arXiv:2108.03471v1 [nucl-th] 7 Aug 2021

Bayesian approach to long-range correlations and multiplicity fluctuations in nucleus-nucleus collisions

Kianusch Vahid Yousefnia,¹ Atharva Kotibhaskar,² Rajeev Bhalerao,³ and Jean-Yves Ollitrault¹

¹Institut de physique théorique, Université Paris Saclay, CNRS, CEA, F-91191 Gif-sur-Yvette, France

²Department of Physics, Fergusson College, F.C. Road, Pune 411004, India

³Department of Physics, Indian Institute of Science Education and Research (IISER), Homi Bhabha Road, Pune 411008, India

(Dated: August 10, 2021)

The number of particles detected in a nucleus-nucleus collision strongly depends on the impact parameter of the collision. Therefore, multiplicity fluctuations, as well as rapidity correlations of multiplicities, are dominated by impact parameter fluctuations. We present a method based on Bayesian inference which allows for a robust reconstruction of fluctuations and correlations at fixed impact parameter. We apply the method to ATLAS data on the distribution of charged multiplicity and transverse energy. We argue that multiplicity fluctuations are smaller at large rapidity than around central rapidity. We suggest simple, new analyses, in order to confirm this effect.

I. INTRODUCTION

The effective theory of strong interactions predicts that in ultrarelativistic nucleus-nucleus collisions, particle production occurs through the formation of color flux tubes parallel to the collision axis [1], which are analogous to strings [2]. Since these tubes extend over a wide range in rapidity, one expects that particle production at different rapidities is strongly correlated [3, 4]. The measurement of this rapidity-rapidity correlation [5] has remained elusive so far because it is hidden by the trivial correlation induced by the variation of impact parameter b within a given centrality class, since a more central collision produces more particles at all rapidities. Methods have been proposed to overcome this limitation [6], but in practice, experimental results are essentially limited to the forward-backward correlation [7-10], which has reduced sensitivity to impact parameter fluctuations.

We introduce a simple Bayesian method which allows us to reconstruct correlations at fixed impact parameter, and therefore eliminates the trivial correlation due to variations of b. More specifically, if (N_1, N_2, \dots, N_p) denote multiplicities, or transverse energies, in p rapidity windows, we show that their $p \times p$ covariance matrix can be reconstructed for b = 0 (central collisions), and a $(p-1) \times (p-1)$ projection of this matrix for b > 0. Note that Bayesian methods in the field of heavy-ion collisions usually involve sophisticated models and extensive calculations [12–15]. By contrast, the simple Bayesian approach implemented in this article does not involve any specific model of the collision dynamics. It could be implemented directly by experimental collaborations.

We apply the reconstruction to ATLAS data [11] on the distribution of the transverse energy E_T measured in a calorimeter located at forward rapidity, and of the charged multiplicity $N_{\rm ch}$ measured in the central rapidity region. The histogram of $(E_T, N_{\rm ch})$ for Pb+Pb collisions is represented in Fig. 1. The impact parameter of a Pb+Pb collision is not measured, and this histogram is the superposition of the contributions of all impact parameters. However, we are able to reconstruct accurately



FIG. 1. (Color online) Normalized histogram of the distribution of $(E_T, N_{\rm ch})$ measured by the ATLAS collaboration [11] in Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV. E_T denotes the transverse energy in the pseudorapidity window $3.2 < |\eta| < 4.9$, and $N_{\rm ch}$ the number of reconstructed tracks of charged particles in the pseudorapidity window $|\eta| < 2.5$. The black ellipse is one of the results given by our Bayesian reconstruction. It is the 99% confidence ellipse for collisions at zero impact parameter.

the location of central collisions on this diagram, which is represented in the figure as a black curve corresponding to the 99% confidence ellipse for collisions at b = 0. This is a specific example of what can be achieved through the Bayesian reconstruction.

The method is explained in Sec. II. In Sec. III, it is validated using a realistic model calculation in which the impact parameter is known, using the TRENTO 3D model of initial conditions [16]. In Sec. IV, we apply it to ATLAS data. In Sec V, we present our results for the fluctuations of E_T and $N_{\rm ch}$, and discuss what they tell us about the early collision dynamics. In Sec. VI, we suggest analyses which could be done easily with existing data, and which would shed new light on the rapidity dependence of fluctuations, and on long-range correlations.

II. METHOD

Detectors of heavy-ion experiments typically measure multiplicities of charged particles or energies deposited by these particles. We assume that there are p such observables in every event, which we denote by (N_1, N_2, \cdots, N_p) . Experimentally, one typically measures the probability distribution of (N_1, N_2, \cdots, N_p) in minimum-bias nucleus-nucleus collisions. Our goal is to reconstruct as much as we can of the probability distribution of (N_1, N_2, \cdots, N_p) at fixed impact parameter, without relying on any specific model of the collision dynamics. The single-variable case, p = 1, has been studied in [17]. The present work generalizes this earlier study to several variables. We illustrate this generalization on ATLAS data shown in Fig. 1. These data correspond to the case p = 2, where $N_1 = E_T$ and $N_2 = N_{ch}$. Below, we will use either the general notation (N_1, N_2) , or the specific notation (E_T, N_{ch}) , depending on the context.

The key assumption of our method is that for fixed impact parameter, fluctuations of N_i are Gaussian. This can be viewed as a consequence of the central limit theorem. In order for it to hold, we first need N_i to be large enough. In the case of the ATLAS data represented in Fig. 1, the multiplicity $N_{\rm ch}$ exceeds 3000 in central collisions, and an even larger number of particles contribute to the transverse energy E_T , so that this assumption is verified. Even though the particles seen in detectors are emitted in the last stages of the collision dynamics [18], there is a consensus that the origin of fluctuations lies in the early collision dynamics [19–24]. Now, due to the strong Lorentz contraction at ultrarelativistic energies, the processes through which particle production occurs at various points in the transverse plane are causally disconnected and hence independent from each other. Therefore, the multiplicity or transverse energy N_i in some detector can be seen as the sum of a large number of independent contributions, which is the condition under which the central limit theorem applies.

More specifically, we assume that the probability distribution of (N_1, N_2, \dots, N_p) at fixed impact parameter is a multivariate normal distribution.

$$P(N_1, ..., N_p|b) = \frac{\exp\left(-\frac{1}{2}(N_i - \bar{N}_i(b))\Sigma_{ij}^{-1}(b)(N_j - \bar{N}_j(b))\right)}{\sqrt{(2\pi)^p |\Sigma(b)|}}$$
(1)

where, in the exponential, we use the Einstein summation convention over the repeated indices i and j. In this equation, \bar{N}_i is the mean, or average, value of N_i , and Σ_{ij} is the symmetric covariance matrix:

$$\bar{N}_{i} = \langle N_{i} \rangle
\Sigma_{ij} = \langle (N_{i} - \bar{N}_{i})(N_{j} - \bar{N}_{j}) \rangle
= \langle N_{i}N_{j} \rangle - \bar{N}_{i}\bar{N}_{j},$$
(2)

where angular brackets denote an average over events

with the same impact parameter b. Σ^{-1} denotes the inverse matrix and $|\Sigma|$ the determinant. The validity of this Gaussian approximation will be discussed further in Sec. III.

The measured distribution is integrated over all values of impact parameter b. We carry out a simple change of variables and integrate instead over the cumulative probability distribution of b,

$$c_b \simeq \frac{\pi b^2}{\sigma_{\rm PbPb}},\tag{3}$$

where σ_{PbPb} is the inelastic nucleus-nucleus cross section [17].¹ c_b is referred to as the centrality fraction, or just centrality in the heavy-ion literature. The probability distribution of c_b is uniform in the interval [0, 1]. Therefore, the measured distribution is a simple integral over c_b :

$$P(N_1,\cdots,N_p) = \int_0^1 P(N_1,\cdots,N_p|c_b)dc_b.$$
(4)

We further assume that the parameters in Eq. (1), namely, the mean values \bar{N}_i and the elements of the covariance matrix Σ_{ij} , are smooth positive functions of c_b , which we choose to parametrize as the exponential of a polynomial [17]:

$$\bar{N}_i(c_b) = \bar{N}_i(0) \exp\left(-\sum_{n=1}^{n_{\max}} a_{i,n} c_b^n\right)$$
$$\Sigma_{ij}(c_b) = \Sigma_{ij}(0) \exp\left(-\sum_{m=1}^{m_{\max}} A_{i,j,m} c_b^m\right)$$
(5)

where $\bar{N}_i(0)$, $a_{i,n}$, $\Sigma_{ij}(0)$, $A_{i,j,m}$ are free parameters, and n_{\max} and m_{\max} are the degrees of the polynomials used to parametrize the mean and the covariance. The parameters are fitted in such a way that the distribution (4) matches data, and the degree of the polynomial is adjusted so as to obtain a satisfactory fit. We carry out a standard χ^2 fit, keeping all non-empty boxes in the histogram of $(E_T, N_{\rm ch})$.

As we shall see in Secs. III and IV, the simple procedure defined by Eqs. (1), (4) and (5) allows one to obtain a good fit to $P(N_1, \dots, N_p)$. The mean values $\overline{N}_i(c_b)$ returned by the fit closely match those obtained directly by fixing the impact parameter and averaging N_i over events, which implies that $\overline{N}_i(c_b)$ can be accurately reconstructed from data. By contrast, one cannot reconstruct the impact parameter dependence of the whole covariance matrix $\Sigma_{ij}(c_b)$, and discrepancies are expected between the best fit value, and the value calculated directly by fixing the impact parameter. We conclude this section by explaining which information can be reconstructed, and which cannot.

¹ Deviations from Eq. (3) only appear for very peripheral collisions which are excluded from our study.

First, consider the tip of the distribution, corresponding to the largest values of N_i . In this region, the integral in Eq. (4) is dominated by the contribution of central collisions, that is, $c_b = 0$. Therefore, both $\bar{N}_i(0)$ and $\Sigma_{ij}(0)$ can be reconstructed using the tip of the distribution. Looking at Fig. 1, one clearly sees that the ellipse, which is defined by $\bar{N}_i(0)$ and $\Sigma_{ij}(0)$, closely fits the tip of the distribution of (E_T, N_{ch}) .



FIG. 2. (Color online) Illustration of Eq. (7). The line is the parametric curve $(\bar{N}_1(c_b), \bar{N}_2(c_b))$, or ridge line (see text). For a given point (N_1, N_2) close to the ridge line, c_1 and c_2 defined by Eq. (7) are obtained by projections on the ridge line.

Consider now the bulk of the distribution, corresponding to less central collisions. The mean value $\bar{N}_i(c_b)$ can be reconstructed from data in the following way. Define $c(N_i)$ as the cumulative probability distribution of N_i :

$$c(N_i) \equiv \int_{N_i}^{\infty} P(N_i) dN_i.$$
(6)

This is the usual definition of centrality in experiment [25]. Except for very peripheral and very central collisions, $c(N_i)$ defined by Eq. (6) is almost equal to the inverse function of $\bar{N}_i(c_b)$ [17, 26]. Deviations are typically well below 1%. The functions $\bar{N}_i(c_b)$ define a ridge line which is displayed in Fig. 2.

By contrast, the centrality dependence of the covariance matrix cannot be fully reconstructed. This can be understood qualitatively as follows. A displacement parallel to the ridge line can be due either to a change in impact parameter, or to a fluctuation at fixed impact parameter. Therefore, only the projection of the covariance matrix orthogonal to the ridge line can be reconstructed.

We now explain this more quantitatively in the case p = 2. The only points where the probability is not negligible are those close to the ridge line (see Fig. 1). For a point (N_1, N_2) close to the ridge line, we denote by c_1 and c_2 the centralities defined by N_1 and N_2 (see Fig. 2):

$$N_1(c_1) = N_1$$

$$\bar{N}_2(c_2) = N_2.$$
 (7)

On the ridge line, $c_1 = c_2$. The only sizable contribution to the integral in Eq. (4) is from values of c_b close to c_1 and c_2 . We then replace N_i with $\bar{N}_i(c_i)$ in Eq. (1) and expand around c_b :

$$N_i - \bar{N}_i(c_b) = \bar{N}_i(c_i) - \bar{N}_i(c_b) = (c_i - c_b)\bar{N}'_i(c_b), \quad (8)$$

where \bar{N}'_i is the derivative of \bar{N}_i with respect to c_b . Since the integrand in Eq. (4) is sizable only in a narrow range, one can neglect the variation of $\bar{N}'_i(c_b)$ and of $\Sigma_{ij}^{-1}(c_b)$ in this range. Then, the integral over c_b is a simple Gaussian integral, which is worked out in Appendix A. In the case p = 2, one obtains:

$$P(N_1, N_2) \propto \exp\left(-\frac{\delta c^2}{2\sigma_{\perp}^2(c_b)}\right),$$
 (9)

where $\delta c \equiv c_2 - c_1$ (see Fig. 2), and

$$\sigma_{\perp}(c_b) \equiv \sqrt{\frac{\Sigma_{11}(c_b)}{\bar{N}_1'(c_b)^2} + \frac{\Sigma_{22}(c_b)}{\bar{N}_2'(c_b)^2} - 2\frac{\Sigma_{12}(c_b)}{\bar{N}_1'(c_b)\bar{N}_2'(c_b)}}.$$
 (10)

In writing Eq. (9), we have neglected the pre-exponential factor, which is independent of δc . c_b can denote any point between c_1 and c_2 , since this is a narrow range in which we neglect the variation of $\sigma_{\perp}(c_b)$. Up to a normalization constant, $\sigma_{\perp}^2(c_b)$ is the projection of the covariance matrix perpendicular to the ridge line.

In the case of the ATLAS analysis, δc represents the difference between the centralities determined according to E_T and $N_{\rm ch}$, and $\sigma_{\perp}(c_b)$ represents the width of the distribution of $(E_T, N_{\rm ch})$, in units of this centrality difference. As we shall see in Sec. III, the reconstruction of $\sigma_{\perp}(c_b)$ from the distribution of E_T and $N_{\rm ch}$ is robust for all c_b , while the elements of the covariance matrix $\Sigma_{ij}(c_b)$ are well reconstructed only for $c_b = 0$.

III. VALIDATION

In order to validate the reconstruction outlined in Sec. II. we test it using fake data generated using the $T_{R}ENTO 3D \mod [16]$. This is a state of the art Monte Carlo generator of the initial state of proton-nucleus and nucleus-nucleus collisions, which has been tuned to reproduce several rapidity-dependent observables. It returns, for each collision event, an entropy density profile $s(x, y, \eta)$ in three dimensions at an early time after the collision, where (x, y) are cartesian coordinates in the transverse plane and η is the space-time rapidity. We generate 10^7 minimum-bias Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV, and we convert the model predictions into values of E_T and $N_{\rm ch}$, that can be compared with the ATLAS data in Fig. 1 in the following way. We assume that $N_{\rm ch}$ is proportional to the initial entropy, obtained by integrating $s(x, y, \eta)$ over (x, y) and over the interval $|\eta| < 2.5$. We assume that E_T is proportional



FIG. 3. (Color online) Normalized histogram of $(E_T, N_{\rm ch})$ for 10⁷ minimum-bias Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV generated using the T_RENTo 3D model [16]. Note that the number of events is smaller by a factor ~ 16 than in the data shown in Fig. 1, and that we use a slightly finer binning in $(E_T, N_{\rm ch})$. The grey square at the bottom left represents the part of the histogram which is excluded from the fit, corresponding to the most peripheral collisions. The ellipses correspond to the 99% confidence ellipses for fixed values of b = 0, 2, 4, 6 fm, calculated directly (full lines, circles at the center) or from the Bayesian reconstruction (dashed lines, crosses at the center).

to the initial energy, obtained by integrating $s(x, y, \eta)^{4/3}$ over (x, y) and over the interval $3.2 < |\eta| < 4.9.^2$ Finally, we normalize E_T and $N_{\rm ch}$ in such a way that their mean values for central collisions coincide with those reconstructed from ATLAS data (Sec. IV). The resulting histogram of $(E_T, N_{\rm ch})$ is displayed in Fig. 3. It is roughly similar to the experimental distribution in Fig. 1.

In order to check the validity of the assumption of Gaussian fluctuations, Eq. (1), we also run the model for fixed impact parameter. We have generated 10^5 events for each value of b. The right panels of Fig. 4 display Gaussian fits to the left panels. The quality of the fits, measured by their χ^2 , decreases as the impact parameter increases. The Gaussian approximation captures the distribution except for the bottom panel, corresponding to b = 12 fm, where deviations are sizable. For this value of b, the Gaussian extends to negative values of E_T and $N_{\rm ch}$ which are unphysical.³ Therefore, the assumption of Gaussian fluctuations breaks down for peripheral collisions, which we exclude from the fit. Excluded events are



FIG. 4. (Color online) Left: Histogram of the distribution of $(E_T, N_{\rm ch})$ from the T_RENTo 3D model of initial conditions [16] (see Sec. III for details) in Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV for fixed values of the impact parameter, from top to bottom: b = 0, 4, 8, 12 fm. Right: two-dimensional Gaussian fits to these distributions. We indicate the χ^2 of each fit.

represented as a grey square at the bottom left of Fig. 3, corresponding to events with $E_T < E_{T,\min} = 0.289$ TeV and $N_{\rm ch} < N_{\rm ch,\min} = 194.39$.

We then fit the remaining distribution using the method outlined in Sec. II. We choose $n_{\text{max}} = 3$ and $m_{\text{max}} = 2$ in Eq. (5). Therefore, there are 4 parameters for $\bar{E}_T(c_b)$ and $\bar{N}_{ch}(c_b)$, and 3 parameters for each element of the covariance matrix, which gives a total of 17 fit parameters. Note that we need to evaluate the Pb+Pb cross section in order to convert the impact parameter into a centrality fraction using Eq. 3. This is done by computing the fraction f(b) of events whose impact parameter is below some value b, small enough that the probability of collision is very close to unity. The cross section is then evaluated as $\sigma_{\text{PbPb}} = \pi b^2/f(b)$ and we obtain $\sigma_{\text{PbPb}} = (800 \pm 2) \text{ fm}^2$.

In order to assess the quality of the reconstruction, we first evaluate the mean values of E_T and $N_{\rm ch}$ and their covariance matrix using Eq. (2), where one fixes the impact parameter before averaging over events. We compare these results with those obtained by fitting the dis-

 $^{^2}$ We thereby assume that the equation of state is conformal, which is approximately true at the high temperatures achieved in the early stages of the collision.

³ In the single-variable case this could in principle be fixed by replacing the Gaussian with a gamma distribution [27], but we have not found a simple way of generalizing this approach to the multi-variable case.



FIG. 5. (Color online) Left to right: $\bar{N}_{ch}(b)$, $\bar{E}_T(b)/\bar{N}_{ch}(b)$ and $\sigma_{\perp}(b)$ (defined by Eq. (10)) versus impact parameter. The vertical grey bands indicate the limits above which the reconstruction can no longer be trusted because peripheral events have been excluded from the fit (see text for details). The full lines correspond to the direct calculation, where one averages over events at fixed b. The derivatives $\bar{N}'_{ch}(c_b)$ and $\bar{E}'_T(c_b)$ in Eq. (10) are evaluated by fitting $\bar{N}_{ch}(c_b)$ and $\bar{E}_T(c_b)$ with a smooth curve and taking the derivative of the fit. The dashed lines in the left and middle panels correspond to the values reconstructed using the projected distributions $P(E_T)$ and $P(N_{ch})$, following the same method as in Ref. [17]. The dotted lines correspond to the values reconstructed using the distribution $P(E_T, N_{ch})$, which is done as explained in Sec. II. The bottom panels display the ratio between the reconstructed value and the direct calculation.

tribution as explained in Sec. II. Since we have excluded peripheral events, we do not expect the reconstruction to be valid if b is too large. However, the exclusion of peripheral events is done according to the values of E_T and $N_{\rm ch}$, not to the value of impact parameter. We expect that the reconstruction is robust for values of b such that at least 90% of events are included. The limiting value of b, above which this condition is no longer satisfied, is represented as light grey bands in Fig. 5. The dark grey band is defined by $\overline{E}_T(b) < E_{T,\min}$.

Figure 5 shows that the mean values $\bar{N}_{ch}(b)$ and $\bar{E}_T(b)$ are accurately reconstructed (we plot the ratio $\bar{E}_T(b)/\bar{N}_{ch}(b)$ rather than $\bar{E}_T(b)$ for reasons that will appear in Sec. IV), as well as the width $\sigma_{\perp}(b)$ defined by Eq. (10). We have argued on general grounds that one cannot expect to reconstruct the elements of the covariance matrix Σ_{ij} , except for central collisions. This is confirmed by the results displayed in Fig. 6. All three elements of the covariance matrix are reconstructed with an error smaller than 6% for b = 0. However, as b in-

creases, the reconstructed values quickly deviate from the direct calculation. This can also be seen by looking at the ellipses in Fig. 3. They are the 99% confidence ellipses (that is, ellipses that contain 99% of the events) for several values of b. They are defined by

$$(N_i - \bar{N}_i(b))\Sigma_{ij}^{-1}(b)(N_j - \bar{N}_j(b)) = -2\ln(1 - 0.99).$$
 (11)

For b = 0, the reconstructed ellipse almost coincides with the direct calculation, but the discrepancy quickly increases as b increases.

IV. APPLICATION TO ATLAS DATA

We finally apply the Bayesian reconstruction to the ATLAS data shown in Fig. 1. Some of the low multiplicity events correspond to photo-nuclear events which do not contribute to the inelastic cross section. Therefore, we evaluate the total number of inelastic collisions $N_{\rm eve}$ using the same centrality calibration as the ATLAS



FIG. 6. (Color online) Elements of the covariance matrix as a function of impact parameter in the T_RENTo 3D model. Top to bottom: Variance of $N_{\rm ch}$, variance of E_T , covariance of $N_{\rm ch}$ and E_T . The full lines correspond to the direct calculation, where one averages over events at fixed b. The dotted lines correspond to the reconstructed values.

collaboration, which gives $N_{\rm eve} = 158, 568, 641$. We exclude peripheral events using the same cuts as in the T_RENTo 3D calculation. Roughly 40% of the events are excluded. Figure 7 displays the ratio of the fit to data. The fit is excellent, except far out in the tail of the distribution. The deviation between fit and data is at the percent level between the two lines, which are drawn is such a way that they encompass 99% of the events.

The values of the fit parameters are listed in Table I. As explained in Sec. II, the output of the fit consists of the following information:

- Impact parameter dependence of the mean transverse energy \bar{E}_T and of the mean charged multiplicity \bar{N}_{ch} .
- Variances of E_T and N_{ch} , and covariance of E_T and N_{ch} , for central collisions at b = 0.
- Impact parameter dependence of the width σ_{\perp} defined by Eq. (10).



FIG. 7. (Color online) Ratio of the fit to ATLAS data in Fig. 1. As in Fig. 3, the grey square at the bottom left represents the part of the histogram which is excluded from the fit, corresponding to the most peripheral collisions. The lines correspond to contours with $\delta c = 2.6\sigma_{\perp}$ in Eq. (9), so that the region between the lines contains ~ 99% of the events. Specifically, the upper and lower lines are the parametric curves $(\bar{E}_T(c_b) + 2.6\sigma_{\perp}(c_b)\bar{E}'_T(c_b), \bar{N}_{ch}(c_b))$, and $(\bar{E}_T(c_b), \bar{N}_{ch}(c_b) + 2.6\sigma_{\perp}(c_b)\bar{N}'_{ch}(c_b))$.

We estimate the error on these quantities as follows. First, we check the robustness of the fit by varying the cutoffs in E_T and $N_{\rm ch}$, and we exclude a fraction of events varying from 35% to 45% of the events. The variation of the results is significantly smaller than the difference between the direct calculation and the reconstruction in the T_RENTo 3D calculation. Therefore, we estimate the relative error on our result as the maximum relative error between the direct calculation and the reconstructed value in the T_RENTo 3D calculation, shown in Figs. 5 and 6.

In the remainder of this section, we discuss our results for the mean values of $N_{\rm ch}$ and E_T . Note that these results could also have been obtained directly from the one-dimensional projections of the two dimensional histogram in Fig. 1, i.e., from the histograms of $N_{\rm ch}$ and E_T alone [17]. The most important results, which involve the covariance matrix of $N_{\rm ch}$ and E_T , will be discussed in Sec. V.

Figure 8 displays the mean charged multiplicity as a function of impact parameter. For the conversion between centrality fraction and impact parameter, we have used the value of the inelastic Pb+Pb cross section $\sigma_{\rm PbPb} = 767 \text{ fm}^2$ extracted from a Glauber calculation [28].⁴ Note that $N_{\rm ch}$ counts reconstructed tracks,

⁴ Note that this value of $\sigma_{\rm PbPb}$ differs from that returned by the T_RENTo 3D calculation. We compare model and experiment at the same impact parameter, not at the same centrality fraction.

(1	
	ATLAS	$T_{R}ENTO 3D$
$\bar{E}_T(0)$	4.424(33) TeV	$4.457(33) { m TeV}$
$\bar{N}_{ m ch}(0)$	3104(1)	3103(1)
$\Sigma_{E_T E_T}(0)$	$0.0224(11) \text{ TeV}^2$	$0.0663(34) \text{ TeV}^2$
$\Sigma_{N_{\rm ch}N_{\rm ch}}(0)$	19236(759)	19664(776)
$\Sigma_{E_T N_{\rm ch}}(0)$	$14.71(28) { m TeV}$	$34.62(70) { m TeV}$
$a_{E_T,1}$	4.00	4.43
$a_{N_{\rm ch},1}$	3.77	4.29
$a_{E_T,2}$	-1.49	-1.19
$a_{N_{\rm ch},2}$	-0.91	-1.78
$a_{E_{T},3}$	3.97	4.17
$a_{N_{\rm ch},3}$	3.86	4.15
$A_{E_T,E_T,1}$	4.08	17.33
$A_{N_{\rm ch},N_{\rm ch},1}$	7.18	11.71
$A_{E_T,N_{\rm ch},1}$	7.48	15.94
$A_{E_T,E_T,2}$	1.18	5.55
$A_{N_{\rm ch},N_{\rm ch},2}$	-1.89	-6.90
$A_{E_T,N_{\rm ch},2}$	1.32	-2.14

TABLE I. Values of fit parameters, as defined by Eq. (5), for ATLAS data shown in Fig. 1, and for the T_RENTo 3D calculation shown in Fig. 3. The uncertainties on the first five parameters are estimated using the relative difference between the direct calculation and its reconstruction in the T_RENTo 3D model. We omitted error bars on the polynomial coefficients in Eq. (5) because the individual uncertainties are not reflective of the uncertainty of the corresponding physical observable, e.g. \bar{E}_T .



FIG. 8. (Color online) Variation of the mean charged multiplicity as a function of impact parameter. The full line is the value reconstructed using ATLAS data. The shaded band is our estimate of the error band (see text). The dotted line is the result of the direct T_RENTO 3D calculation, shown as a full line in Fig. 5.

and that the efficiency of the reconstruction quickly decreases for $p_t < 0.8 \text{ GeV/c}$ [11]. Therefore, the value of $N_{\rm ch}$ seen by ATLAS is only ~ 29% of that seen by ALICE [29] in the corresponding pseudorapidity interval. For the sake of comparison, we also show the results from the T_RENTO 3D calculation (corresponding to the full lines in Fig. 5). The variation of $\bar{N}_{\rm ch}$ as a function of



FIG. 9. (Color online) Same as Fig. 8 for the ratio of the mean transverse energy to the mean charged multiplicity.

 \boldsymbol{b} is well reproduced by the model, although not perfectly.

Figure 9 displays the ratio of the mean transverse energy to the mean charged multiplicity, as a function of impact parameter. Let us first comment on the order of magnitude of the result for b = 0. The ratio $\bar{E}_T/\bar{N}_{\rm ch}$ can be decomposed as $(\bar{E}_T/\bar{N}_{\rm FCAL})(\bar{N}_{\rm FCAL}/\bar{N}_{\rm ch})$, where $N_{\rm FCAL}$ is the hadron multiplicity (neutral and charged) falling into the calorimeter acceptance. We evaluate $N_{\rm FCAL}$ by integrating the pseudorapidity spectra of charged particles [29] over the acceptance covered by the calorimeter, multiplying by a factor $\frac{3}{2}$ to take into account neutral particles. We then evaluate its average value at b = 0 by extrapolating linearly the values in the centrality intervals 0 - 5% and 5 - 10%, that is, $\bar{N}_{\text{FCAL}}(0) \simeq \frac{3}{2} N_{\text{FCAL}}(0 - 5\%) - \frac{1}{2} N_{\text{FCAL}}(5 - 10\%).$ We obtain $\bar{N}_{\text{FCAL}}(b = 0) \simeq 7930$. The contribution of a hadron to the transverse energy is roughly the transverse mass $m_t = \sqrt{p_t^2 + m^2}$. Using the value of $\bar{E}_T(0)$ in Table I, we obtain $\bar{E}_T(0)/\bar{N}_{\text{FCAL}}(0) = 0.558 \text{ GeV}$, which is the expected order of magnitude for the average transverse mass.

We next comment on the dependence of the ratio $\bar{E}_T/\bar{N}_{\rm ch}$ on impact parameter. ATLAS data show a non-monotonic behavior, which is not reproduced by the T_RENTo 3D model. This variation can be understood by decomposing again $E_T/N_{\rm ch}$ as $(E_T/N_{\rm FCAL})(N_{\rm FCAL}/N_{\rm ch})$. The ratio $E_T/N_{\rm FCAL}$ is the average transverse mass, which is determined by the mean transverse momentum $\langle p_t \rangle$, which itself decreases mildly as a function of impact parameter [30]. This effect is responsible for the decrease seen in the model, and in data at small b. The ratio $N_{\rm FCAL}/N_{\rm ch}$ is the ratio of the multiplicity in the forward rapidity region, covered by the calorimeter, and the central rapidity region, where $N_{\rm ch}$ is measured. Now, the rapidity distribution becomes slightly broader as the impact parameter increases [29]. The intuitive picture for this phenomenon is that the stopping between the nuclei is not as strong as in central collisions. This implies that the ratio $N_{\rm FCAL}/N_{\rm ch}$ increases as a function of impact parameter. This effect

is not reproduced by the T_RENTo 3D model. It overrides the decrease of $\langle p_t \rangle$ for large values of b, leading to the increase seen in ATLAS data.

V. FLUCTUATIONS AT LARGE AND CENTRAL RAPIDITIES

In this Section, we present results involving the covariance matrix of E_T and $N_{\rm ch}$. These are our most important results, as they shed light on the fluctuations in the central rapidity region probed by $N_{\rm ch}$, in the large rapidity region probed by E_T , and on their mutual correlations. We first isolate dynamical fluctuations of E_T and $N_{\rm ch}$ by subtracting out Poisson fluctuations in Sec. V A. We then present our results for central collisions in Sec. V B, and finally the results on the impact parameter dependence in Sec. V C.

A. Subtraction of Poisson fluctuations

We generally consider quantities of the form $N = \sum_{i=1}^{M} x_i$, where M is the particle multiplicity. Our goal is to identify non-trivial fluctuations and correlations of these quantities. The baseline is the case where the probabilities of finding a particle in different regions of phase space are independent variables [31]. In this case, x_i are independent variables and M follows a Poisson distribution. A simple calculation shows that the variance of N is the expectation value of $\sum_{i=1}^{M} x_i^2$. We refer to this contribution loosely as "the variance of Poisson fluctuations", and we subtract it from the observed variance in order to isolate the non-trivial part.

For the charged multiplicity, one simply counts particles, so that $x_i = 1$, and the subtraction is straightforward:

$$\Sigma_{N_{\rm ch}N_{\rm ch}}(b) \to \Sigma_{N_{\rm ch}N_{\rm ch}}(b) - \bar{N}_{\rm ch}(b).$$
 (12)

There is however an uncertainty on this subtraction due to hadronic decays. If a hadron decays into two charged particles, which both fall into the detector acceptance, then the contribution of that hadron is $x_i = 2$, not $x_i = 1$. The magnitude of this effect is modest, and cannot be evaluated accurately, because it depends on the hadronization mechanism. In the case of a fluiddynamical model, it depends on the freeze-out temperature [32]. We include this uncertainty in our error bar, by multiplying $\bar{N}_{ch}(b)$ in Eq. (12) by a coefficient which can vary between 1 and 1.2.

The transverse energy can be written as $E_T = \sum_i x_i$, where the sum has N_{FCAL} terms, corresponding to all hadrons, neutral and charged, falling in the calorimeter, and $x_i \simeq m_t$, where $m_t = \sqrt{p_t^2 + m^2}$ is the transverse mass. The average value of E_T at fixed b is $\bar{E}_T(b) =$ $\langle m_t \rangle \bar{N}_{\text{FCAL}}(b)$, and the variance of Poisson fluctuations is $\langle m_t^2 \rangle \bar{N}_{\text{FCAL}}(b)$, where angular brackets denote an average value over hadrons in the calorimeter. This quantity can be decomposed as

$$\langle m_t^2 \rangle \bar{N}_{\text{FCAL}}(b) = \frac{\langle m_t^2 \rangle}{\langle m_t \rangle^2} \frac{\bar{N}_{\text{ch}}(b)}{\bar{N}_{\text{FCAL}}(b)} \frac{\bar{E}_T(b)^2}{\bar{N}_{\text{ch}}(b)}.$$
 (13)

We evaluate the first factor on the right-hand side using ALICE data on identified particle spectra [33], which give $\langle m_t^2 \rangle / \langle m_t \rangle^2 \simeq 1.52$ for central collisions at $\sqrt{s_{\rm NN}} = 2.76$ TeV. We neglect the dependence of this ratio on impact parameter, rapidity and $\sqrt{s_{\rm NN}}$. The second factor can be estimated for central collisions using the value of $\bar{N}_{\rm ch}(0)$ in Table I, and the estimate $\bar{N}_{\rm FCAL}(0) \simeq 7930$ obtained in Sec. IV. We also neglect its dependence on rapidity and impact parameter. Putting these factors together, the subtraction of Poisson fluctuations is done according to the formula:

$$\Sigma_{E_T E_T}(b) \to \Sigma_{E_T E_T}(b) - \alpha \frac{E_T(b)^2}{\bar{N}_{\rm ch}(b)}, \qquad (14)$$

where $\alpha \simeq 0.59$. We assign an uncertainty of $\pm 20\%$ to this factor α .

Finally, since E_T and $N_{\rm ch}$ are measured in separate rapidity regions, no hadron contributes simultaneously to E_T and $N_{\rm ch}$, and Poisson fluctuations do not contribute to the covariance $\Sigma_{E_T N_{\rm ch}}(b)$.

Using the values in Table I, for central collisions, the subtracted quantities in Eqs. (12) and (14) are 16% and 17% of the total. This implies that at least 83% of the variances of E_T and $N_{\rm ch}$ can be attributed to dynamical fluctuations.

B. Fluctuations of N_{ch} and E_T in central collisions

	E_T	$N_{ m ch}$
	$1.14(6) \times 10^{-3}$	$1.07(2) \times 10^{-3}$
E_T	$0.95(7) imes 10^{-3}$	$1.07(2) imes 10^{-3}$
	(3.57×10^{-3})	(2.52×10^{-3})
	$1.07(2) \times 10^{-3}$	$2.00(8) \times 10^{-3}$
$N_{\rm ch}$	$1.07(2) imes 10^{-3}$	$1.68(9) imes 10^{-3}$
	(2.52×10^{-3})	(1.96×10^{-3})

TABLE II. Relative covariance matrix σ_{ij} of E_T and N_{ch} , defined by Eq. (15), for central Pb+Pb collisions (b = 0). For each element, the first line is the total covariance returned by the fit to data. The second line is the dynamical covariance obtained after subtracting Poisson fluctuations (see Sec. VA). The third line is the value obtained in the TRENTO 3D calculation by simulating events at b = 0, which differs by a few percent from the reconstructed value in Table I. The TRENTO 3D calculation returns a continuous density profile, and does not include hadronization. Therefore, the fluctuations calculated in this model are dynamical, and the subtraction of Poisson fluctuations explained in Sec. V A does not apply.

We now discuss the fluctuations of E_T and N_{ch} for central collisions, and their mutual correlation. We use as a measure of these fluctuations the relative covariance matrix [34], defined as

$$\sigma_{ij} \equiv \frac{\langle N_i N_j \rangle}{\langle N_i \rangle \langle N_j \rangle} - 1 = \frac{\Sigma_{ij}}{\bar{N}_i \bar{N}_j}.$$
 (15)

Its diagonal elements are the relative variances of N_i . Values extracted from ATLAS data (Table II) show that the relative fluctuations of E_T are smaller than those of $N_{\rm ch}$, both before and after isolating dynamical fluctuations. This is the reason why E_T is a better estimator of the centrality than $N_{\rm ch}$ [11, 35]. It is a non-trivial observation, which is not reproduced by the T_RENTO 3D calculation. The T_RENTO 3D calculation reproduces the variance of $N_{\rm ch}$ to a good approximation, which is not surprising since it was fitted to the distribution of the charged multiplicity measured near mid-rapidity. On the other hand, it overestimates the variance of E_T by at least a factor 3.

There is no trivial relation between the relative fluctuations of E_T and the relative fluctuations of the multiplicity in the same rapidity window. In a hydrodynamic picture of the collision [36], however, one expects them to be of the same order of magnitude, because the multiplicity is proportional to the entropy, and the energy and the entropy are related through the equation of state. In Sec. VI, we will suggest a specific analysis in order to check this assumption. If it is correct, our results imply that multiplicity fluctuations are smaller at large rapidity than around central rapidity.

The values of the relative variance of E_T in Table II can be compared with those previously extracted from ATLAS data at 2.76 TeV [17]. After subtracting Poisson fluctuations with help of multiplicity densities measured by ALICE [37], we obtain $\sigma_{11} = 0.76 \times 10^{-3}$. This is smaller than the value at 5.02 TeV in Table II, which implies that multiplicity fluctuations increase as a function of the rapidity gap between incoming nuclei and the detector, which increases by ~ 0.6 between 2.76 TeV and 5.02 TeV. It is tempting to postulate that multiplicity fluctuations depend on the rapidity difference with incoming nuclei, much as average multiplicities themselves [38, 39].

The second important observation is that E_T and $N_{\rm ch}$ are strongly correlated, even at fixed impact parameter. This is illustrated by the elongated covariance ellipse represented in Fig. 10. The correlation is significantly stronger after Poisson fluctuations have been subtracted out.

We now interpret these results within a simple model. We assume that a nucleus-nucleus collision at impact parameter b produces N_s sources, and that E_T and $N_{\rm ch}$ are obtained by summing the contributions of all sources in the event [35, 40, 41]. We assume that the sources



FIG. 10. (Color online) Full line: 99% confidence ellipse of zero impact parameter collisions, already shown in Fig. 1. The ellipse is defined by Eq. (11). The dashed line represents the ellipse, after subtraction of the contribution of Poisson fluctuations. The dotted line is the value from the T_RENTO 3D calculation.

are independent.⁵ We denote the contribution of a single source to $N_1 = E_T$ and $N_2 = N_{\rm ch}$ by n_1 and n_2 , respectively. Under these assumptions, one easily obtains:

$$\sigma_{ij} = \frac{\langle N_s^2 \rangle - \langle N_s \rangle^2}{\langle N_s \rangle^2} + \frac{1}{\langle N_s \rangle} \frac{\langle n_i n_j \rangle - \langle n_i \rangle \langle n_j \rangle}{\langle n_i \rangle \langle n_j \rangle} \quad (16)$$

The first term in the right-hand side is the relative variance of the number of sources, which gives the same contribution to all elements of the covariance matrix. It is a trivial correlation arising from a global change in the system, much as correlation stemming from the variation of impact parameter [42]. It does not shed light on rapidity correlations in the emission process. The last term in the right-hand side, which involves the covariance matrix of a single source, is physically more interesting.

It is not possible to uniquely determine the quantities in the right-hand side of Eq. (16), because there are more unknowns than equations. But one can obtain non-trivial information. For instance, we have noted that the relative variance of E_T is smaller than that of $N_{\rm ch}$. Since they get the same contribution from the variance of the number of sources, it implies that the difference between the relative variances of E_T and $N_{\rm ch}$ is larger at the level of a single source. The suppression of fluctuations at large rapidity is by no means a small effect. One can also obtain non-trivial information on the correlation between n_1 and n_2 , which represent the contributions of a single source to E_T and $N_{\rm ch}$. For in-

⁵ Note that a similar picture underlies our assumption of Gaussian fluctuations, as explained in Sec. II.

stance, one can readily exclude that they are uncorrelated, because this would imply $\sigma_{11} > \sigma_{12}$, at variance with the values in Table II, after Poisson fluctuations have been subtracted out. One can obtain a lower bound $r_{\rm min}$ on the Pearson correlation coefficient r between n_1 and n_2 , defined by $\langle n_1 n_2 \rangle - \langle n_1 \rangle \langle n_2 \rangle = r \sigma(n_1) \sigma(n_2)$, where $\sigma(n_i) \equiv \sqrt{\langle n_i^2 \rangle - \langle n_i \rangle^2}$ is the standard deviation of n_i . Elementary algebra shows that $r^2 - r_{\rm min}^2$ is a perfect square, where $r_{\rm min}$ is defined by:

$$r_{\min} \equiv 2 \frac{\sqrt{(\sigma_{12} - \sigma_{11})(\sigma_{22} - \sigma_{12})}}{\sigma_{22} - \sigma_{11}}$$
(17)

With the values in Table II, and taking into account the error bars (from the reconstruction and from the subtraction of Poisson fluctuations) we obtain $r_{\rm min} = 0.72\pm0.15$. Two effects may contribute to this strong correlation: Some sources may be stronger than others, for instance those situated in the center of the interaction region. Stronger sources yield larger values of both n_1 and n_2 , which induces a mutual correlation. The second effect is the dynamical effect that one would like to isolate, that particle production from a single source is strongly correlated across rapidities. More detailed modeling will be necessary to disentangle the relative contributions of these two effects.



FIG. 11. (Color online) Same as Fig. 8 for the width of the distribution, defined by Eq. (10). Full line: before subtracting Poisson fluctuations. Dashed line: after subtracting Poisson fluctuations according to Eqs. (14) and (12). The shaded area around the curve is our estimate of the error on the reconstruction. Dotted line: value calculated in the T_RENT0 3D model.

C. Impact parameter dependence of fluctuations

As explained in Sec. II, one cannot reconstruct the impact parameter dependence of the full covariance matrix. Only the specific linear combination σ_{\perp} , defined

by Eq. (10), can be reconstructed for all impact parameters. Its impact parameter dependence is represented in Fig. 11, before and after subtracting Poisson fluctuations (Sec. V A). The T_RENTO 3D model underpredicts σ_{\perp} for all values of impact parameter, which is the reason why the distribution in Fig. 3 is narrower than that in Fig. 1. This shows that the width of the distribution, as measured by σ_{\perp} , contains non-trivial dynamical information, which can be used to discriminate between models.

VI. CONCLUSIONS AND PERSPECTIVES

We have introduced a simple Bayesian method which allows for a robust reconstruction of multiplicity fluctuations and rapidity correlations in nucleus-nucleus collisions at fixed impact parameter b. Starting from the distribution $p(N_1, \dots, N_p)$ of multiplicities (or transverse energies, or number of hits) N_1, \dots, N_p measured in different parts of the detector, one can reconstruct the full covariance matrix of (N_1, \dots, N_p) at b = 0, as well as the impact parameter dependence of a $(p-1) \times (p-1)$ projection of the matrix. We have applied the method to ATLAS data on the joint distribution of transverse energy and charged multiplicity. We have shown that dynamical fluctuations are smaller at large rapidity than around central rapidity, and that particle production is strongly correlated across rapidity, even at fixed impact parameter. We have also shown that the width of the distribution and its impact parameter dependence are not reproduced by our model calculation, so that this observable can be used to rule out three-dimensional models of initial fluctuations.

Rapidity correlations have so far been studied using multiplicities of charged particle tracks [8, 9]. However, particle tracks are typically reconstructed only around central rapidity, so that the rapidity coverage of such studies is limited. We have shown that track reconstruction is not needed. The transverse energy in a calorimeter, the multiplicity of pixel clusters [43] or the number of hits in a scintillator [44] also give a quantitative information on fluctuations, provided that the detector is large enough that dynamical fluctuations dominate over Poisson fluctuations. A few tests should however be made in order to interpret correctly the information from calorimeters. Specifically, one should first figure out how energy fluctuations relate to multiplicity fluctuations in the same rapidity window. This can be done easily using one of the central detectors of ALICE, CMS or ATLAS. These detectors measure the multiplicity of charged particle tracks and their momenta. One can use the sum of transverse momenta of charged particles, $P_T \equiv \sum_{N_{\rm ch}} p_t$, as a proxy for their transverse energy. By comparing the tails of the distributions of $N_{\rm ch}$ and P_T , one can readily check how their relative variances compare for central collisions.

We conclude by listing analyses that would shed ad-

ditional light on long-range correlation. The transverse energy is usually defined as $E_T = E_F + E_B$, where E_F and E_B are transverse energies in forward and backward calorimeters [9, 43]. If measured, the distribution $p(E_F, E_B)$ would provide direct information on forwardbackward rapidity correlations. This analysis would be even simpler than that carried out in this paper because backward and forward calorimeters are symmetric around mid-rapidity, and this symmetry reduces the number of fit parameters in the Bayesian analysis. Similarly, the ALICE collaboration could measure the distribution $p(V_{0A}, V_{0C})$, where V_{0A} and V_{0C} are the multiplicities in backward and forward scintillators [25]. These analyses can be readily carried out by experimental collaborations using existing data, by following the exact same steps as our analysis of ATLAS public data.

More detailed information on the rapidity structure could be obtained by extending the analysis to three variables. For instance, in the case of the ATLAS data studied in this paper [11], one could split the calorimeter into its forward and backward components and measure $p(E_F, E_B, N_{ch})$. Similar analyses could be done with the CMS and ALICE detectors. They would yield detailed information on the long-range rapidity structure of correlations. Finally, the same method could be applied to heavy-ion collisions at lower energies [45–47].

ACKNOWLEDGEMENTS

We thank J. Jia for sending us the centrality calibration of data. AK thanks Bhavya Bhatt for technical assistance in the early stages of this work. JYO thanks F. Gelis and G. Giacalone for discussions.

Appendix A: Reconstruction of the covariance matrix in non-central collisions

In this Appendix, we specify which information about the covariance matrix $\Sigma_{ij}(b)$ can be reconstructed for b > 0. Inserting Eq. (8) into Eq. (1) and neglecting the pre-

- T. Lappi and L. McLerran, Nucl. Phys. A 772, 200-212 (2006) [arXiv:hep-ph/0602189 [hep-ph]].
- [2] B. Andersson, G. Gustafson, G. Ingelman and T. Sjostrand, Phys. Rept. 97, 31-145 (1983)
- [3] F. Gelis, T. Lappi and R. Venugopalan, Phys. Rev. D 79, 094017 (2009) [arXiv:0810.4829 [hep-ph]].
- [4] T. Lappi and A. Ramnath, Phys. Rev. D 100, no.5, 054003 (2019) [arXiv:1904.00782 [hep-ph]].
- [5] A. Bzdak and D. Teaney, Phys. Rev. C 87, no.2, 024906 (2013) [arXiv:1210.1965 [nucl-th]].
- [6] A. Olszewski and W. Broniowski, Phys. Rev. C 96, no.5, 054903 (2017) [arXiv:1706.02862 [nucl-th]].

exponential factor, one obtains:

$$P(N_1, ..., N_p | b) \propto \exp\left(-\frac{1}{2}\sum_{ij}(c_i - c_b)\bar{N}'_i \Sigma_{ij}^{-1}\bar{N}'_j(c_j - c_b)\right)$$
where we omit the dependence of \bar{N}'_i and Σ^{-1} on h for

where we omit the dependence of N'_i and Σ_{ij}^{-1} on b for simplicity. Then, the integral in Eq. (4) is a Gaussian integral over c_b . Evaluating this integral, and neglecting again the pre-exponential factor, one obtains

$$P(N_1, ..., N_p) \propto \exp\left(-\frac{1}{2}\sum_{ij} c_i \Pi_{ij} c_j\right), \qquad (A2)$$

where

$$\Pi_{ij} \equiv \bar{N}_i' \Sigma_{ij}^{-1} \bar{N}_j' - \frac{\left(\sum_{\alpha} \bar{N}_i' \Sigma_{i\alpha}^{-1} \bar{N}_{\alpha}'\right) \left(\sum_{\beta} \bar{N}_j' \Sigma_{j\beta}^{-1} \bar{N}_{\beta}'\right)}{\sum_{\alpha\beta} \bar{N}_{\alpha}' \Sigma_{\alpha\beta}^{-1} \bar{N}_{\beta}'}.$$
(A3)

 Π is the quantity which can be reconstructed for all impact parameters. It is a $(p \times p)$ symmetric matrix which verifies the property $\sum_{i} \prod_{ij} = 0$ for all j, that is, all lines and columns sum up to zero. Due to this property, one can rewrite Eq. (A2) as:

$$P(N_1, ..., N_p) \propto \exp\left(-\frac{1}{2}\sum_{i>1, j>1} (c_i - c_1)\Pi_{ij}(c_j - c_1)\right).$$
(A4)

This form shows that the distribution of $(N_1, ..., N_p)$ is a function of p-1 variables c_i-c_1 . Physically, Π represents the projection of the covariance matrix onto the (p-1)-dimensional subspace orthogonal to the ridge line. In the case p = 2, Eq. (A3) gives

$$\mathbf{\Pi} = \frac{1}{\sigma_{\perp}^2} \begin{pmatrix} 1 & -1 \\ -1 & 1 \end{pmatrix},\tag{A5}$$

where σ_{\perp} is defined by Eq. (10). Inserting Eq. (A5) into Eq. (A4), one recovers Eq. (9).

- [7] B. B. Back *et al.* [PHOBOS], Phys. Rev. C 74, 011901 (2006) [arXiv:nucl-ex/0603026 [nucl-ex]].
- [8] B. I. Abelev *et al.* [STAR], Phys. Rev. Lett. **103**, 172301 (2009) [arXiv:0905.0237 [nucl-ex]].
- [9] M. Aaboud *et al.* [ATLAS], Phys. Rev. C **95**, no.6, 064914 (2017) [arXiv:1606.08170 [hep-ex]].
- [10] S. Acharya *et al.* [ALICE], Phys. Lett. B **781**, 20-32 (2018) [arXiv:1710.07975 [nucl-ex]].
- [11] M. Aaboud *et al.* [ATLAS], JHEP **01**, 051 (2020) [arXiv:1904.04808 [nucl-ex]].
- [12] J. E. Bernhard, J. S. Moreland, S. A. Bass, J. Liu and U. Heinz, Phys. Rev. C 94, no.2, 024907 (2016) [arXiv:1605.03954 [nucl-th]].

12

- [13] G. Nijs, W. van der Schee, U. Gürsoy and R. Snellings, Phys. Rev. C 103, no.5, 054909 (2021) [arXiv:2010.15134 [nucl-th]].
- [14] D. Everett *et al.* [JETSCAPE], Phys. Rev. C 103, no.5, 054904 (2021) [arXiv:2011.01430 [hep-ph]].
- [15] J. E. Parkkila, A. Onnerstad and D. J. Kim, [arXiv:2106.05019 [hep-ph]].
- [16] W. Ke, J. S. Moreland, J. E. Bernhard and S. A. Bass, Phys. Rev. C 96, no.4, 044912 (2017) [arXiv:1610.08490 [nucl-th]].
- [17] S. J. Das, G. Giacalone, P. A. Monard and J. Y. Ollitrault, Phys. Rev. C 97, no.1, 014905 (2018) [arXiv:1708.00081 [nucl-th]].
- [18] W. Busza, K. Rajagopal and W. van der Schee, Ann. Rev. Nucl. Part. Sci. 68, 339-376 (2018) [arXiv:1802.04801 [hep-ph]].
- [19] M. L. Miller, K. Reygers, S. J. Sanders and P. Steinberg, Ann. Rev. Nucl. Part. Sci. 57, 205-243 (2007) [arXiv:nucl-ex/0701025 [nucl-ex]].
- [20] W. Broniowski, M. Rybczynski and P. Bozek, Comput. Phys. Commun. 180, 69-83 (2009) [arXiv:0710.5731 [nucl-th]].
- [21] J. S. Moreland, J. E. Bernhard and S. A. Bass, Phys. Rev. C 92, no.1, 011901 (2015) [arXiv:1412.4708 [nucl-th]].
- [22] C. Loizides, J. Kamin and D. d'Enterria, Phys. Rev. C 97, no.5, 054910 (2018) [erratum: Phys. Rev. C 99, no.1, 019901 (2019)] [arXiv:1710.07098 [nucl-ex]].
- [23] C. Bierlich, G. Gustafson, L. Lönnblad and H. Shah, JHEP 10, 134 (2018) [arXiv:1806.10820 [hep-ph]].
- [24] P. Bożek, W. Broniowski, M. Rybczynski and G. Stefanek, Comput. Phys. Commun. 245, 106850 (2019) [arXiv:1901.04484 [nucl-th]].
- [25] B. Abelev *et al.* [ALICE], Phys. Rev. C 88, no.4, 044909 (2013) [arXiv:1301.4361 [nucl-ex]].
- [26] W. Broniowski and W. Florkowski, Phys. Rev. C 65, 024905 (2002) [arXiv:nucl-th/0110020 [nucl-th]].
- [27] R. Rogly, G. Giacalone and J. Y. Ollitrault, Phys. Rev. C 98, no.2, 024902 (2018) [arXiv:1804.03031 [nucl-th]].
- [28] J. Adam *et al.* [ALICE], Phys. Rev. Lett. **116**, no.22, 222302 (2016) [arXiv:1512.06104 [nucl-ex]].

- [29] J. Adam *et al.* [ALICE], Phys. Lett. B **772**, 567-577 (2017) [arXiv:1612.08966 [nucl-ex]].
- [30] S. Acharya *et al.* [ALICE], Phys. Lett. B **788**, 166-179 (2019) [arXiv:1805.04399 [nucl-ex]].
- [31] P. Di Francesco, M. Guilbaud, M. Luzum and J. Y. Ollitrault, Phys. Rev. C 95, no.4, 044911 (2017) [arXiv:1612.05634 [nucl-th]].
- [32] A. Mazeliauskas, S. Floerchinger, E. Grossi and D. Teaney, Eur. Phys. J. C **79**, no.3, 284 (2019) [arXiv:1809.11049 [nucl-th]].
- [33] B. Abelev et al. [ALICE], Phys. Rev. C 88, 044910 (2013) [arXiv:1303.0737 [hep-ex]].
- [34] C. Pruneau, S. Gavin and S. Voloshin, Phys. Rev. C 66, 044904 (2002) [arXiv:nucl-ex/0204011 [nucl-ex]].
- [35] M. Zhou and J. Jia, Phys. Rev. C 98, no.4, 044903 (2018) [arXiv:1803.01812 [nucl-th]].
- [36] J. Y. Ollitrault, Eur. J. Phys. 29, 275-302 (2008) [arXiv:0708.2433 [nucl-th]].
- [37] E. Abbas *et al.* [ALICE], Phys. Lett. B **726**, 610-622 (2013) [arXiv:1304.0347 [nucl-ex]].
- [38] B. B. Back *et al.* [PHOBOS], Phys. Rev. C 74, 021901 (2006) [arXiv:nucl-ex/0509034 [nucl-ex]].
- [39] F. Gelis, A. M. Stasto and R. Venugopalan, Eur. Phys. J. C 48, 489-500 (2006) [arXiv:hep-ph/0605087 [hep-ph]].
- [40] P. Bozek and W. Broniowski, Phys. Rev. C 97, no.3, 034913 (2018) [arXiv:1711.03325 [nucl-th]].
- [41] J. Jia, C. Zhang and J. Xu, Phys. Rev. Res. 2, no.2, 023319 (2020) [arXiv:2001.08602 [nucl-th]].
- [42] Y. V. Kovchegov and D. E. Wertepny, Nucl. Phys. A 906, 50-83 (2013) [arXiv:1212.1195 [hep-ph]].
- [43] S. Chatrchyan *et al.* [CMS], JHEP **02**, 088 (2014) [arXiv:1312.1845 [nucl-ex]].
- [44] B. Abelev *et al.* [ALICE], Phys. Rev. C 88, no.4, 044909 (2013) [arXiv:1301.4361 [nucl-ex]].
- [45] J. D. Frankland *et al.* [INDRA], [arXiv:2011.04496 [nuclex]].
- [46] P. Parfenov, D. Idrisov, V. B. Luong and A. Taranenko, Particles 4, no.2, 275-287 (2021)
- [47] F. Li, Y. Wang, Z. Gao, P. Li, H. Lv, Q. Li, C. Y. Tsang and M. B. Tsang, [arXiv:2105.08912 [nucl-th]].