

Gravity Scattering Amplitudes

Amplitudes 2014

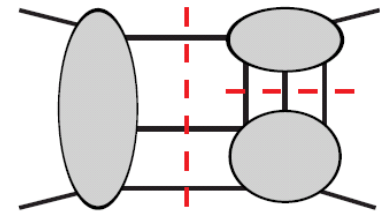
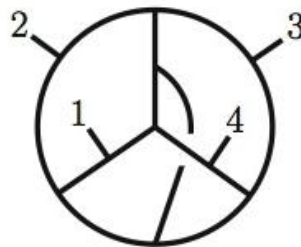
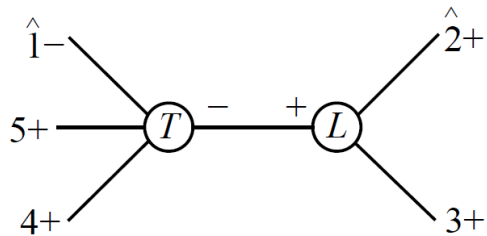
June 13, 2014

Zvi Bern, UCLA

Review + new work:

ZB, Scott Davies, Tristan Dennen, Sasha Smirnov, Volodya Smirnov: [arXiv:1309.2498](https://arxiv.org/abs/1309.2498)

ZB, Scott Davies, Josh Nohle, [arXiv:1405.1015](https://arxiv.org/abs/1405.1015)



Outline

1. Review of some basic properties of gravity amplitudes:

- Various new representations of gravity amplitudes.
- Relations between gravity and gauge theory.
- Duality between color and kinematics.
- Tame UV behavior of $N = 8$ supergravity.

2. Some recent developments discussed here:

- New surprises in UV properties in $N = 4$ supergravity.
- Subleading soft limits and extended BMS symmetry. **Also talk from Huang**

3. Other talks on gravity amplitudes:

- String theory multiloop amplitudes. **talk from Green**
- New extensions of color kinematics duality. **talks from Johansson, Roiban**
- Scattering equations and implications for gravity. **talks from Monteiro, Yuan**
- Causality and 3 graviton amplitudes **talk from Maldacena**

Review of Some Basic Properties

Various Representations of Gravity Amplitudes

There are a plethora of representations of gravity tree amplitude:

- 1) **Classic KLT relations: gravity tree amplitudes from gauge theory.** Kawai, Lewellen, Tye
- 2) **Duality between color and kinematics.** ZB, Carrasco and Johansson
- 3) **Twistor forms.** Hodges; Cachazo and Skinner; Adamo and Mason; Litsey and Stankowicz
- 4) **Ambitwistor string form.** Mason and Skinner
- 5) **Grassmannian forms.** Cachazo and Geyer; Cheung
- 6) **Scattering equations and gravity amplitudes.** Cachazo, He, Yuan
Montiero and O'Connell

Loop level less well developed, but we have powerful tools and many results in gravity theory.

- 1) **Generalized unitarity.** ZB, Dixon, Dunbar, Kosower;
ZB, Dixon, Dunbar, Perelstein, Rozowsky
- 2) **Duality between color and kinematics.** ZB, Carrasco, Johansson
ZB, Dixon, Carrasco, Johansson, Roiban

KLT Relations

At tree level, Kawai, Lewellen and Tye presented a relationship between closed and open string amplitudes. In field theory limit, relationship is between gravity and gauge theory.

$$M_4^{\text{tree}}(1, 2, 3, 4) = s_{12} A_4^{\text{tree}}(1, 2, 3, 4) A_4^{\text{tree}}(1, 2, 4, 3),$$

$$M_5^{\text{tree}}(1, 2, 3, 4, 5) = s_{12} s_{34} A_5^{\text{tree}}(1, 2, 3, 4, 5) A_5^{\text{tree}}(2, 1, 4, 3, 5) \\ + s_{13} s_{24} A_5^{\text{tree}}(1, 3, 2, 4, 5) A_5^{\text{tree}}(3, 1, 4, 2, 5)$$

Gravity amplitude

Color stripped gauge theory amplitude

where we have stripped all coupling constants

Full gauge theory amplitude

$$A_4^{\text{tree}} = g^2 \sum_{\text{non-cyclic}} \text{Tr}(T^{a_1} T^{a_2} T^{a_3} T^{a_4}) A_4^{\text{tree}}(1, 2, 3, 4)$$



Holds for any external states.

Progress in gauge theory can be imported into gravity theories

Gravity and Gauge Theory Amplitudes

$$\begin{aligned}
 M_4^{\text{tree}}(1_h^-, 2_h^-, 3_h^+, 4_h^+) &= \left(\frac{\kappa}{2}\right)^2 s_{12} A_4^{\text{tree}}(1_g^-, 2_g^-, 3_g^+, 4_g^+) \times A_4^{\text{tree}}(1_g^-, 2_g^-, 4_g^+, 3_g^+) \\
 &= \left(\frac{\kappa}{2}\right)^2 s_{12} \frac{\langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle} \times \frac{\langle 12 \rangle^4}{\langle 12 \rangle \langle 24 \rangle \langle 43 \rangle \langle 31 \rangle}
 \end{aligned}$$

gravity

gauge theory

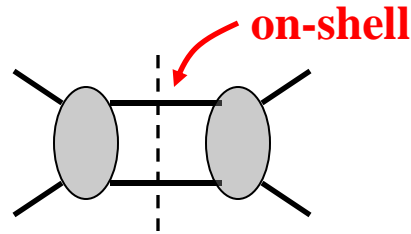
$$\langle jl \rangle = \langle k_j^- | k_l^+ \rangle = \frac{1}{2} \bar{u}(k_j)(1 + \gamma_5)u(k_l) = \sqrt{2k_j \cdot k_l} e^{i\phi}$$

This can be used to generate n -point graviton amplitudes.

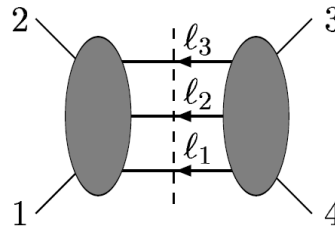
Berends, Giele, Kuijf

Unitarity Method

Two-particle cut:

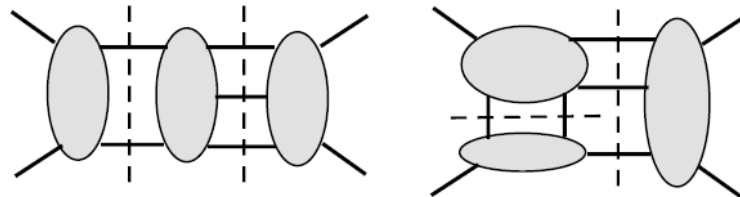


Three-particle cut:



Systematic assembly of complete amplitudes from cuts for any number of particles or loops.

Generalized unitarity as a practical tool:



Different cuts merged to give an expression with correct cuts in all channels.

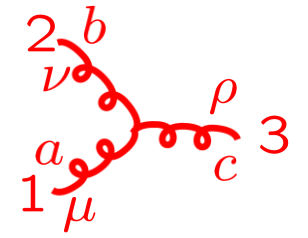
Bern, Dixon and Kosower
 Britto, Cachazo and Feng; Forde;
 Ossala, Pittau, Papadopolous, and many others

Standard general purpose tool for turning tree advances into loop advances

Duality Between Color and Kinematics

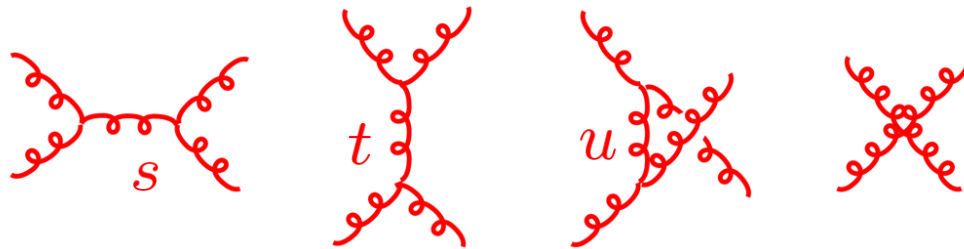
ZB, Carrasco and Johansson

coupling constant \rightarrow $-g$ color factor f^{abc} momentum dependent kinematic factor $(\eta_{\mu\nu}(k_1 - k_2)_\rho + \text{cyclic})$



Color factors based on a Lie algebra: $[T^a, T^b] = i f^{abc} T^c$

Jacobi Identity $f^{a_1 a_2 b} f^{b a_4 a_3} + f^{a_4 a_2 b} f^{b a_3 a_1} + f^{a_4 a_1 b} f^{b a_2 a_3} = 0$



Use $1 = s/s = t/t = u/u$ to assign 4-point diagram to others.

$$\mathcal{A}_4^{\text{tree}} = g^2 \left(\frac{n_s C_s}{s} + \frac{n_t C_t}{t} + \frac{n_u C_u}{u} \right)$$

$$s = (k_1 + k_2)^2 \quad t = (k_1 + k_4)^2 \quad u = (k_1 + k_3)^2$$

**Color factors satisfy Jacobi identity:
 Numerator factors satisfy similar identity:**

$$C_u = C_s - C_t$$

$$n_u = n_s - n_t$$

Color and kinematics satisfy the same identity

Duality Between Color and Kinematics

ZB, Carrasco, Johansson

Consider five-point tree amplitude:

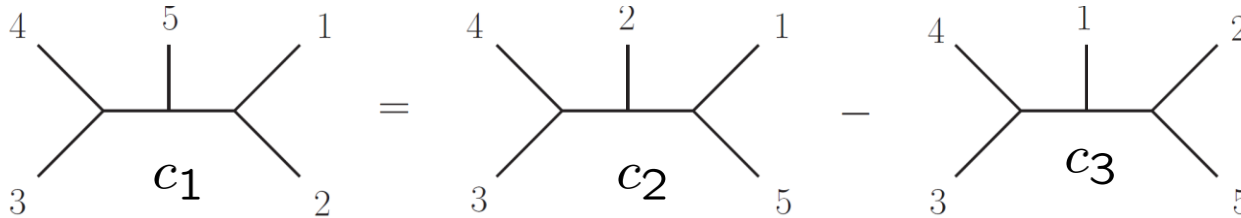
gauge theory $\mathcal{A}_5^{\text{tree}} = \sum_{i=1}^{15} \frac{c_i n_i}{\prod_{\alpha_i} p_{\alpha_i}^2}$

sum is over diagrams (blue arrow pointing to the sum)

color factor (red arrow pointing to c_i)

kinematic numerator factor (red arrow pointing to n_i)

Feynman propagators (red arrow pointing to the denominator)



$$c_1 \equiv f^{a_3 a_4 b} f^{b a_5 c} f^{c a_1 a_2}, \quad c_2 \equiv f^{a_3 a_4 b} f^{b a_2 c} f^{c a_1 a_5}, \quad c_3 \equiv f^{a_3 a_4 b} f^{b a_1 c} f^{c a_2 a_5}$$

$$n_i \sim k_4 \cdot k_5 k_2 \cdot \varepsilon_1 \varepsilon_2 \cdot \varepsilon_3 \varepsilon_4 \cdot \varepsilon_5 + \dots$$

$$c_1 - c_2 + c_3 = 0 \Leftrightarrow n_1 - n_2 + n_3 = 0$$

Claim: At n-points we can always find a rearrangement where color and kinematics satisfy the same algebraic constraint equations.

Nontrivial constraints on amplitudes in field theory and string theory

BCJ, Bjerrum-Bohr, Feng, Damgaard, Vanhove, ; Mafra, Stieberger, Schlotterer; Cachazo; Tye and Zhang; Feng, Huang, Jia; Chen, Du, Feng; Du, Feng, Fu; Naculich, Nastase, Schnitzer

Gravity and Gauge Theory

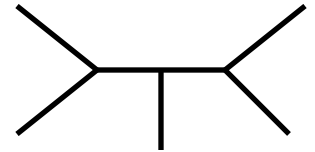
gauge theory: $\frac{1}{g^{n-2}} \mathcal{A}_n^{\text{tree}}(1, 2, 3, \dots, n) = \sum_i \frac{n_i c_i}{\prod_{\alpha_i} p_{\alpha_i}^2}$ **sum over diagrams with only 3 vertices**

kinematic numerator \curvearrowright n_i c_i \curvearrowleft color factor

$$c_i \sim f^{a_1 a_2 b_1} f^{b_1 b_2 a_5} f^{b_2 a_4 a_5}$$

Assume we have:

$$c_1 + c_2 + c_3 = 0 \iff n_1 + n_2 + n_3 = 0$$



Then: $c_i \Rightarrow \tilde{n}_i$ **kinematic numerator of second gauge theory**

Proof: ZB, Dennen, Huang, Kiermaier

gravity: $-i \left(\frac{2}{\kappa}\right)^{(n-2)} \mathcal{M}_n^{\text{tree}}(1, 2, \dots, n) = \sum_i \frac{n_i \tilde{n}_i}{\prod_{\alpha_i} p_{\alpha_i}^2}$ **Encodes KLT tree relations**

Gravity numerators are a double copy of gauge-theory ones.

This works for ordinary Einstein gravity and susy versions.

Cries out for a unified description of the sort given by string theory!

Gravity From Gauge Theory

BCJ

$$-i \left(\frac{2}{\kappa} \right)^{(n-2)} \mathcal{M}_n^{\text{tree}}(1, 2, \dots, n) = \sum_i \frac{n_i \tilde{n}_i}{\prod_{\alpha_i} p_{\alpha_i}^2}$$

n \tilde{n}

$N = 8$ sugra: $(N = 4 \text{ sYM}) \times (N = 4 \text{ sYM})$

$N = 4$ sugra: $(N = 4 \text{ sYM}) \times (N = 0 \text{ sYM})$

**Theory controlled by simple tensor product of YM theories.
Recent papers show more sophisticated cases.**

Cases with fewer supersymmetries.

Carrasco, Chiodaroli, Günaydin and Roiban; ZB, Davies, Dennen, Huang and Nohle

Bagger, Lambert, Gustavsson (BLG) theory and ABJM theories.

Bargheer, He and McLoughlin; Huang and Johansson; Allic Sivaramakrishnan

Fundamental representation matter.

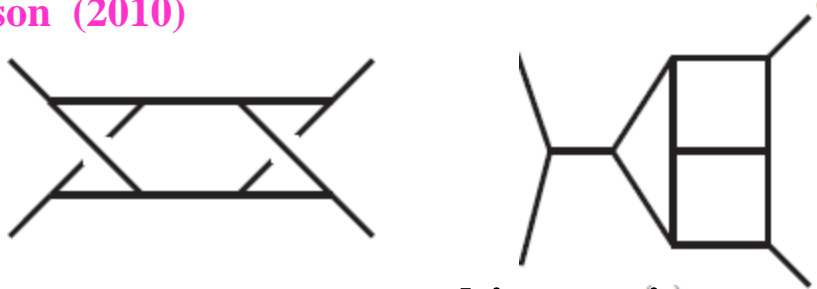
See Johansson's talk

Applications to gauged supergravity.

See Roiban's talk

Loop-Level Conjecture

ZB, Carrasco, Johansson (2010)



$$c_i + c_j + c_k = 0$$

$$n_i + n_j + n_k = 0$$

sum is over diagrams

kinematic numerator

color factor

gauge theory

propagators

gravity

symmetry factor

$$\frac{(-i)^L}{g^{n-2+2L}} \mathcal{A}_n^{\text{loop}} = \sum_j \int \prod_{l=1}^L \frac{d^D p_l}{(2\pi)^D} \frac{1}{S_j} \frac{n_j c_j}{\prod_{\alpha_j} p_{\alpha_j}^2}$$

$$\frac{(-i)^{L+1}}{(\kappa/2)^{n-2+2L}} \mathcal{M}_n^{\text{loop}} = \sum_j \int \prod_{l=1}^L \frac{d^D p_l}{(2\pi)^D} \frac{1}{S_j} \frac{n_j \tilde{n}_j}{\prod_{\alpha_j} p_{\alpha_j}^2}$$

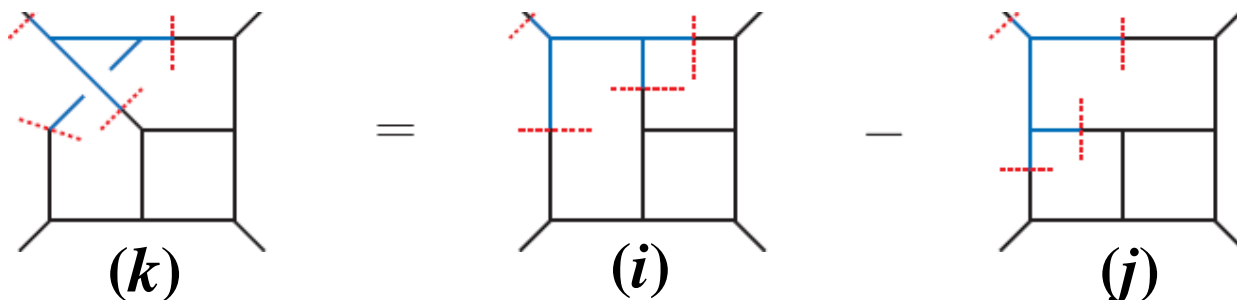
Loop-level is identical to tree-level one except for symmetry factors and loop integration.

This works if numerator satisfies duality.

Gravity integrands are free!

BCJ

Ideas generalize to loops:



color factor

$$C_k = C_i - C_j$$

$$n_k = n_i - n_j$$

kinematic numerator

If you have a set of duality satisfying numerators.

To get:

gauge theory \rightarrow gravity theory

simply take

color factor \rightarrow kinematic numerator

$$C_k \rightarrow n_k$$

Gravity loop integrands are trivial to obtain once we have gauge theory in a form where duality works.

One diagram numerator to rule them all

ZB, Carrasco, Johansson (2010)

$N = 4$ super-Yang-Mills integrand

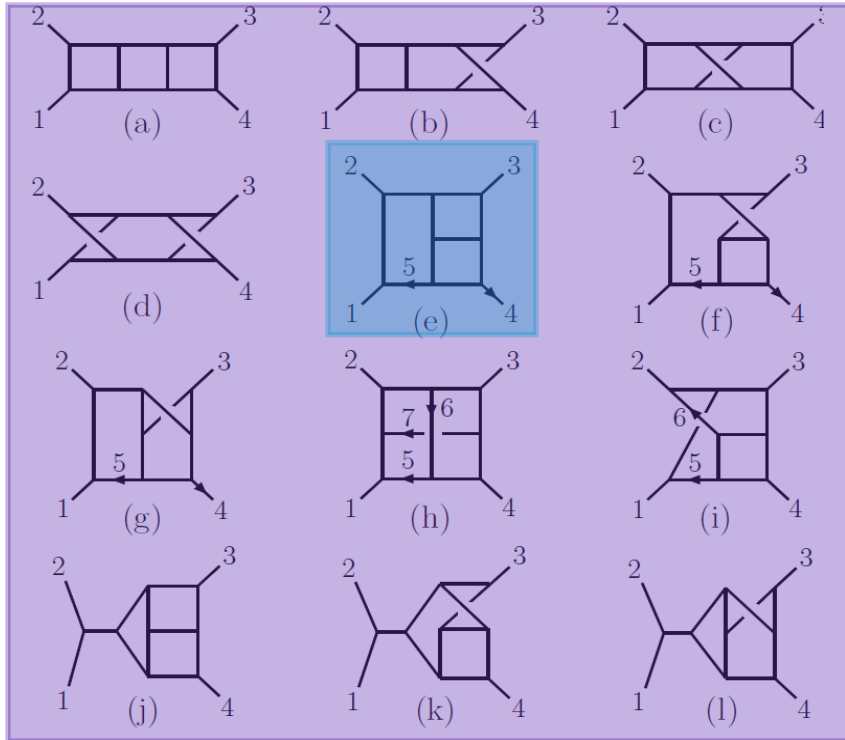


Diagram (e) is the master diagram.

Determine the master integrand in proper form and duality gives all others.

$N = 8$ sugra given by replacing color with kinematic numerators.

Integral $I^{(x)}$	$\mathcal{N} = 4$ Super-Yang-Mills ($\sqrt{\mathcal{N}} = 8$ supergravity) numerator
(a)–(d)	s^2
(e)–(g)	$(s(-\tau_{35} + \tau_{45} + t) - t(\tau_{25} + \tau_{45}) + u(\tau_{25} + \tau_{35}) - s^2)/3$
(h)	$(s(2\tau_{15} - \tau_{16} + 2\tau_{26} - \tau_{27} + 2\tau_{35} + \tau_{36} + \tau_{37} - u) + t(\tau_{16} + \tau_{26} - \tau_{37} + 2\tau_{36} - 2\tau_{15} - 2\tau_{27} - 2\tau_{35} - 3\tau_{17}) + s^2)/3$
(i)	$(s(-\tau_{25} - \tau_{26} - \tau_{35} + \tau_{36} + \tau_{45} + 2t) + t(\tau_{26} + \tau_{35} + 2\tau_{36} + 2\tau_{45} + 3\tau_{46}) + u\tau_{25} + s^2)/3$
(j)–(l)	$s(t - u)/3$

$$\tau_{ij} = 2k_j \cdot l_j$$

Recent Developments: Soft Graviton Behavior

See Yu-tin Huang's talk

Subleading Soft Gravity

See Yu-tin Huang's talk

BMS symmetry of asymptotically flat space time: diffeomorphisms leaving null infinity (Scri) invariant.

Bondi, Burg, Metzner; Sachs

The Weinberg soft graviton theorem is Ward identity of BMS symmetry.

He, Lysov, Pitra and Strominger

Using extended BMS symmetry Strominger obtained subleading soft limit identity.

Barnich and Troessaert; Strominger; Adamo, Casali and Skinner

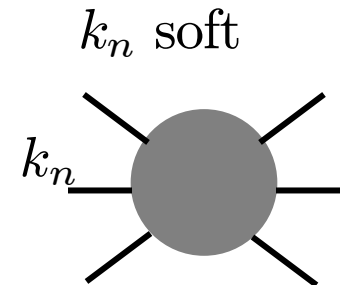
Verified at tree level by Cachazo and Strominger :

$$M_n^{\text{tree}} \rightarrow \left(S_n^{(0)} + S_n^{(1)} + S_n^{(2)} \right) M_{n-1}^{\text{tree}} + \mathcal{O}(k_n^2)$$

Weinberg term 

- Analogous to Low's 1958 soft-photon theorem.
- Gross and Jackiw (1968) discussed behavior through 2nd subleading order for a soft graviton coupled to scalars.
- Subleading corrections to soft gluons and gravitons had also been studied more recently.

Laenen, Magnea, Stavenga, White; White



Subleading Soft Graviton Behavior at Tree Level

Cachazo and Strominger

$$M_n^{\text{tree}} \rightarrow \left(S_n^{(0)} + S_n^{(1)} + S_n^{(2)} \right) M_{n-1}^{\text{tree}} + \mathcal{O}(k_n^2)$$

$$S_n^{(0)} = \sum_{i=1}^{n-1} \frac{\varepsilon_{\mu\nu} k_i^\mu k_i^\nu}{k_n \cdot k_i},$$

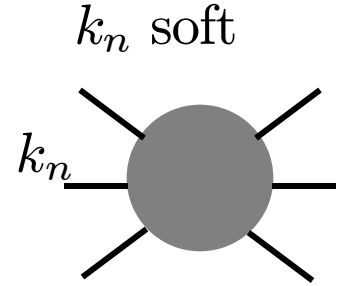
$$S_n^{(0)} = - \sum_{i=1}^{n-1} \frac{[n i] \langle x i \rangle \langle y i \rangle}{\langle n i \rangle \langle x n \rangle \langle y n \rangle},$$

$$S_n^{(1)} = \sum_{i=1}^{n-1} \frac{\varepsilon_{\mu\nu} k_i^\mu k_{n\rho} J_i^{\rho\nu}}{k_n \cdot k_i},$$

$$S_n^{(1)} = -\frac{1}{2} \sum_{i=1}^{n-1} \frac{[n i]}{\langle n i \rangle} \left(\frac{\langle x i \rangle}{\langle x n \rangle} + \frac{\langle y i \rangle}{\langle y n \rangle} \right) \tilde{\lambda}_n^{\dot{\alpha}} \frac{\partial}{\partial \tilde{\lambda}_i^{\dot{\alpha}}},$$

$$S_n^{(2)} = \frac{1}{2} \sum_{i=1}^{n-1} \frac{\varepsilon_{\mu\nu} k_{n\rho} J_i^{\rho\mu} k_{n\sigma} J_i^{\sigma\nu}}{k_n \cdot k_i}$$

$$S_n^{(2)} = -\frac{1}{2} \sum_{i=1}^{n-1} \frac{[n i]}{\langle n i \rangle} \tilde{\lambda}_n^{\dot{\alpha}} \tilde{\lambda}_n^{\dot{\beta}} \frac{\partial^2}{\partial \tilde{\lambda}_i^{\dot{\alpha}} \partial \tilde{\lambda}_i^{\dot{\beta}}},$$



Holds in D -dimensions

Schwab and Volovich, Afkhami-Jeddi

Similar behavior for gluons:

Casali

$$A_n^{\text{tree}} \rightarrow \left(S_{n \text{ YM}}^{(0)} + S_{n \text{ YM}}^{(1)} \right) A_{n-1}^{\text{tree}} + \mathcal{O}(k_n)$$

$$S_{n \text{ YM}}^{(0)} = \frac{\langle (n-1) 1 \rangle}{\langle (n-1) n \rangle \langle n 1 \rangle},$$

$$S_{n \text{ YM}}^{(1)} = \frac{1}{\langle (n-1) n \rangle} \tilde{\lambda}_n^{\dot{\alpha}} \frac{\partial}{\partial \tilde{\lambda}_{n-1}^{\dot{\alpha}}} - \frac{1}{\langle 1 n \rangle} \tilde{\lambda}_n^{\dot{\alpha}} \frac{\partial}{\partial \tilde{\lambda}_1^{\dot{\alpha}}}$$

Loop-Level Subleading Soft Behavior

Is subleading soft behavior renormalized in gravity? **Yes!**

However, corrections in gravity are fairly tame:

- No corrections beyond tree for leading behavior. ZB, Dixon, Perelstein, Rozowsky
- No corrections beyond one loop for 1st subleading behavior.
- No corrections beyond two loops for 2nd subleading behavior.

ZB, Davies, Nohle

Two types of corrections:

- Ones connected to IR singularities.
- Ones connected to nontrivial complex factorization channels where residues that are not simple products of amplitudes.

ZB, Davies, Nohle; He, Huang, Wen

Loop-Level Subleading Soft Behavior

Loop corrections linked to IR singularities:

ZB, Davies, Nohle

$$M_n^{1\text{-loop}} \rightarrow \left(S_n^{(0)} + S_n^{(1)} + S_n^{(2)} \right) M_{n-1}^{1\text{-loop}} + \left(S_n^{(0)1\text{-loop}} + S_n^{(1)1\text{-loop}} + S_n^{(2)1\text{-loop}} \right) M_{n-1}^{\text{tree}}$$

Straightforward to work out because the IR divergences are all known.

Weinberg; Naculich and Schnitzer; Akhoury, Saotome and Sterman,

$$M_n^{1\text{-loop}} \Big|_{\text{div.}} = \frac{\sigma_n}{\epsilon} M_n^{\text{tree}} \quad \sigma_n = -c_\Gamma \sum_{i=1}^{n-1} \sum_{j=i+1}^n s_{ij} \log\left(\frac{\mu^2}{-s_{ij}}\right)$$

IR discontinuities cause loop corrections

$$\sigma'_n = -c_\Gamma \sum_{i=1}^{n-1} s_{in} \log\left(\frac{\mu^2}{-s_{in}}\right)$$

$$S_n^{(0)1\text{-loop}} \Big|_{\text{div.}} = 0,$$

$$S_n^{(1)1\text{-loop}} \Big|_{\text{div.}} = \frac{1}{\epsilon} \left[\sigma'_n S_n^{(0)} - \left(S_n^{(1)} \sigma_{n-1} \right) \right],$$

$$S_n^{(2)1\text{-loop}} \Big|_{\text{div.}} = \frac{1}{\epsilon} \left[\sigma'_n S_n^{(1)} - \left(S_n^{(2)} \sigma_{n-1} \right) + \sum_{i=1}^{n-1} \frac{[n i]}{\langle n i \rangle} \left(\tilde{\lambda}_n^{\dot{\alpha}} \frac{\partial \sigma_{n-1}}{\partial \tilde{\lambda}_i^{\dot{\alpha}}} \right) \tilde{\lambda}_n^{\dot{\beta}} \frac{\partial}{\partial \tilde{\lambda}_i^{\dot{\beta}}} \right]$$

Good choice of momentum conservation prescription greatly simplifies expressions.

There are also IR finite pieces associated with these.

with P. Di Vecchia

Loop Corrections from Nontrivial Loop Factorizations

See Huang's talk

ZB, Davies and Nohle; He, Huang and Wen

At loop level in general there are factorization channels with no simple factorization properties also causing corrections. These come from careful implementation of dim. reg. $\epsilon/\epsilon \quad D = 4 - 2\epsilon$

Most easily seen in IR finite amplitudes of nonsusy Yang-Mills:

single pole double pole “unreal poles”

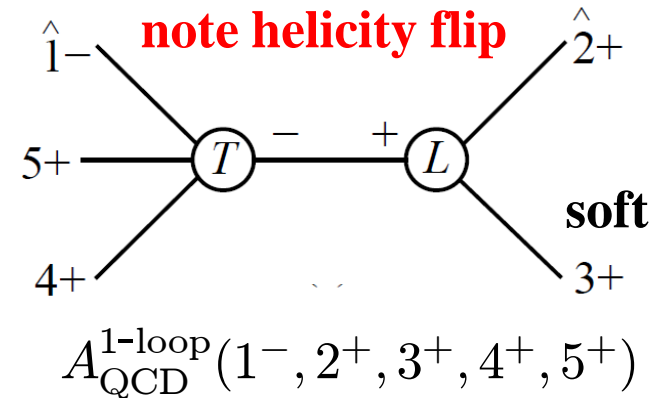
$$\frac{1}{\langle 23 \rangle}$$

standard factorization

$$\frac{[23]}{\langle 23 \rangle^2}$$

nonstandard factorization

$$\frac{[23]}{\langle 23 \rangle}$$



ZB, Dixon, Kosower; Vaman and Yao; Dunbar, Eittle and Perkins

Unreal poles contaminate the subleading soft behavior, causing nontrivial loop corrections in non-susy theories.

Gravity is similar except this happens at 2nd subleading order.

In nonsusy gravity extended BMS symmetry has anomalous 2nd subleading soft behavior

Renormalization of Subleading Soft Behavior

Very recently, Cachazo and Yuan noted that loop corrections can be avoided by reversing the standard order of limits in dimensional regularization.

$$D = 4 - 2\epsilon$$

Recall 1-loop QCD correction to leading soft function:

$$S_{n \text{ YM}}^{(0)1\text{-loop}} = -S_{n \text{ YM}}^{(0)} \frac{C_\Gamma}{\epsilon^2} \left(\frac{-\mu^2 S_{(n-1)1}}{S_{(n-1)n} S_{n1}} \right)^\epsilon \frac{\pi\epsilon}{\sin(\pi\epsilon)}$$

$$s_{ab} = (k_a + k_b)^2 \quad s_{n1} \rightarrow 0 \text{ for } k_n \rightarrow 0$$

**If $\epsilon < 0$ held fixed
1-loop YM soft
correction vanishes
for $k_n \rightarrow 0$**

However, serious problems:

- Soft physics comes out wrong.
- In QCD it ruins leading IR cancellations between real emission and virtual contributions.
- Keeping ϵ finite is in general impractical.

Keeping standard formulation of soft limits is greatly preferred.

Sample Application: UV properties of Gravity

UV Finiteness of $N = 8$ Supergravity?

If $N = 8$ supergravity is perturbatively finite it would imply a new symmetry or non-trivial dynamical mechanism. No known symmetry can render a $D = 4$ gravity theory finite.

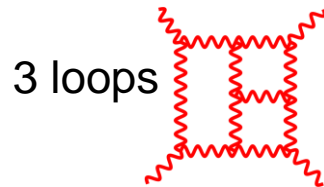
The discovery of such a mechanism would have a fundamental impact on our understanding of gravity.

Of course, perturbative finiteness is not the only issue for consistent gravity: Nonperturbative completions? High-energy behavior of theory? Realistic models?

Consensus opinion for the late 1970's and early 1980's:
All supergravity theories would diverge by three loops and therefore are not viable as fundamental theories.

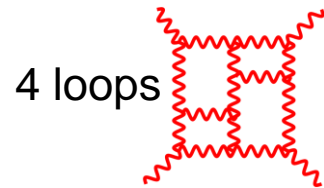
Feynman Diagrams for Gravity

SUPPOSE WE WANT TO CHECK IF CONSENSUS OPINION IS TRUE USING FEYNMAN DIAGRAMS.

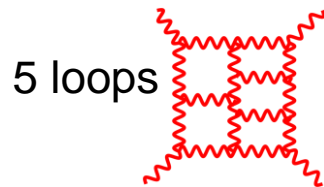


$\sim 10^{20}$
TERMS

No surprise it has never
been calculated via
Feynman diagrams.



$\sim 10^{26}$
TERMS



$\sim 10^{31}$
TERMS

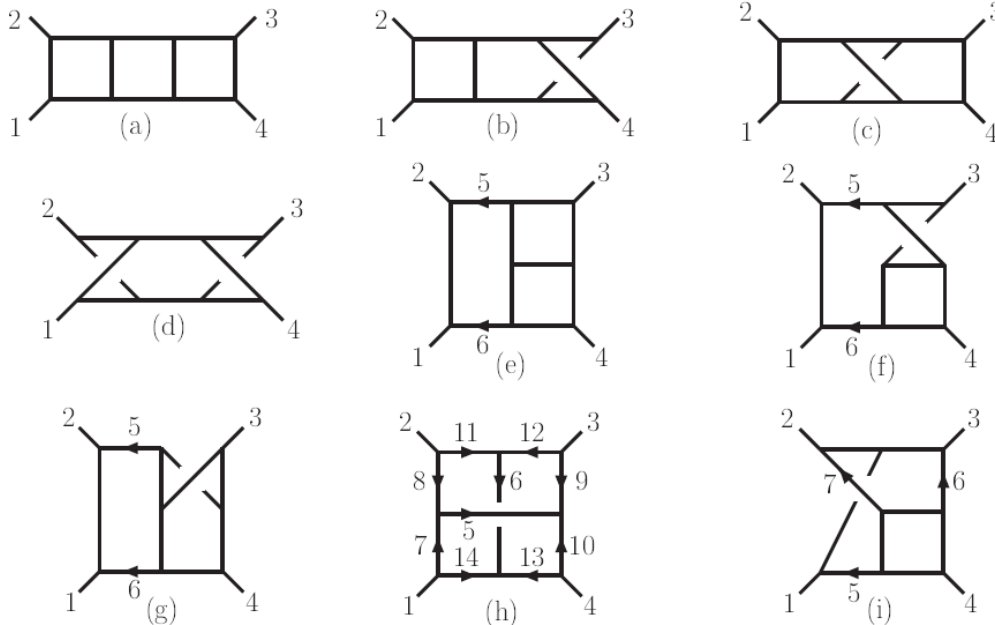
More terms than
atoms in your brain!

- Calculations to settle this seemed utterly hopeless!
- Seemed destined for dustbin of undecidable questions.

Complete Three-Loop Result

Analysis of unitarity cuts shows highly nontrivial all-loop cancellations. ZB, Dixon and Roiban (2006); ZB, Carrasco, Forde, Ita, Johansson (2007)

To test completeness of cancellations, we decided to directly calculate potential three-loop divergence.



ZB, Carrasco, Dixon, Johansson, Kosower, Roiban (2007)

Three loops is not only ultraviolet finite it is “superfinite”—finite for $D < 6$.

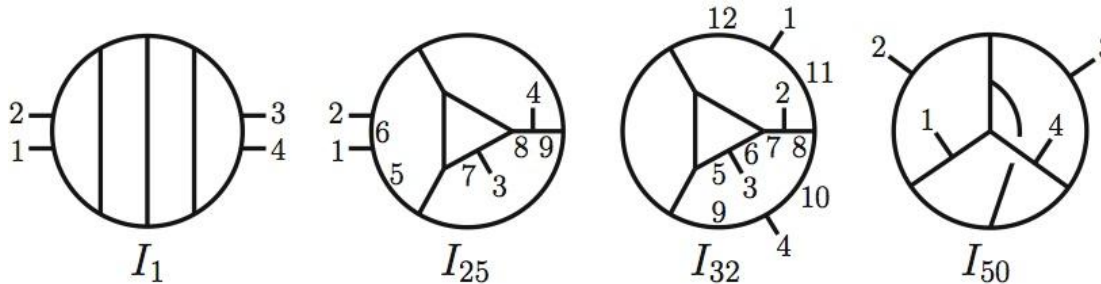
It is very finite!

Obtained via on-shell unitarity method.

Four-Loop $N = 8$ Supergravity Amplitude Construction

ZB, Carrasco, Dixon, Johansson, Roiban (2009)

Get 85 distinct diagrams or integrals.



$$M_4^{4\text{-loop}} = \left(\frac{\kappa}{2}\right)^{10} stu M_4^{\text{tree}} \sum_{S_4} \sum_{i=1}^{50} c_i I_i$$

← **Integral**
← **symmetry factor**
← **leg perms**

UV finite for $D < 11/2$
It's very finite!

Duality between color and kinematic discovered by doing this calculation.

Current Status of $N = 8$ Divergences

Consensus in $N = 8$ supergravity is that trouble starts at 5 loops and by 7 loops we have valid UV counterterm in $D = 4$ under all known symmetries (suggesting divergences).

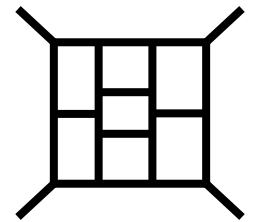
Bossard, Howe, Stelle; Elvang, Freedman, Kiermaier; Green, Russo, Vanhove ; Green and Bjornsson ; Bossard , Hillmann and Nicolai; Ramond and Kallosh; Broedel and Dixon; Elvang and Kiermaier; Beisert, Elvang, Freedman, Kiermaier, Morales, Stieberger

For $N = 8$ sugra in $D = 4$:

- All counterterms ruled out until 7 loops.
- But $D^8 R^4$ available at 7 loops (1/8 BPS) under all known symmetries. No known nonrenormalization theorem.

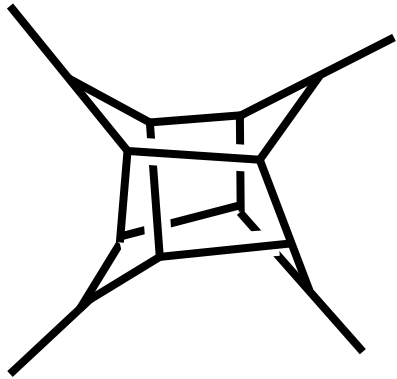
Bossard, Howe, Stelle and Vanhove

Based on this a reasonable person would conclude that $N = 8$ supergravity almost certainly diverges at 7 loops in $D = 4$.



$N = 8$ Supra 5 Loop Calculation

ZB, Carrasco, Johansson, Roiban



~500 such diagrams with ~1000s terms each

Being reasonable and being right are not the same.

Place your bets:

- At 5 loops in $D = 24/5$ does $N = 8$ supergravity diverge?
- At 7 loops in $D = 4$ does $N = 8$ supergravity diverge?

$D^8 R^4$ counterterms



Kelly Stelle:
English wine
“It will diverge”

5 loops



Zvi Bern:
California wine
“It won't diverge”

Recent Progress using Half Maximal Supergravity

While 5 loops $N = 8$ has proven to be formidable even with powerful methods, we have made significant recent progress by reducing the susy.

Examples of Magical Cancellations?

There are a number of curious connections between the divergences of $N = 4$ sYM and $N = 8$ supergravity through 4 loops in their critical dimensions (which I won't discuss here).

ZB, Carrasco, Dixon, Johansson, Roiban

Do we have any examples where a divergence vanishes but the standard symmetries suggest valid counterterms?

Yes!

Two examples in half-maximal supergravity :

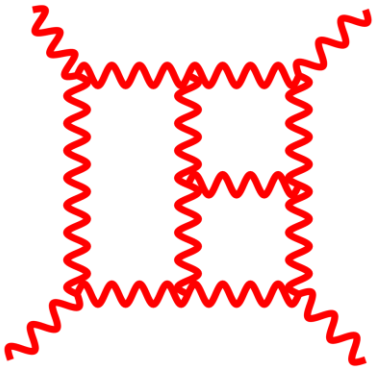
- $D = 5$ at 2 loops.
- $D = 4$ at 3 loops.

$N = 4$ supergravity of Cremmer, Ferrara and Scherk.

$N = 4$ Supergravity in $D = 4$

- $N = 4$ sugra at 3 loops ideal $D = 4$ test case to study.
- Representation where duality between color and kinematics manifest for $N = 4$ sYM 3-loop 4-pt amplitude.

ZB, Carrasco, Johansson (2010)



Consensus had it that a valid R^4 counterterm exists for this theory in $D = 4$.

Bossard, Howe, Stelle; Bossard, Howe, Stelle, Vanhove

$N = 8$ Supergravity \rightarrow $N = 4$ Supergravity

$D^8 R^4, L = 7 \rightarrow R^4, L = 3$

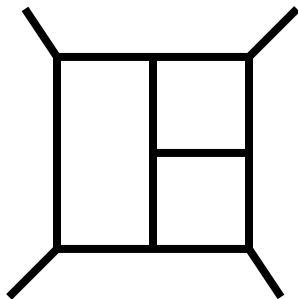
$D^{10} R^4, L = 8 \rightarrow D^2 R^4, L = 4$

Three-Loop Construction

ZB, Davies, Dennen, Huang

$N = 4$ sugra : $(N = 4$ sYM) \times $(N = 0$ YM)

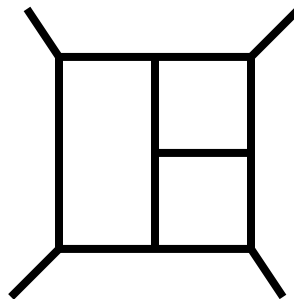
$N = 4$ sYM



$$\sim l \cdot k s^2 t A_4^{\text{tree}}$$

**BCJ
representation**

pure YM



$$\sim (\varepsilon_i \cdot l)^4 l^4$$

**Feynman
representation**

$$c_i \rightarrow n_i$$

$N = 4$ sugra linear
divergent

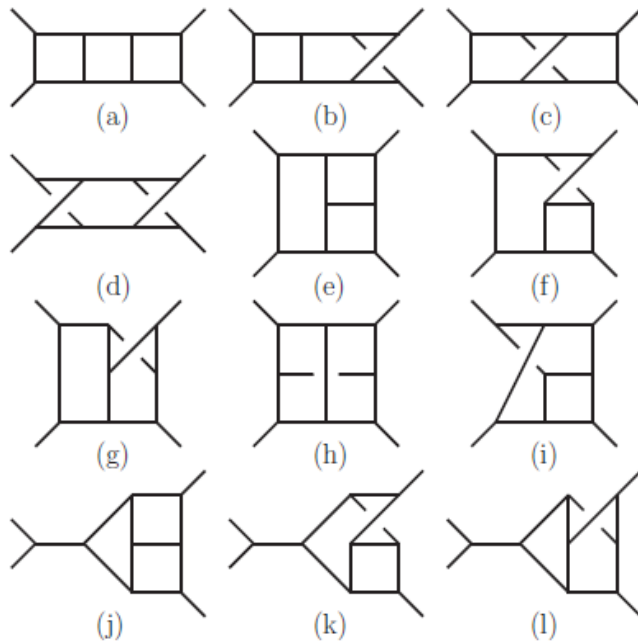
$$\int (d^D l)^3 \frac{k^7 l^9}{l^{20}}$$

simple to see
finite for $N=5,6,8$
sugra

- No covariant formalism exists that can make this manifestly UV finite.
- UV divergences are obtained by series expanding small external momentum (or large loop momentum).
- Introduce mass regulator for IR divergences.
- In general, subdivergences must be subtracted.

The $N = 4$ Supergravity UV Cancellation

ZB, Davies, Dennen, Huang



Graph	$(\text{divergence})/(\langle 12 \rangle^2 [34]^2 st A^{\text{tree}} (\frac{\kappa}{2})^8)$
(a)-(d)	0
(e)	$\frac{263}{768} \frac{1}{\epsilon^3} + \frac{205}{27648} \frac{1}{\epsilon^2} + \left(-\frac{5551}{768} \zeta_3 + \frac{326317}{110592} \right) \frac{1}{\epsilon}$
(f)	$-\frac{175}{2304} \frac{1}{\epsilon^3} - \frac{1}{4} \frac{1}{\epsilon^2} + \left(\frac{593}{288} \zeta_3 - \frac{217571}{165888} \right) \frac{1}{\epsilon}$
(g)	$-\frac{11}{36} \frac{1}{\epsilon^3} + \frac{2057}{6912} \frac{1}{\epsilon^2} + \left(\frac{10769}{2304} \zeta_3 - \frac{226201}{165888} \right) \frac{1}{\epsilon}$
(h)	$-\frac{3}{32} \frac{1}{\epsilon^3} - \frac{41}{1536} \frac{1}{\epsilon^2} + \left(\frac{3227}{2304} \zeta_3 - \frac{3329}{18432} \right) \frac{1}{\epsilon}$
(i)	$\frac{17}{128} \frac{1}{\epsilon^3} - \frac{29}{1024} \frac{1}{\epsilon^2} + \left(-\frac{2087}{2304} \zeta_3 - \frac{10495}{110592} \right) \frac{1}{\epsilon}$
(j)	$-\frac{15}{32} \frac{1}{\epsilon^3} + \frac{9}{64} \frac{1}{\epsilon^2} + \left(\frac{101}{12} \zeta_3 - \frac{3227}{1152} \right) \frac{1}{\epsilon}$
(k)	$\frac{5}{64} \frac{1}{\epsilon^3} + \frac{89}{1152} \frac{1}{\epsilon^2} + \left(-\frac{377}{144} \zeta_3 + \frac{287}{432} \right) \frac{1}{\epsilon}$
(l)	$\frac{25}{64} \frac{1}{\epsilon^3} - \frac{251}{1152} \frac{1}{\epsilon^2} + \left(-\frac{835}{144} \zeta_3 + \frac{7385}{3456} \right) \frac{1}{\epsilon}$

All three-loop divergences cancel completely!

All subdivergences cancel amongst themselves with uniform mass regulator.

Too bad bet wasn't on $N = 4$ supergravity!

Tourkine and Vanhove have understood this result by extrapolating from two-loop heterotic string amplitudes.

Explanations?

Prediction of superspace: If you add $N = 4$ vector multiplets, amplitude should develop no new 2, 3 loop divergences.

Bossard, Howe and Stelle (2013)

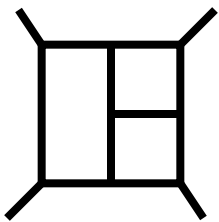
Note that $N = 4$ supergravity with matter already diverges at one loop.

Fischler (1979)

Prediction motivated us to check cases with vector multiplets.

ZB, Davies, Dennen (2013)

Four vector multiplet amplitude diverges at 2, 3 loops!



$$n_V = D_s - 4$$

$$\mathcal{M}^{(3)}(1_H, 2_H, 3_H, 4_H)|_{D=4 \text{ div.}} = 0, \quad \leftarrow \text{external graviton multiplets}$$

$$\mathcal{M}^{(3)}(1_H, 2_H, 3_V, 4_V)|_{D=4 \text{ div.}} = 0,$$

$$\mathcal{M}^{(3)}(1_V, 2_V, 3_V, 4_V)|_{D=4 \text{ div.}} = -\frac{1}{(4\pi)^6} \left(\frac{\kappa}{2}\right)^8 (s^2 + t^2 + u^2) st A_{Q=16}^{(0)}$$

UV divergence

matter multiplet

Similar story in $D = 5$

$$\times \frac{(D_s - 2)^2}{4} \left(\frac{D_s - 2}{2\epsilon^3} - \frac{1}{\epsilon^2} + \frac{1}{\epsilon} \right)$$

Adding vector multiplets causes new divergences both at 2, 3 loops.

Conclusion: currently, no viable complete standard symmetry understanding of 3-loop finiteness in pure $N = 4$ supergravity.

Explanations?

Key Question:

Is there an ordinary symmetry explanation for this?
Or is something extraordinary happening?

Bossard, Howe and Stelle (2013) showed that 3-loop finiteness can be explained by ordinary superspace +duality symmetries, *assuming* a 16 supercharge off-shell superspace exists.

$$\int d^4x d^{16}\theta \frac{1}{\epsilon} \mathcal{L}$$

More θ s implies more derivatives in operators

If true, then there is a perfectly good “ordinary” explanation.

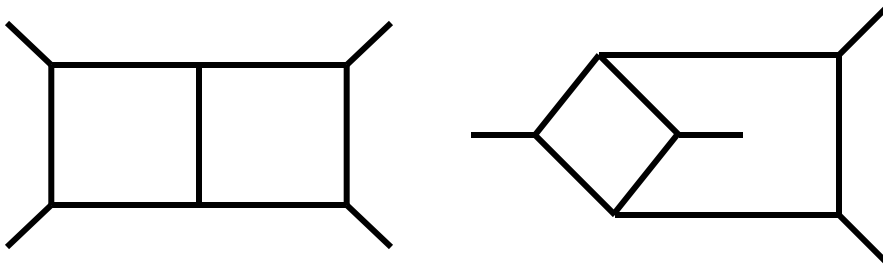
Does this superspace exist in $D = 5$ or $D = 4$?

Not easy to construct: A non-Lorentz covariant harmonic superspace .

What is the new “magic”?

To shed light on the source of magic we need a simpler example:

Half maximal supergravity at 2 loop in $D = 5$



**Susy + duality symmetry
does not appear to protect
against divergence in $D = 5$.**

Bossard, Howe and Stelle
ZB, Davies, Dennen

Story is similar to $D = 4, N = 4$ sugra at 3 loops except that it is much simpler to work with.

One-Loop Warmup in Half-Maximal Sugra

ZB, Boucher-Veronneau, Johansson
ZB, Davies, Dennen, Huang

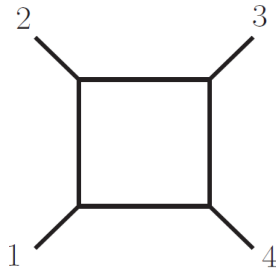
Generic color decomposition:

Dixon, Del Duca, Maltoni

$$\mathcal{A}_Q^{(1)} = ig^4 \left[c_{1234}^{(1)} A_Q^{(1)}(1, 2, 3, 4) + c_{1342}^{(1)} A_Q^{(1)}(1, 3, 4, 2) + c_{1423}^{(1)} A_Q^{(1)}(1, 4, 2, 3) \right]$$

$Q = \#$ supercharges

$Q = 0$ is pure non-susy YM



$c_{1234}^{(1)}$

is color factor of this box diagram

$$s = (k_1 + k_2)^2$$

$$t = (k_2 + k_3)^2$$

To get $Q + 16$ supercharge supergravity take 2nd copy $N = 4$ sYM

$N = 4$ sYM numerators very simple: independent of loop momentum

$$n_{1234} = n_{1342} = n_{1423} = st A_{Q=16}^{\text{tree}}(1, 2, 3, 4) \quad c_{1234}^{(1)} \rightarrow n_{1234}$$

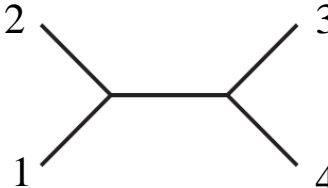
$$\mathcal{M}_{Q+16}^{(1)} = i \left(\frac{\kappa}{2} \right)^4 st A_{Q=16}^{\text{tree}}(1, 2, 3, 4) \left[A_Q^{(1)}(1, 2, 3, 4) + A_Q^{(1)}(1, 3, 4, 2) + A_Q^{(1)}(1, 4, 2, 3) \right]$$

One-loop divergences in pure YM

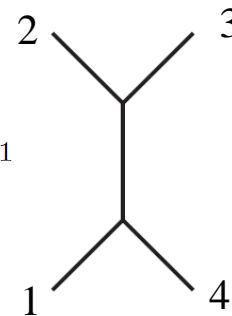
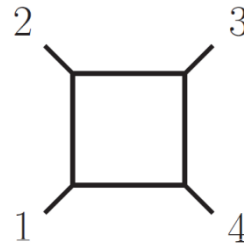
ZB, Davies, Dennen, Huang

Go to a basis of color factors

Three independent one-loop color tensors

$$b_1^{(0)} = \tilde{f}^{a_1 a_2 b} \tilde{f}^{b a_3 a_4}$$


$$b_2^{(0)} = \tilde{f}^{a_2 a_3 b} \tilde{f}^{b a_4 a_1}$$



$$b_1^{(1)} \equiv c_{1234}^{(1)} = \tilde{f}^{a_1 b_2 b_1} \tilde{f}^{a_2 b_3 b_2} \tilde{f}^{a_3 b_4 b_3} \tilde{f}^{a_4 b_1 b_4}$$

All other color factors expressible in terms of these three:

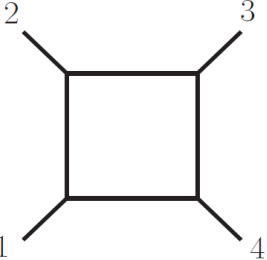
$$\mathcal{A}_Q^{(1)} = ig^4 \left[\overset{\text{one-loop color tensor}}{\curvearrowright} b_1^{(1)} \left(A_Q^{(1)}(1, 2, 3, 4) + A_Q^{(1)}(1, 3, 4, 2) + A_Q^{(1)}(1, 4, 2, 3) \right) \right. \\ \left. - \frac{1}{2} C_A b_1^{(0)} A_Q^{(1)}(1, 3, 4, 2) - \frac{1}{2} C_A b_2^{(0)} A_Q^{(1)}(1, 4, 2, 3) \right]$$

\curvearrowleft tree color tensor \curvearrowright

$C_A = 2 N_c$ for $SU(N_c)$

One-loop divergences in pure YM

In a basis of color factors:



↖ one-loop color tensor

$$\mathcal{A}_Q^{(1)} = ig^4 \left[b_1^{(1)} \left(A_Q^{(1)}(1, 2, 3, 4) + A_Q^{(1)}(1, 3, 4, 2) + A_Q^{(1)}(1, 4, 2, 3) \right) - \frac{1}{2} C_A b_1^{(0)} A_Q^{(1)}(1, 3, 4, 2) - \frac{1}{2} C_A b_2^{(0)} A_Q^{(1)}(1, 4, 2, 3) \right]$$

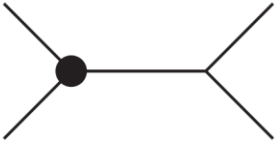
↖ tree color tensor

Q supercharges (mainly interested in $Q = 0$)

**$D = 4$: F^2 is only allowed counterterm by renormalizability
1-loop color tensor *not* allowed.**

$D = 6$: F^3 counterterm: 1-loop color tensor again *not* allowed.

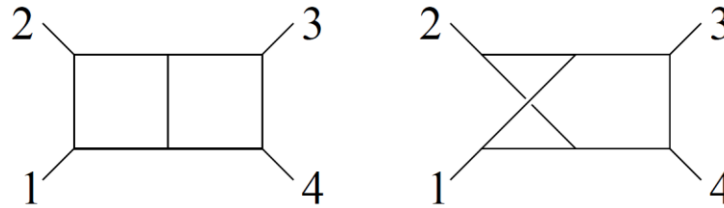
$$F^3 = f^{abc} F_\nu^{a\mu} F_\sigma^{b\nu} F_\mu^{c\sigma}$$



$$A_Q^{(1)}(1, 2, 3, 4) + A_Q^{(2)}(1, 3, 4, 2) + A_Q^{(1)}(1, 4, 2, 3) \Big|_{D=4,6 \text{ div.}} = 0$$

$$M_{Q+16}^{(1)}(1, 2, 3, 4) \Big|_{D=4,6 \text{ div.}} = 0$$

Two Loop Half Maximal Sugra in $D = 5$



ZB, Davies, Dennen, Huang

$$\mathcal{A}_Q^{(2)} = -g^6 \left[c_{1234}^P A_Q^P(1, 2, 3, 4) + c_{3421}^P A_Q^P(3, 4, 2, 1) \right. \\ \left. + c_{1234}^{NP} A_Q^{NP}(1, 2, 3, 4) + c_{3421}^{NP} A_Q^{NP}(3, 4, 2, 1) + \text{cyclic} \right]$$

$D = 5 F^3$ counterterm: 1,2-loop color tensors forbidden!



- 1) Go to color basis.
- 2) Demand no forbidden color tensors in pure YM divergence.
- 3) Replace color factors with kinematic numerators.

gravity $\mathcal{M}_{16+Q}^{(2)}(1, 2, 3, 4) \Big|_{D=5 \text{ div.}} = 0$

Half-maximal supergravity four-point divergence vanishes because forbidden color tensor cancels in pure YM theory.

Note: this cancellation is mysterious from standard symmetries.

Two Loop $D = 5$ UV Magic

ZB, Davies, Dennen, Huang

At least for 2 loops in $D = 5$ we have identified the source of unexpected UV cancellations in half-maximal supergravity:

It is the *same* magic found by 't Hooft and Veltman 40 years ago preventing forbidden divergences appearing in ordinary non-susy gauge theory!

Completely explains the $D = 5$ two-loop half-maximal sugra case, which still remains mysterious from standard supergravity symmetry viewpoint.

Half maximal supergravity at $L = 2, D = 5$ or $L = 3, D = 4$ are borderline cases.

We need to go beyond these and also develop a general understanding of the structure.

Four-loop $N = 4$ Supergravity Divergences

ZB, Davies, Dennen, Smirnov, Smirnov

Four loops done same way as three loops.

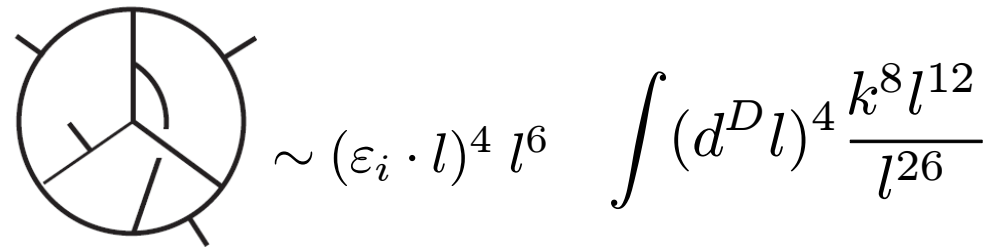
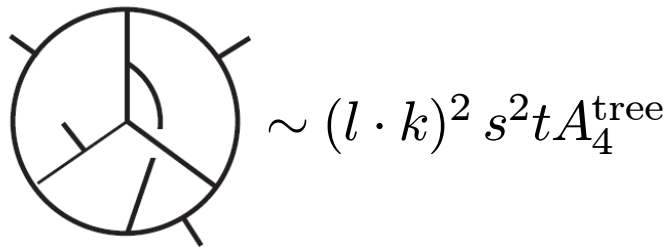
Similar to three loops except industrial level: C++ FIRE 4

$N = 4$ sugra: $(N = 4 \text{ sYM}) \times (N = 0 \text{ YM})$

$N = 4 \text{ sYM}$

pure YM

$N = 4$ sugra diagrams quadratically divergent



BCJ representation

Feynman representation

$D^2 R^4$ counterterm

82 nonvanishing diagram types using $N = 4$ sYM BCJ form.

Are there cancellations between the pieces?

The 4 loop Divergence of $N = 4$ Supergravity

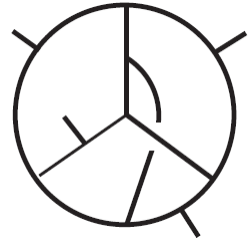
ZB, Davies, Dennen, Smirnov, Smirnov

Pure $N = 4$ supergravity is divergent at 4 loops with divergence

Result is
for Siegel
dimensional
reduction.

$$\mathcal{M}^{4\text{-loop}} \Big|_{\text{div.}} = \frac{1}{(4\pi)^8} \frac{1}{\epsilon} \left(\frac{\kappa}{2}\right)^{10} \frac{1}{144} (1 - 264\zeta_3) \mathcal{T}$$

dim. reg. UV pole



$$\mathcal{T} = st A_{\mathcal{N}=4}^{\text{tree}} (\mathcal{O}_1 - 28\mathcal{O}_2 - 6\mathcal{O}_3)$$

$$\mathcal{O}_1 = \sum_{S_4} (D_\alpha F_{1\mu\nu}) (D^\alpha F_2^{\mu\nu}) F_{3\rho\sigma} F_4^{\rho\sigma}$$

$$\mathcal{O}_2 = \sum_{S_4} (D_\alpha F_{1\mu\nu}) (D^\alpha F_2^{\nu\sigma}) F_{3\sigma\rho} F_4^{\rho\mu}$$

$$\mathcal{O}_3 = \sum_{S_4} (D_\alpha F_{1\mu\nu}) (D_\beta F_2^{\mu\nu}) F_{3\sigma}{}^\alpha F_4^{\sigma\beta}$$

$$s = (k_1 + k_2)^2$$

$$t = (k_2 + k_3)^2$$

$$F_j^{\mu\nu} \equiv i(k_j^\mu \varepsilon_j^\nu - k_j^\nu \varepsilon_j^\mu),$$

$$D^\alpha F_j^{\mu\nu} \equiv -k_j^\alpha (k_j^\mu \varepsilon_j^\nu - k_j^\nu \varepsilon_j^\mu)$$

Valid for all nonvanishing 4-point amplitudes of pure $N = 4$ sugra

Some Peculiar Properties



ZB, Davies, Dennen, Smirnov, Smirnov

Linear combinations to expose $D = 4$ helicity structure

Refers to helicities of pure YM component

$$\mathcal{O}^{--++} = \mathcal{O}_1 - 4\mathcal{O}_2$$

$$\mathcal{O}^{-+++} = \mathcal{O}_1 - 4\mathcal{O}_3$$

$$\mathcal{O}^{++++} = \mathcal{O}_2$$

$$\mathcal{O}^{--++} = 4s^2t \frac{\langle 12 \rangle^4}{\langle 12 \rangle \langle 23 \rangle \langle 34 \rangle \langle 41 \rangle},$$

$$\mathcal{O}^{-+++} = -12s^2t^2 \frac{[24]^2}{[12] \langle 23 \rangle \langle 34 \rangle [41]},$$

$$\mathcal{O}^{++++} = 3st(s+t) \frac{[12][34]}{\langle 12 \rangle \langle 34 \rangle},$$

The latter two configurations would vanish if a $U(1)$ symmetry were not anomalous.

See Carrasco, Kallosh, Tseytlin and Roiban

All three independent configurations have similar divergence.

Very peculiar because the nonanomalous sector should have a very different analytic structure. Not related by any supersymmetry Ward identities.

For anomalous sectors:

- $D = 4$ generalized cuts decomposing into tree amplitudes vanish.**
- Anomaly is ϵ/ϵ (UV divergence suppressed by ϵ).** $D = 4 - 2\epsilon$

Relation to $U(1)$ Anomaly



Anomalous sector feeds poor UV behavior into non-anomalous sector

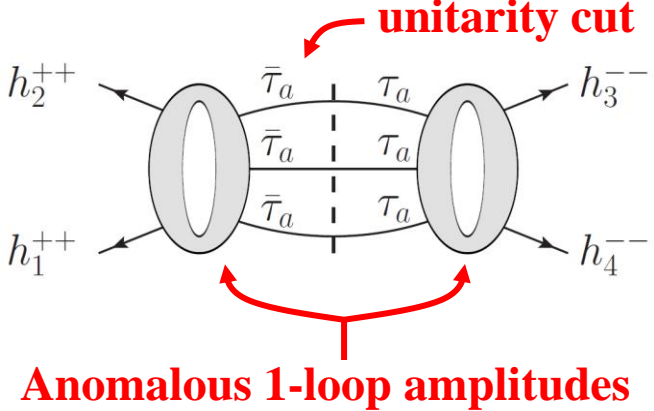


Figure from arXiv:1303.6219
Carrasco, Kallosh, Tseytlin and Roiban

- As pointed out by Carrasco, Kallosh, Roiban, Tseytlin the anomalous amplitudes are poorly behaved and contribute to a 4-loop UV divergence (unless somehow canceled as they are at 3 loops).
- Via the anomaly it is easy to understand why all three sectors can have similar divergence structure.
- The dependence of the divergence on vector multiplets matches anomaly.

anomaly has exactly this factor

$$\mathcal{M}_{n_V}^{4\text{-loop}} \Big|_{\text{div.}} = \frac{1}{(4\pi)^8} \left(\frac{\kappa}{2}\right)^{10} \frac{n_V + 2}{2304} \left[\frac{6(n_V + 2)n_V}{\epsilon^2} + \frac{(n_V + 2)(3n_V + 4) - 96(22 - n_V)\zeta_3}{\epsilon} \right] \mathcal{T}$$

n_V is number vector multiplets

Bottom line: The 4 loop divergence looks specific to $N = 4$ sugra and likely due to anomaly. Won't be present in $N \geq 5$ sugra.

$N = 5$ Supergravity

ZB, Davies, Dennen, Smirnov, Smirnov

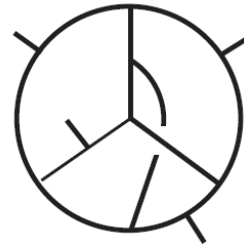
**An obvious test is to check $N = 5$ supergravity at four loops.
No anomalies and it is therefore expected to be finite.**

$N = 4$ sYM



**BCJ
representation**

$N = 1$ sYM



**Feynman
representation**

**Straightforward to obtain by adding a single adjoint fermion
to the pure Yang-Mills side of double copy.**

Almost finished, but didn't quite make it in time for conference.

Summary

- Gravity amplitudes are currently under intense study with many new representations and uncovered structures.
- Soft graviton behavior of amplitudes are corrected by loops.
- Surprisingly good UV behavior of supergravity uncovered. In $N = 4$ supergravity as yet no standard symmetry explanation for observed cancellations. Finiteness of $N \geq 5$ supergravity still an open problem.
- In the case of half maximal sugra at 2 loops in $D = 5$, where we could analyze in detail, “magical cancellations” follow from “standard magic” of Yang-Mills preventing forbidden divergences from appearing.

We can expect many more structures and surprises as we probe (super)gravity theories using modern tools and ideas.