

# Relating neutrino masses, LFV and SUSY breaking



F. R. Joaquim  
(Univ. di Padova and INFN Padua)



Work done in collaboration with Anna Rossi ([hep-ph/0604083](https://arxiv.org/abs/hep-ph/0604083) and [hep-ph/0606xxx](https://arxiv.org/abs/hep-ph/0606xxx))

## PLAN:

- Predictions for LFV processes in the triplet seesaw mechanism
- Sensitivity to low-energy neutrino parameters
- General analysis of the parameter space
- Conclusions


## A kind of introduction...


- SUSY breaking is communicated to the observable sector via Gauge and Yukawa interactions


$$W = \frac{1}{\sqrt{2}}(\mathbf{Y}_T L T L + \mathbf{Y}_S d^c S d^c) + \mathbf{Y}_Z d^c Z L + \mathbf{Y}_d d^c Q H_1 + \mathbf{Y}_e e^c L H_1 + \mathbf{Y}_u u^c Q H_2 \\ + \frac{1}{\sqrt{2}}\lambda H_2 \bar{T} H_2 + M_T T \bar{T} + M_Z Z \bar{Z} + M_S S \bar{S} + \mu H_1 H_2$$

- Soft SUSY breaking terms generated at one and two-loop level (finite contributions at the messenger decoupling scale)

$$-\mathcal{L}_{\text{soft}} \supset (H_1 \tilde{e}^c \mathbf{A}_e \tilde{L} + H_1 \tilde{d}^c \mathbf{A}_d \tilde{Q} - H_2 \tilde{u}^c \mathbf{A}_u \tilde{Q} + \frac{1}{2} M_a \lambda_a \lambda_a + B_H \mu H_1 H_2 + \text{h.c.}) + \\ \tilde{L}^\dagger \mathbf{m}_{\tilde{L}}^2 \tilde{L} + \tilde{e}^c \mathbf{m}_{\tilde{e}^c}^2 \tilde{e}^{c*} + \tilde{d}^c \mathbf{m}_{\tilde{d}^c}^2 \tilde{d}^{c*} + \tilde{u}^c \mathbf{m}_{\tilde{u}^c}^2 \tilde{u}^{c*} + \tilde{Q}^\dagger \mathbf{m}_{\tilde{Q}}^2 \tilde{Q} + m_{H_1}^2 H_1^\dagger H_1 + m_{H_2}^2 H_2^\dagger H_2$$

$M_T, B_T, \lambda$    $m_{H_2}^2|_{M_T}, m_{H_1}^2|_{M_T}, B_H|_{M_T}$  and  $\mathbf{m}_X^2|_{M_T}, \mathbf{A}_X|_{M_T}, M_a|_{M_T}$

  $\mu_{\text{EWSB}}$




$$|\mu|^2 = \frac{1}{\cos(2\beta)}(m_{H_2}^2 \sin^2 \beta - m_{H_1}^2 \cos^2 \beta) - \frac{1}{2} m_Z^2, \quad B_H^{\text{EW}} \mu = \frac{1}{2} \sin(2\beta)(m_{H_1}^2 + m_{H_2}^2 + 2|\mu|^2)$$

tan  $\beta$  scanning and matching between  $B_H^{\text{EW}}$  and  $B_H$  required  $\longrightarrow$  tan  $\beta$ , sgn( $\mu$ ) and  $|\mu|$  fixed

Typically the dominant contributions to the off-diagonal slepton masses are:  $(\mathbf{m}_{\tilde{L}}^2)_{ij} \propto \frac{B_T^2}{(16\pi^2)^2} (\mathbf{Y}_T^\dagger \mathbf{Y}_T)_{ij}$

Direct relation with the effective neutrino mass matrix:  $(\mathbf{m}_{\tilde{L}}^2)_{ij} \propto (\mathbf{Y}_T^\dagger \mathbf{Y}_T)_{ij} \sim \left(\frac{M_T}{\lambda v_2^2}\right)^2 (\mathbf{m}_\nu^\dagger \mathbf{m}_\nu)_{ij}, i \neq j$

In terms of the PMNS mixing matrix and neutrino masses:

  $\mathbf{U} \equiv \mathbf{U}^\delta \text{diag}(e^{i\phi_1}, e^{i\phi_2}, 1), (\mathbf{m}_\nu^\dagger \mathbf{m}_\nu)_{ij} = \mathbf{U}_{ik}^\delta m_k^2 \mathbf{U}_{kj}^{\delta\dagger}$

Normal ordering:  
 $(m_1 < m_2 < m_3)$

$$m_2^2 = m_1^2 + \Delta m_\odot^2$$

$$m_3^2 = m_1^2 + \Delta m_\odot^2 + \Delta m_{\text{atm}}^2$$

$$(\mathbf{m}_\nu^\dagger \mathbf{m}_\nu)_{ij} = m_1^2 \delta_{ij} + \Delta m_\odot^2 \mathbf{U}_{i2} \mathbf{U}_{j2}^* + (\Delta m_{\text{atm}}^2 + \Delta m_\odot^2) \mathbf{U}_{i3} \mathbf{U}_{j3}^*$$

Inverted ordering:  
 $(m_3 < m_1 < m_2)$

$$m_2^2 = m_3^2 + \Delta m_{\text{atm}}^2$$

$$m_1^2 = m_3^2 - \Delta m_\odot^2 + \Delta m_{\text{atm}}^2$$

$$(\mathbf{m}_\nu^\dagger \mathbf{m}_\nu)_{ij} = m_3^2 \delta_{ij} + \Delta m_{\text{atm}}^2 \mathbf{U}_{i2} \mathbf{U}_{j2}^* + (\Delta m_{\text{atm}}^2 - \Delta m_\odot^2) \mathbf{U}_{i1} \mathbf{U}_{j1}^*$$

For  $i \neq j$ ,  $(\mathbf{m}_\nu^\dagger \mathbf{m}_\nu)_{ij}$  does not depend on the lightest neutrino mass.

Normal ordering:  $m_1 < m_2 < m_3$

$$\rho \equiv \frac{\Delta m_{\odot}^2}{\Delta m_{\text{atm}}^2}$$

$$\left\{ \begin{array}{l} (\mathbf{m}_{\tilde{L}}^2)_{\mu e} \propto c_{13} [\rho c_{23} \sin(2\theta_{12}) + 2 s_{13} s_{23}] \\ (\mathbf{m}_{\tilde{L}}^2)_{\tau\mu} \propto c_{13}^2 \sin(2\theta_{23}) \\ (\mathbf{m}_{\tilde{L}}^2)_{\tau e} \propto c_{13} [2(1 + \rho c_{12}^2) s_{13} c_{23} - \rho s_{23} \sin(2\theta_{12})] \end{array} \right. \xrightarrow{\text{green arrow}} \left\{ \begin{array}{l} \left. \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau\mu}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|_{s_{13} \ll \rho/2} \simeq \frac{\Delta m_{\text{atm}}^2}{\Delta m_{\odot}^2} \frac{\sin(2\theta_{23})}{c_{23} \sin(2\theta_{12})} \simeq 40 \\ \left. \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau e}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|_{s_{13} \ll \rho/2} \simeq -\tan \theta_{23} \simeq -1 \end{array} \right.$$

Inverted ordering:  $m_3 < m_1 < m_2$

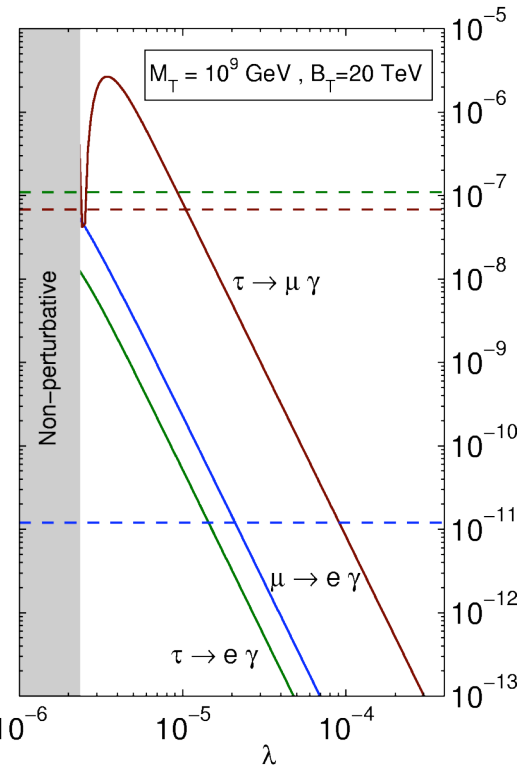
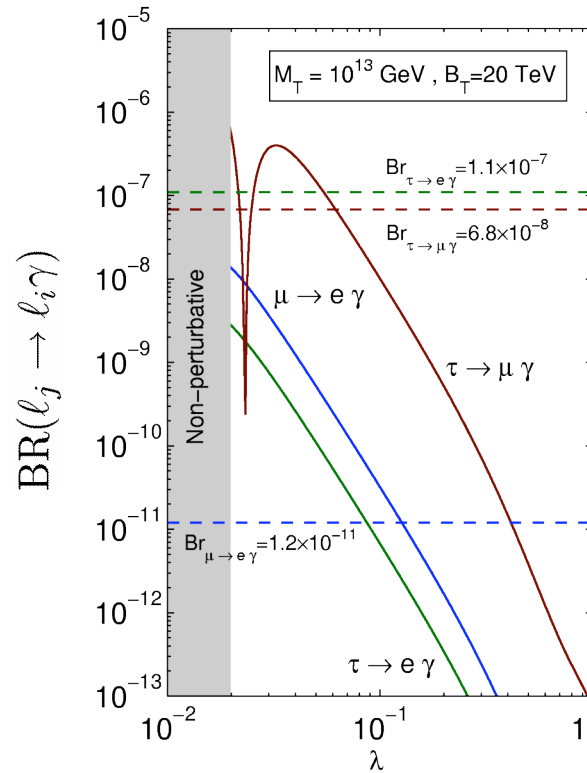
$$\left\{ \begin{array}{l} (\mathbf{m}_{\tilde{L}}^2)_{\mu e} \propto c_{13} [\rho c_{23} \sin(2\theta_{12}) - 2(1 - \rho c_{12}^2) s_{13} s_{23}] \\ (\mathbf{m}_{\tilde{L}}^2)_{\tau\mu} \propto -c_{13}^2 \sin(2\theta_{23}) \\ (\mathbf{m}_{\tilde{L}}^2)_{\tau e} \propto -c_{13} [\rho s_{23} \sin(2\theta_{12}) + 2 s_{13} c_{23}] \end{array} \right. \xrightarrow{\text{green arrow}} \left\{ \begin{array}{l} \left. \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau\mu}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|_{s_{13} \ll \rho/2} \simeq -\frac{\Delta m_{\text{atm}}^2}{\Delta m_{\odot}^2} \frac{\sin(2\theta_{23})}{c_{23} \sin(2\theta_{12})} \simeq -40 \\ \left. \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau e}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|_{s_{13} \ll \rho/2} \simeq -\tan \theta_{23} \simeq -1 \end{array} \right.$$

# Predictions for $\ell_j \rightarrow \ell_i \gamma$

$$\frac{\text{BR}(\tau \rightarrow \mu \gamma)}{\text{BR}(\mu \rightarrow e \gamma)} \simeq \left| \frac{(m_L^2)_{\tau\mu}}{(m_L^2)_{\mu e}} \right|^2 \frac{\text{BR}(\tau \rightarrow \mu \nu_\tau \bar{\nu}_\mu)}{\text{BR}(\mu \rightarrow e \nu_\mu \bar{\nu}_e)}$$

$$\frac{\text{BR}(\tau \rightarrow e \gamma)}{\text{BR}(\mu \rightarrow e \gamma)} \simeq \left| \frac{(m_L^2)_{\tau e}}{(m_L^2)_{\mu e}} \right|^2 \frac{\text{BR}(\tau \rightarrow e \nu_\tau \bar{\nu}_e)}{\text{BR}(\mu \rightarrow e \nu_\mu \bar{\nu}_e)}$$

$\theta_{13} = 0$



$\theta_{13} = 0$

$$\frac{\text{BR}(\tau \rightarrow \mu \gamma)}{\text{BR}(\mu \rightarrow e \gamma)} \sim 300$$

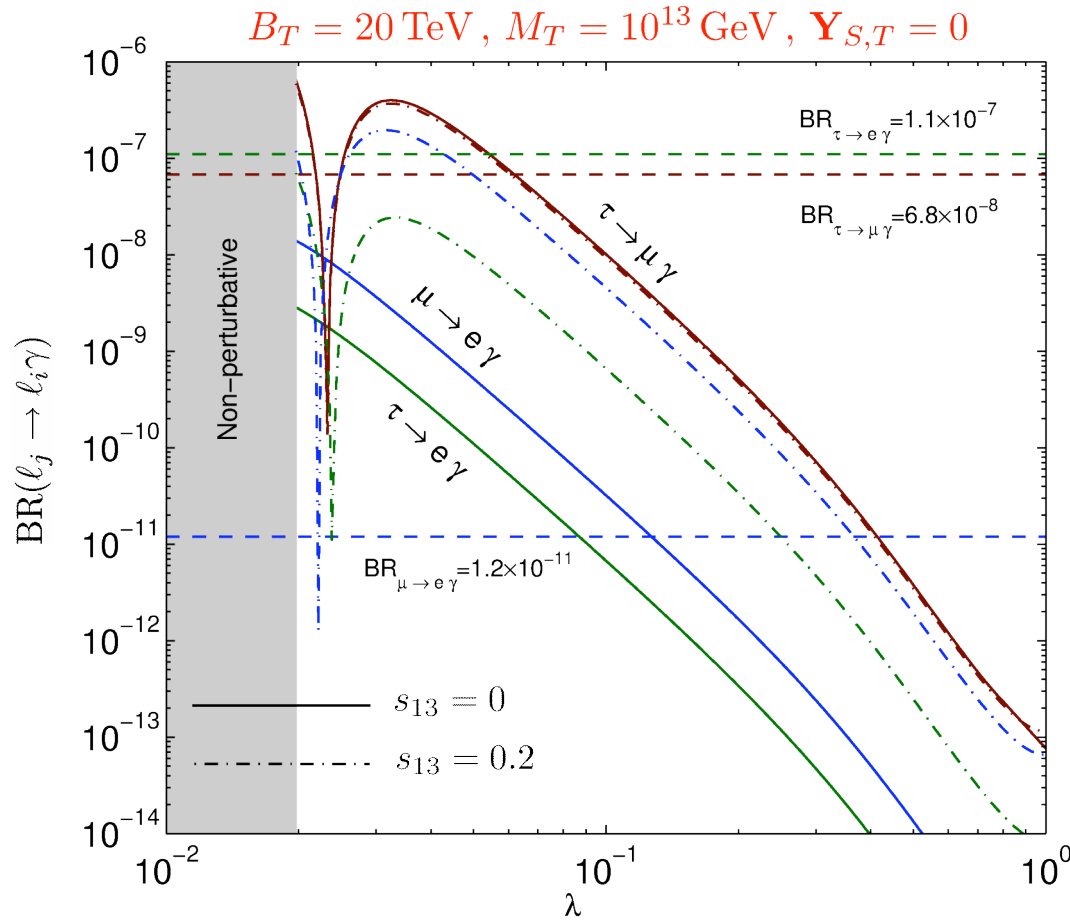
$$\text{BR}(\mu \rightarrow e \gamma) < 1.1 \times 10^{-11} \quad [\text{MEGA}'99]$$

$$\text{BR}(\tau \rightarrow \mu \gamma) < 6.8 \times 10^{-8} \quad [\text{BaBar}'05]$$

$$\text{BR}(\tau \rightarrow e \gamma) < 1.1 \times 10^{-7} \quad [\text{BaBar}'05]$$

$$\frac{\text{BR}(\tau \rightarrow e \gamma)}{\text{BR}(\mu \rightarrow e \gamma)} \sim 0.2$$

# Sensitivity to 1-3 neutrino mixing



$$\frac{\text{BR}(\tau \rightarrow e\gamma)}{\text{BR}(\mu \rightarrow e\gamma)} \simeq \left| \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau e}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|^2 \frac{\text{BR}(\tau \rightarrow e\nu_\tau\bar{\nu}_e)}{\text{BR}(\mu \rightarrow e\nu_\mu\bar{\nu}_e)}$$

$$\frac{\text{BR}(\tau \rightarrow \mu\gamma)}{\text{BR}(\mu \rightarrow e\gamma)} \simeq \left| \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau\mu}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|^2 \frac{\text{BR}(\tau \rightarrow \mu\nu_\tau\bar{\nu}_\mu)}{\text{BR}(\mu \rightarrow e\nu_\mu\bar{\nu}_e)}$$

For  $s_{13} = 0.2$  :

$$\left| \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau\mu}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|_{\theta_{13}=0.2}^2 \simeq 9.0, \quad \left| \frac{(\mathbf{m}_{\tilde{L}}^2)_{\tau e}}{(\mathbf{m}_{\tilde{L}}^2)_{\mu e}} \right|_{\theta_{13}=0.2}^2 \simeq 0.7$$

$$(\mathbf{m}_{\tilde{L}}^2)_{\mu e} \propto c_{13} [\rho c_{23} \sin(2\theta_{12}) + 2s_{13}s_{23}]$$

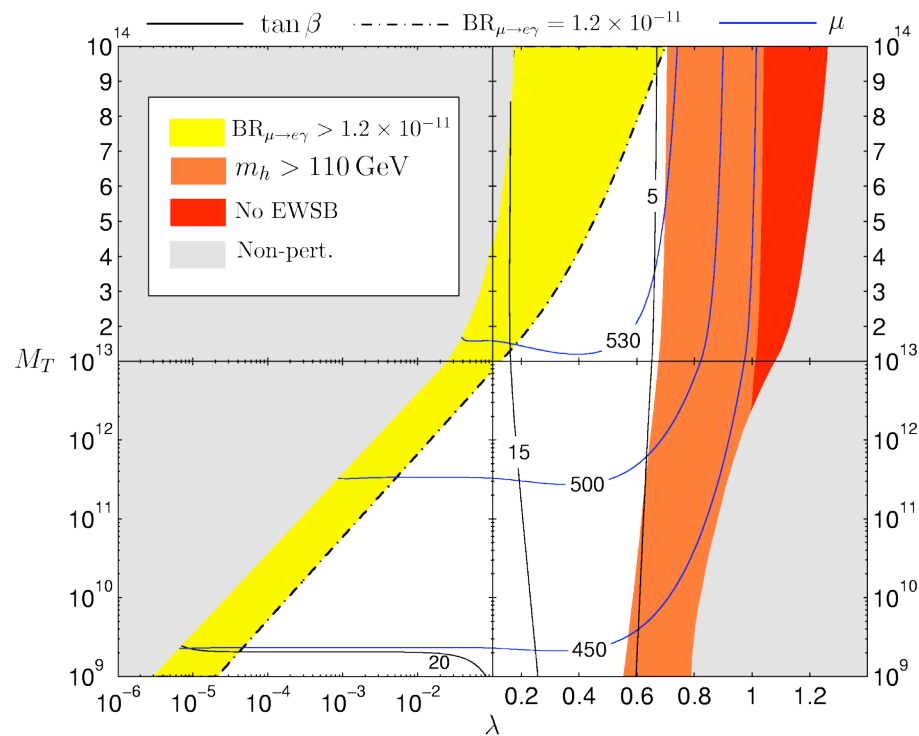
$$(\mathbf{m}_{\tilde{L}}^2)_{\tau\mu} \propto c_{13}^2 \sin(2\theta_{23})$$

$$(\mathbf{m}_{\tilde{L}}^2)_{\tau e} \propto c_{13} [2(1 + \rho c_{12}^2)s_{13}c_{23} - \rho s_{23} \sin(2\theta_{12})]$$

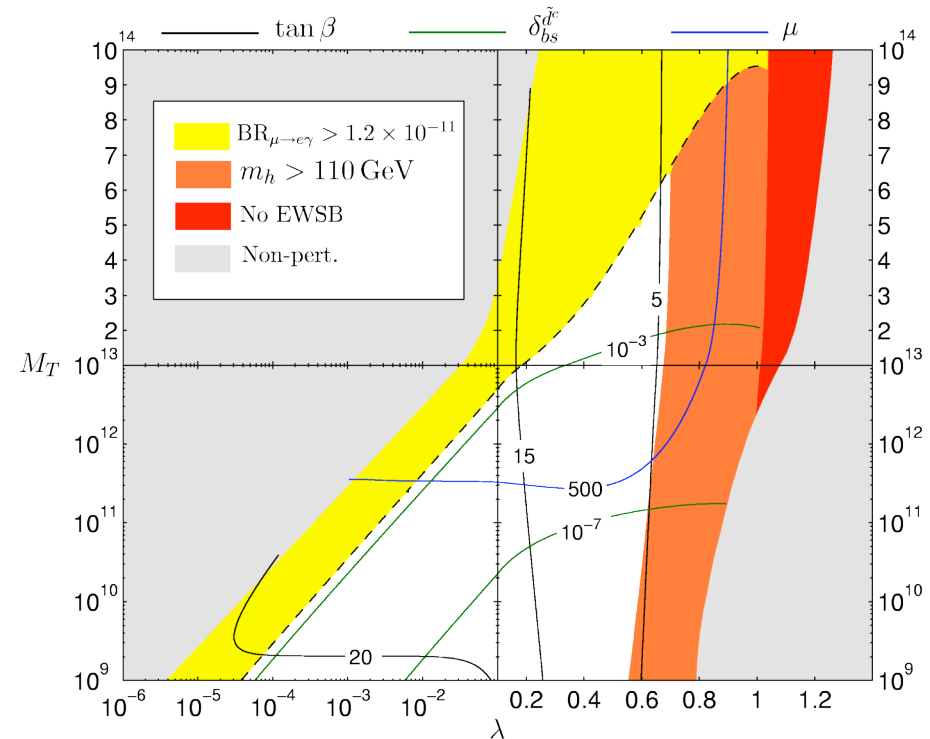
$$\left. \frac{\text{BR}(\tau \rightarrow \mu\gamma)}{\text{BR}(\mu \rightarrow e\gamma)} \right|_{s_{13}=0.2} \simeq 2.0$$

$$\left. \frac{\text{BR}(\tau \rightarrow e\gamma)}{\text{BR}(\mu \rightarrow e\gamma)} \right|_{s_{13}=0.2} \simeq 0.1$$

$$B_T = 20 \text{ TeV}, \mathbf{Y}_S = \mathbf{Y}_Z = 0$$



$$B_T = 20 \text{ TeV}, \mathbf{Y}_S = \mathbf{Y}_Z = \mathbf{Y}_T$$



## Conclusions:

- ➔ Very predictive framework: strict relation between LFV, REWSB and low-energy neutrino physics.
- ➔ More information from low-energy neutrino physics needed: neutrino masses, 1-3 mixing angle, (leptonic CP-violation)
- ➔ Information about the SUSY spectrum, Higgs mass, etc...