop source



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with

• The 2-point function $\mathcal{G}^{\mu\nu}$ can be written as

$$\mathcal{G}^{\mu\nu}(x,y) = \int \frac{d^3 \vec{k}}{(2\pi)^3 2E_k} \ \eta^{\mu}_{-k}(x) \ \eta^{\nu}_{+k}(y)$$

$$\begin{cases} \left[\mathcal{D}_{\mu}, \left[\mathcal{D}^{\mu}, \eta_{\pm \boldsymbol{k}}^{\nu}\right]\right] - \left[\mathcal{D}_{\mu}, \left[\mathcal{D}^{\nu}, \eta_{\pm \boldsymbol{k}}^{\mu}\right]\right] + ig\left[\mathcal{F}_{\mu}{}^{\nu}, \eta_{\pm \boldsymbol{k}}^{\mu}\right] = 0\\ \lim_{t \to -\infty} \eta_{\pm \boldsymbol{k}}^{\mu}(t, \vec{\boldsymbol{x}}) = \epsilon^{\mu}(\boldsymbol{k}) \ e^{\pm ik \cdot \boldsymbol{x}} \end{cases}$$

(obtained by writing the YM equation for $\mathcal{A} + \eta_{\pm k}$, linearized in $\eta_{\pm k}$)

• The equation of motion for β^{μ} reads

$$\begin{split} \left[\mathcal{D}_{\mu}, \left[\mathcal{D}^{\mu}, \beta^{\nu} \right] \right] &- \left[\mathcal{D}_{\mu}, \left[\mathcal{D}^{\nu}, \beta^{\mu} \right] \right] + ig \left[\mathcal{F}_{\mu}{}^{\nu}, \beta^{\mu} \right] = \\ &= \underbrace{\frac{\partial^{3} \mathcal{L}_{YM}(\mathcal{A})}{\partial \mathcal{A}^{\nu}(x) \partial \mathcal{A}^{\rho}(x) \partial \mathcal{A}^{\sigma}(x)}}_{3g \text{ vertex in the background } \mathcal{A}} \underbrace{\frac{1}{2} \int \frac{d^{3} \vec{k}}{(2\pi)^{3} 2E_{k}} \eta^{\mu}_{-k}(x) \eta^{\nu}_{+k}(x)}_{\text{value of the loop}} \end{split}$$



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Expression as a perturbation of the initial classical field



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Single gluon spectrum at NLO

The retarded nature of the field fluctuations allows a factorization between the initial condition (calculable analytically) and the evolution on top of \mathcal{A}^{μ} (complicated) :

$$a^{\mu}(x) = \underbrace{\left[\int_{\vec{u} \in \mathrm{LC}} a(u) \cdot \mathbb{T}_{u}\right]}_{\mathbf{u} \in \mathrm{LC}} \mathcal{A}^{\mu}(x)$$

initial condition

- 'LC' is a surface just above the backward light-cone
- \mathbb{T}_u is the generator of shifts of the initial value of the fields on this surface :

$$\mathcal{F}[\mathcal{A}_{\text{initial}} + a] \equiv \exp\left[\int_{\vec{u} \in \text{LC}} a(u) \cdot \mathbb{T}_{u}\right] \mathcal{F}[\mathcal{A}_{\text{initial}}]$$

Note : this construction is possible only because the objects involved in the problem obey retarded boundary conditions

(u)



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This factorization can be applied to the NLO gluon spectrum:







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This factorization can be applied to the NLO gluon spectrum:



They can be written as a perturbation of the LC initial fields :

$$\frac{dN}{d^{3}\vec{p}}\Big|_{_{\rm NLO}} = \begin{bmatrix} \frac{1}{2} \int \mathcal{G}(\vec{u},\vec{v}) \,\mathbb{T}_{u} \,\mathbb{T}_{v} \end{bmatrix} \frac{dN}{d^{3}\vec{p}}\Big|_{_{\rm LO}}$$

below the LC above the LC
$$\mathcal{G}(\vec{u},\vec{v}) \equiv \int \frac{d^{3}\vec{k}}{(2\pi)^{3}2E_{k}} \,\eta_{-k}(u) \,\eta_{+k}(v)$$





This factorization can be applied to the NLO gluon spectrum:



■ The loop can also be below the light-cone :

χ 👝

$$\frac{dN}{d^{3}\vec{p}}\Big|_{\rm NLO} = \left[\frac{1}{2}\int\limits_{\vec{u},\vec{v}\in \rm LC} \mathcal{G}(\vec{u},\vec{v}) \,\mathbb{T}_{u} \,\mathbb{T}_{v} + \int\limits_{\vec{u}\in \rm LC} \beta(\vec{u}) \,\mathbb{T}_{u}\right] \left.\frac{dN}{d^{3}\vec{p}}\right|_{\rm LO}$$

below the LC above the LC

 \triangleright the functions $\mathcal{G}(\vec{u}, \vec{v})$ and $\beta(\vec{u})$ can be evaluated analytically



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If \vec{u}, \vec{v} belong to the same branch of the LC (e.g. $u^- = v^- = \epsilon$), the function $\mathcal{G}(\vec{u}, \vec{v})$ contains

$$\mathcal{G}(\vec{u},\vec{v}) \sim \int_0^{+\infty} \frac{dk^+}{k^+} \cdots e^{ik^-(u^+ - v^+)} \quad \text{with} \quad k^- \equiv \frac{k_\perp^2}{2k^+}$$

> the integral converges at $k^+ = 0$ but not when $k^+ \to +\infty$

Note : the log is a $\log(\Lambda^+/p^+)$, where Λ^+ is the boundary between the hard color sources and the fields, and p^+ the longitudinal momentum of the produced gluon





Leading Log approximation

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In the LC gauge
$$\mathcal{A}^+ = 0$$
, the operator $\eta(u) \cdot \mathbb{T}_u$ is

$$\eta(u) \cdot \mathbb{T}_{\boldsymbol{u}} \equiv (\partial^{-} \eta_{a}^{i}(u)) \frac{\delta}{\delta(\partial^{-} \mathcal{A}_{a}^{i}(u))} + \eta_{a}^{-}(u) \frac{\delta}{\delta \mathcal{A}_{a}^{-}(u)} + (\partial_{\mu} \eta_{a}^{\mu}(u)) \frac{\delta}{\delta(\partial_{\mu} \mathcal{A}_{a}^{\mu}(u))}$$

- An explicit calculation of $\partial^- \eta^i_{\pm k}$ and $\eta^-_{\pm k}$ shows that these components have an extra $1/k^+$ when $k^+ \to +\infty$
- At leading log, it seems sufficient to consider :

$$\eta(u) \cdot \mathbb{T}_{\boldsymbol{u}} \underset{\text{LLog}}{=} (\partial_{\mu} \eta^{\mu}_{a}(u)) \frac{\delta}{\delta(\partial_{\mu} \mathcal{A}^{\mu}_{a}(u))}$$

This is almost correct, but not quite...



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Leading Log approximation

The region above the LC contains a classical background field of the form

$$\mathcal{A}^{\pm} = 0$$
 , $\mathcal{A}^{i} = \frac{i}{g} \Omega^{\dagger} \partial^{i} \Omega$

by the interaction of the fluctuation with a background field can turn terms that are not divergent on the LC into divergent terms ! (factors of k^+ can arise in the 3-gluon derivative coupling)

Because this background is a pure gauge, this problem is easily circumvented by using $[\Omega \eta]_a$ instead of η_a :

$$\eta(u) \cdot \mathbb{T}_{u} \equiv (\partial^{-}[\Omega\eta]_{b}^{i}(u)) \frac{\delta}{\delta(\partial^{-}[\Omega\mathcal{A}]_{b}^{i}(u))} + [\Omega\eta]_{b}^{-}(u) \frac{\delta}{\delta[\Omega\mathcal{A}]_{b}^{-}(u)} + (\partial_{\mu}[\Omega\eta]_{b}^{\mu}(u)) \frac{\delta}{\delta(\partial_{\mu}[\Omega\mathcal{A}]_{b}^{\mu}(u))}$$

 \triangleright at leading log, only the last term matters



JIMWLK Hamiltonian

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• Derivatives with respect to $\partial_{\mu}[\Omega \mathcal{A}]^{\mu}_{b}(u)$ can be mapped to derivatives with respect to the slowest color sources :

$$\int du^{+} \frac{\delta}{\delta(\partial_{\mu}[\Omega \mathcal{A}]^{\mu}_{b}(u))} = \int d^{2}\vec{x}_{\perp} \langle \vec{u}_{\perp} | \frac{1}{\partial_{\perp}^{2}} | \vec{x}_{\perp} \rangle \frac{\delta}{\delta \widetilde{\mathcal{A}}^{+}_{a}(\epsilon, \vec{x}_{\perp})}$$

with $-\partial_{\perp}^{2} \widetilde{\mathcal{A}}^{+}(\epsilon, \vec{x}_{\perp}) = \rho(\epsilon, \vec{x}_{\perp})$

• When \vec{u}, \vec{v} are on the same branch of the LC, we have

$$\frac{1}{2} \int \mathcal{G}(\vec{\boldsymbol{u}}, \vec{\boldsymbol{v}}) \mathbb{T}_{\boldsymbol{u}} \mathbb{T}_{\boldsymbol{v}} = \frac{1}{2} \log \left(\frac{\Lambda^{+}}{p^{+}} \right) \int \eta_{ab}(\vec{\boldsymbol{x}}_{\perp}, \vec{\boldsymbol{y}}_{\perp}) \frac{\delta^{2}}{\delta \widetilde{\mathcal{A}}_{a}^{+}(\epsilon, \vec{\boldsymbol{x}}_{\perp}) \delta \widetilde{\mathcal{A}}_{b}^{+}(\epsilon, \vec{\boldsymbol{y}}_{\perp})}$$

$$\begin{array}{l} \text{with} \hspace{0.2cm} \eta_{ab}(\vec{\boldsymbol{x}}_{\perp},\vec{\boldsymbol{y}}_{\perp}) \equiv \frac{1}{\pi} \int \frac{d^{2}\vec{\boldsymbol{z}}_{\perp}}{(2\pi)^{2}} \frac{(\vec{\boldsymbol{x}}_{\perp}-\vec{\boldsymbol{z}}_{\perp}) \cdot (\vec{\boldsymbol{y}}_{\perp}-\vec{\boldsymbol{z}}_{\perp})}{(\vec{\boldsymbol{x}}_{\perp}-\vec{\boldsymbol{z}})^{2} (\vec{\boldsymbol{y}}_{\perp}-\vec{\boldsymbol{z}}_{\perp})^{2}} \\ \times \Big[1 + \Omega(x) \Omega^{\dagger}(y) - \Omega(x) \Omega^{\dagger}(z) - \Omega(z) \Omega^{\dagger}(y) \Big]_{ab} \end{array}$$



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By using the Green's formula for β^{μ} , one can show that

$$\int_{\vec{\boldsymbol{u}}\in\mathrm{LC}}\boldsymbol{\beta}(\vec{\boldsymbol{u}}) \mathbb{T}_{\boldsymbol{u}} \stackrel{=}{=} \frac{1}{2} \log\left(\frac{\Lambda^{+}}{p^{+}}\right) \int_{\vec{\boldsymbol{x}}_{\perp}} \left(\int_{\vec{\boldsymbol{y}}_{\perp}} \frac{\delta\eta_{ab}(\vec{\boldsymbol{x}}_{\perp}, \vec{\boldsymbol{y}}_{\perp})}{\delta\tilde{\mathcal{A}}_{b}^{+}(\epsilon, \vec{\boldsymbol{y}}_{\perp})}\right) \frac{\delta}{\delta\tilde{\mathcal{A}}_{a}^{+}(\epsilon, \vec{\boldsymbol{x}}_{\perp})}$$

Combining the real and virtual terms :

$$\frac{1}{2} \int_{\vec{u},\vec{v}\in LC} \mathcal{G}(\vec{u},\vec{v}) \mathbb{T}_{u} \mathbb{T}_{v} + \int_{\vec{u}\in LC} \mathcal{\beta}(\vec{u}) \mathbb{T}_{u} \Big]$$

$$= \log \left(\frac{\Lambda^{+}}{p^{+}}\right) \underbrace{\frac{1}{2} \int_{\vec{y}_{\perp}} \frac{\delta}{\delta \widetilde{\mathcal{A}}_{b}^{+}(\epsilon,\vec{y}_{\perp})} \eta_{ab}(\vec{x}_{\perp},\vec{y}_{\perp}) \frac{\delta}{\delta \widetilde{\mathcal{A}}_{a}^{+}(\epsilon,\vec{x}_{\perp})}}_{JIMWLK \mathcal{H}}$$

(Note : \mathcal{H} is Hermitian)



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The configuration where \vec{u}, \vec{v} are on the first branch of the LC can be rewritten as

$$\frac{dN}{d^{3}\vec{\boldsymbol{p}}}\Big|_{\rm NLO} \stackrel{=}{=} \log\left(\frac{\Lambda^{+}}{p^{+}}\right) \mathcal{H}_{1} \left.\frac{dN}{d^{3}\vec{\boldsymbol{p}}}\right|_{\rm LO}$$

with \mathcal{H}_1 the JIMWLK Hamiltonian for the first nucleus

Including also the configuration where both \vec{u}, \vec{v} are on the second branch of the LC, we get

$$\frac{dN}{d^{3}\vec{\boldsymbol{p}}}\Big|_{_{\rm NLO}} \stackrel{=}{_{\rm LLog}} \left[\log\left(\frac{\Lambda^{+}}{p^{+}}\right)\mathcal{H}_{1} + \log\left(\frac{\Lambda^{-}}{p^{-}}\right)\mathcal{H}_{2}\right] \left.\frac{dN}{d^{3}\vec{\boldsymbol{p}}}\right|_{_{\rm LO}}$$



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The only remaining possibility is to have \vec{u} and \vec{v} on different branches of the LC



However, there is no log divergence in this case, since the k^+ integral is of the form :

$$\int \frac{dk^+}{k^+} \cdots e^{ik^+(u^- - v^-)} e^{ik^-(u^+ - v^+)}$$

no mixing of the divergences of the two nuclei



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All the above discussion is for given sources $\rho_{1,2}$ (or given fields $\widetilde{\mathcal{A}}_{1,2}^{\pm}$). Averaging over all the configurations of the sources in the two projectiles, and using the hermiticity of \mathcal{H} , we obtain

$$\left\langle \frac{dN}{d^{3}\vec{\boldsymbol{p}}} \right\rangle_{\text{LO+NLO}} \stackrel{=}{\underset{\text{LLog}}{=}} \int \left[D\widetilde{\mathcal{A}}_{1}^{+} D\widetilde{\mathcal{A}}_{2}^{-} \right] \\ \times \left(\left[1 + \log\left(\frac{\Lambda^{+}}{p^{+}}\right) \mathcal{H}_{1} + \log\left(\frac{\Lambda^{-}}{p^{-}}\right) \mathcal{H}_{2} \right] W[\widetilde{\mathcal{A}}_{1}^{+}] W[\widetilde{\mathcal{A}}_{2}^{-}] \right) \left. \frac{dN}{d^{3}\vec{\boldsymbol{p}}} \right|_{\text{LO}}$$

This is a 1-loop result. Using RG arguments, we get the following factorized formula for the resummation of the leading log terms to all orders :

$$\left\langle \frac{dN}{d^3 \vec{p}} \right\rangle_{\text{LLog}} = \int \left[D \widetilde{\mathcal{A}}_1^+ D \widetilde{\mathcal{A}}_2^- \right] W_{Y_1} [\widetilde{\mathcal{A}}_1^+] W_{Y_2} [\widetilde{\mathcal{A}}_2^-] \left. \frac{dN}{d^3 \vec{p}} \right|_{\text{LO}}$$
with $\frac{\partial}{\partial Y} W_Y = \mathcal{H} W$, $Y_1 = \log(\sqrt{s}/p^+)$, $Y_2 = \log(\sqrt{s}/p^-)$



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Requirements for factorization

The fact that the observable is bilinear in the fields is not essential. The formula

$$\mathcal{O}_{_{\mathrm{NLO}}} = \left[\frac{1}{2} \int_{\vec{u}, \vec{v} \in \mathrm{LC}} \mathcal{G}(\vec{u}, \vec{v}) \, \mathbb{T}_{\boldsymbol{u}} \mathbb{T}_{\boldsymbol{v}} + \int_{\vec{u} \in \mathrm{LC}} \boldsymbol{\beta}(\vec{u}) \, \mathbb{T}_{\boldsymbol{u}}\right] \mathcal{O}_{_{\mathrm{LO}}}$$

can be established for more general observables, provided their expectation value depends on retarded fields only

- Crucial ingredients for factorization :
 - Only connected diagrams contribute
 - One should have an initial value problem
 retarded boundary conditions are essential
 - The observable should involve only one rapidity scale. Otherwise, there are extra large corrections in $\alpha_s(y_1 - y_2)$ that are not captured in the evolution of the $W[\widetilde{\mathcal{A}}^{\pm}]$'s



Quantities that do factorize

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Energy-momentum tensor $T^{\mu
u}(au,\eta,ec{x}_{\perp})$:

$$\begin{split} \langle T^{\mu\nu}(\tau,\eta,\vec{x}_{\perp}) \rangle &= \int \left[D\widetilde{\mathcal{A}}_{1}^{+} D\widetilde{\mathcal{A}}_{2}^{-} \right] W_{Y_{1}}[\widetilde{\mathcal{A}}_{1}^{+}] W_{Y_{2}}[\widetilde{\mathcal{A}}_{2}^{-}] \left[T^{\mu\nu}(\tau,\eta,\vec{x}_{\perp}) \right]_{\text{LO}} \\ \text{with} \quad Y_{1} = \ln(\sqrt{s}) - \eta \quad , \quad Y_{2} = \ln(\sqrt{s}) + \eta \end{split}$$

- CGC initial conditions for hydrodynamics
 Note : this cannot be used for studying fluctuations
- Higher moments of the multiplicity distribution in a small slice of rapidity. These moments are expressible in terms of retarded quantities (FG, Venugopalan)
- For some quantities, an extension of the above form of factorization may be able to resum all the leading logs. Example : 2-gluon correlations with a large rapidity separation between the gluons (work in progress with T. Lappi and R. Venugopalan)



Quantities that do not factorize

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- More exclusive quantities seem out of reach of this form of factorization :
 - Example : survival probability of rapidity gaps
- ▷ for such quantities, the main obstruction is the impossibility to write them in terms of retarded objects

The problem is not factorization (which should follow from causality to a large extent) per se, but that our description of the wavefunction of the incoming projectiles does not contain enough information to answer the question we are asking

 $(W[\widetilde{\mathcal{A}}]$ is only the diagonal part of the initial density matrix of the incoming nucleus)



Summary

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- Factorization works (at leading log) in the saturation regime of nucleus-nucleus collisions : all the leading logarithms of 1/x_{1,2} can be absorbed into the evolution of the distribution of color sources of the corresponding nucleus
- Restriction : the observable must be sufficiently inclusive (so that it can be expressed in terms of fields with retarded boundary conditions)
- The proof becomes straightforward once one has rewritten the observable in a way that exhibits the causal nature of the involved fields
- Extensions :
 - The "non leading log" terms still contain pieces that trigger the Weibel instability > resummation ?
 - Factorization for the inclusive 2-gluon spectrum
 - Factorization in an exclusive quantity