Kurtosis of elliptic flow fluctuations

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Elliptic flow (v_2) in ultrarelativistic nucleus-nucleus collisions fluctuates event to event, both in magnitude and in orientation with respect to the reaction plane. Even though the reaction plane is not known event to event in experiment, we show that the statistical properties of v_2 fluctuations in the reaction plane can be precisely extracted from experimental data. Previous studies have shown how to measure the mean, variance and skewness using the first three cumulants $v_2\{2\}$, $v_2\{4\}$ and $v_2\{6\}$. We complement these studies by providing a formula for the kurtosis, which requires an accurate determination of the next cumulant $v_2\{8\}$. Using existing data, we show that the kurtosis is positive for most centralities, in contrast with the kurtosis of triangular flow fluctuations, which is negative. We argue that these features are robust predictions of fluid-dynamical models.

I. INTRODUCTION

Anisotropic flow is a spectacular phenomenon observed in ultrarelativistic nucleus-nucleus [1-4] and protonnucleus [5, 6] collisions. It is understood as resulting from the hydrodynamic response of the quark-gluon plasma to its anisotropic shape at the early stage of the collision [7, 8]. Event to event fluctuations of anisotropic flow [9] thus give valuable insight [10–12] into the earlystage dynamics, where the origin of these fluctuations The probability distribution of anisotropic lies [13]. flow has been analyzed in detail [14]. Flow fluctuations are Gaussian to a good approximation [15]. Non-Gaussianities are however directly revealed by measurements of higher-order cumulants, such as $v_2{4}$ in protonnucleus collisions [16, 17] or v_3 {4} in Pb+Pb collisions [2]. They are also responsible for the small splitting between v_2 {4} and v_2 {6} in Pb+Pb collisions [18-21]. Non-Gaussianities are generic in such microscopic systems, where they appear as corrections to the central limit theorem [22–24]. Unlike the situation in the early Universe, where primordial non-Gaussianities are compatible with zero [25] and observed non-Gaussianities are generated during the expansion, the natural expectation in heavy-ion collisions is that non-Gaussianities are already present in the early stages, and partially washed out by the subsequent hydrodynamic expansion [18, 26]. Precise data on non-Gaussian flow fluctuations allow one to test the hydrodynamic picture [24, 27] and to constrain models of the initial state [12, 28].

We study fluctuations of elliptic flow, v_2 , in semicentral nucleus-nucleus collisions, which is the largest and most accurately measured flow phenomenon [29]. Cumulants [30] of the magnitude of v_2 , denoted by $v_2\{n\}$, have been measured precisely in Pb+Pb collisions for n = 2, 4, 6, 8 [19, 20, 31]. It has long been known [15] that $v_2\{4\}$ is, to a good approximation, the mean v_2 projected onto the reaction plane, while the splitting between $v_2\{2\}$ and $v_2\{4\}$ gives access to the variance of the fluctuations. More recently, it has been shown that the splitting between $v_2\{4\}$ and $v_2\{6\}$ measures the skewness of elliptic flow fluctuations [18]. Here we show that by combining the information from $v_2\{4\}$, $v_2\{6\}$ and $v_2\{8\}$, one can measure the next cumulant, namely, the kurtosis. The kurtosis of v_2 fluctuations in the reaction plane is defined in Sec. II. We estimate its magnitude and centrality dependence. In Sec. III, we derive a general expression of the kurtosis as a function of the measured cumulants, which is valid for a large system. We test the validity of this expression on models of elliptic flow fluctuations. In Sec. IV, we extract the kurtosis from existing data on Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV.

II. KURTOSIS OF v_2 FLUCTUATIONS IN THE REACTION PLANE

A. Definition

Elliptic flow is the second complex Fourier coefficient of the single-particle distribution $f(\mathbf{p})$ [32]: $V_2 = \int e^{2i\varphi} f(\mathbf{p}) d\mathbf{p} / \int f(\mathbf{p}) d\mathbf{p}$, where integration runs over the detector acceptance. It can be decomposed into real and imaginary parts: $V_2 = v_x + iv_y$. In this section, we choose for $\varphi = 0$ the direction of impact parameter, or reaction plane. We focus on the probability distribution of v_x , the projection of elliptic flow onto the reaction plane. This distribution can be characterized by its cumulants. We denote by κ_{n0} the cumulant of order n. The origin of this notation will be clarified in Sec. III. The first 4 cumulants are:

$$\begin{aligned}
\kappa_{10} &= \langle v_x \rangle, \\
\kappa_{20} &= \langle (v_x - \langle v_x \rangle)^2 \rangle, \\
\kappa_{30} &= \langle (v_x - \langle v_x \rangle)^3 \rangle, \\
\kappa_{40} &= \langle (v_x - \langle v_x \rangle)^4 \rangle - 3\kappa_{20}^2,
\end{aligned}$$
(1)

where angular brackets denote an average over many events in a centrality class. κ_{30} and κ_{40} vanish if the distribution of v_x is a Gaussian [15].

The standardized skewness of the distribution of v_x is

$$\gamma_1 \equiv \frac{\kappa_{30}}{(\kappa_{20})^{3/2}}.\tag{2}$$

1.4

1.2

1.0

0.8

0.6

0.4

0.2

0.0

kurtosis

A generic prediction of hydrodynamics is that the distribution of v_x has negative skew, $\gamma_1 < 0$ [18]. The reason is twofold: First, v_x is proportional to the eccentricity in the reaction plane to a good approximation [33, 34]. Second, the eccentricity is bounded by 1, and this right cutoff skews the distribution to the left. γ_1 has been predicted to become more negative as the centrality percentile increases [18]. Recent experimental analyses [19, 20] confirm the hydrodynamic prediction: γ_1 reaches the value -0.4 at 60% centrality.

Our goal in this paper is to extend this analysis to the next cumulant order. The standardized kurtosis of the fluctuations of v_x is defined by

$$\gamma_2 \equiv \frac{\kappa_{40}}{(\kappa_{20})^2}.\tag{3}$$

This quantity, which vanishes if the fluctuations of v_x are Gaussian, is sometimes referred to as "excess kurtosis" rather than just "kurtosis". A positive γ_2 indicates that the distribution has heavier tails than a Gaussian distribution.

B. Magnitude and centrality dependence

We now investigate the order of magnitude and centrality dependence of the kurtosis γ_2 . Abbasi *et al.* [26] have carried out an extensive event-by-event hydrodynamic calculation (14000 events per centrality bin) using Monte Carlo Glauber initial conditions [35]. Their result for the standardized kurtosis γ_2 (Fig. 3(c) arXiv version v1 of Ref. [26]) is essentially compatible with 0 in the 0-60% centrality range. They also compute the kurtosis of the initial eccentricity, which is obtained by replacing the elliptic flow $V_2 = v_x + iv_y$ with the initial eccentricity $\varepsilon_2 = \varepsilon_x + i\varepsilon_y$ in Eq. (1). It is compatible with 0 for central collisions, but increases with the centrality percentile and reaches 0.4 at 60% centrality. If V_2 was proportional to ε_2 in every event [33], these two quantities would have the exact same γ_2 . The fact that γ_2 is smaller for elliptic flow than for the initial eccentricity is a clear signature of a nonlinear hydrodynamic response. A similar phenomenon is observed for the skewness γ_1 , which is reduced by a factor $\simeq 2$ through the hydrodynamic evolution [18, 26]. Interestingly, this reduction is not seen in transport calculations using the AMPT model [36], where the skewness of elliptic flow is compatible with the skewness of the initial eccentricity.

Even though the hydrodynamic evolution washes out part of the initial non-Gaussianity, it is likely that some of it will remain, which is hidden by statistical errors in the calculation of Ref. [26]. To gain an understanding of the behavior of the kurtosis with the centrality

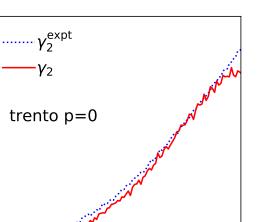


FIG. 1. (Color online) Kurtosis of eccentricity fluctuations along the reaction plane in the T_RENTo model for Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV, as a function of centrality percentile. Full line: Eq. (3). Dotted line: Eq. (15). In these equations, we replace the elliptic flow with the initial eccentricity.

20

10

30

centrality [%]

40

50

60

percentile and of its sensitivity to models of initial conditions, we evaluate it within the T_RENTo model of initial conditions [37]. The T_RENTo model has a parameter pwhich determines how the initial energy density depends on the thickness functions T_A and T_B of colliding nuclei. We choose the value p = 0, which corresponds to a density proportional to $\sqrt{T_A T_B}$. This parametrization reproduces quantitatively the magnitude of anisotropic flow fluctuations in Pb+Pb collisions [12]. We generate 8×10^7 Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV, which we sort into 0.5% centrality bins. The full line in Fig. 1 displays γ_2 of initial eccentricity fluctuations as a function of the centrality percentile. It is significantly larger than with the Glauber initial conditions of Ref. [26], and reaches unity at 60% centrality.

An interesting feature shown by γ_2 in Fig. 1 is its clear change of sign between 10% and 15% centrality. The negative value of the kurtosis in the 0-10% centrality range means that the distribution of ε_x has lighter tails than a Gaussian. In the context of eccentricity fluctuations, this has been shown to be a consequence of the bound $|\varepsilon_x| < 1$ when the distribution of $(\varepsilon_x, \varepsilon_y)$ is azimuthallysymmetric [24], which is the case precisely for the most central collisions. For this reason, one expects the kurtosis of the initial triangularity, ε_3 , and eventually that of triangular flow, v_3 , to be negative. The kurtosis of triangular flow fluctuations can be directly obtained from experimental data [26]

$$\gamma_2 = -\frac{3}{2} \frac{v_3\{4\}^4}{v_3\{2\}^4},\tag{4}$$

where the factor $3/2 = \langle \cos^4 \phi \rangle / \langle \cos^2 \phi \rangle^2$ comes from the projection onto the *x* axis. Thus the observation of a positive $v_3\{4\}$ in Pb+Pb collisions [2] and Xe+Xe collisions [38], in agreement with hydrodynamic predictions [39], implies a negative γ_2 , which lies typically between -0.1 and 0 [26].

We now assess the robustness of the T_RENTo results in Fig. 1 by evaluating the kurtosis for the Elliptic Power distribution [40]. This distribution is the exact [28] distribution of the complex eccentricity $(\varepsilon_x, \varepsilon_y)$ for $N \ge 2$ identical, pointlike sources, randomly distributed in the (x, y) plane with a Gaussian probability distribution. Its analytic form is

$$p(\varepsilon_x, \varepsilon_y) = \frac{N-2}{2\pi} (1-\varepsilon_0^2)^{\frac{N-1}{2}} \frac{(1-\varepsilon_x^2 - \varepsilon_y^2)^{\frac{N}{2}-2}}{(1-\varepsilon_0 \varepsilon_x)^{N-1}}, \quad (5)$$

where the parameter ε_0 is the eccentricity of the distribution of the sources. Eq. (5) provides a reasonable fit of most models of initial conditions for all centralities. The cumulants κ_{n0} of this distribution can be evaluated analytically in terms of hypergeometric functions [26]. Their asymptotic values for large N are:

$$\kappa_{10} = \varepsilon_0 + \mathcal{O}\left(\frac{1}{N}\right), \kappa_{20} = \frac{(1 - \varepsilon_0^2)^2}{N} + \mathcal{O}\left(\frac{1}{N^2}\right), \kappa_{30} = -\frac{6\varepsilon_0(1 - \varepsilon_0^2)^3}{N^2} + \mathcal{O}\left(\frac{1}{N^3}\right), \kappa_{40} = \frac{6(12\varepsilon_0^2 - 1)(1 - \varepsilon_0^2)^4}{N^3} + \mathcal{O}\left(\frac{1}{N^4}\right).$$
(6)

The cumulants decrease by successive powers of 1/N as the order increases. The skewness and kurtosis are given by:

$$\gamma_1 = -\frac{6\varepsilon_0}{\sqrt{N}} + \mathcal{O}\left(\frac{1}{N^{3/2}}\right),$$

$$\gamma_2 = \frac{6\left(12\varepsilon_0^2 - 1\right)}{N} + \mathcal{O}\left(\frac{1}{N^2}\right). \tag{7}$$

They vanish in the limit $N \to \infty$, as expected from the central limit theorem. The skewness γ_1 is negative. The sign of γ_2 is driven by the mean eccentricity ε_0 : it is negative for $\varepsilon_0 < 0.28$ and positive for $\varepsilon_0 > 0.29$. Therefore, the observation that γ_2 is larger in our T_RENTo calculation than in the Glauber model [26] seems naturally explained by the fact that the T_RENTo model presents a larger eccentricity in the reaction plane [41].

In order to check the validity of the asymptotic result (7), we evaluate γ_2 numerically for realistic values of N. Figure 2 (a) displays contour plots of γ_2 in the (N, ε_0) plane. In order to compare with the values of γ_2 from the

T_RENTo calculation, displayed in Fig. 1, we have fitted the eccentricity distributions from the T_RENTo model with Eq. (5) in a few centrality intervals. The resulting values of (N, ε_0) are displayed as symbols in Fig. 2. The values of γ_2 in Figs. 1 and 2 (a) are not in quantitative agreement, which means that the fit of T_RENTo with the Elliptic Power distribution is not perfect. However, the order of magnitude and change of sign are reproduced, and seem to be robust predictions of hydrodynamics.

In conclusion, the positive sign of the kurtosis in noncentral collisions appears to be a generic consequence of a large eccentricity in the reaction plane.

III. EXPRESSING THE KURTOSIS AS A FUNCTION OF MEASURED CUMULANTS

We now explain how the kurtosis γ_2 can be extracted from experiment. This is not trivial because experiments measure the magnitude $v_2 \equiv \sqrt{v_x^2 + v_y^2}$, not v_x and v_y separately. In this Section, we show that one can however reconstruct γ_2 for a large system.¹ We first define the cumulants in two coordinate systems: (1) The reaction plane coordinate system, where x is the direction of impact parameter, or reaction plane, as in the previous section. (2) The detector system, where x denotes a fixed orientation with respect to the detector. This is the natural coordinate system for experiments. The usual cumulant of order n measured in experiments, denoted by $v_2\{n\}$ [30], is defined in this system.

A. Cumulants in the reaction plane coordinate system

Cumulants are a double sequence κ_{n_x,n_y} with $n_x, n_y \ge 0$, which completely specify the probability distribution of (v_x, v_y) . They are defined by [26]:

$$\ln\left\langle e^{k_{x}v_{x}+k_{y}v_{y}}\right\rangle = \sum_{n_{x},n_{y}}\frac{k_{x}^{n_{x}}}{n_{x}!}\frac{k_{y}^{n_{y}}}{n_{y}!}\kappa_{n_{x},n_{y}}.$$
 (8)

The probability distribution of (v_x, v_y) is symmetric with respect to the reaction plane in the absence of parity violation [42, 43]. This implies that it is an even function of v_y , so that the only nonvanishing cumulants are those with even n_y . Let us write the expansion explicitly, keeping all terms with $n_x + n_y \leq 4$:

$$\ln \left\langle e^{k_x v_x + k_y v_y} \right\rangle = k_x \kappa_{10} \\ + \frac{k_x^2}{2} \kappa_{20} + \frac{k_y^2}{2} \kappa_{02}$$

¹ By large system, we mean a nucleus-nucleus collision, and we have tested that our formalism works up to 60% centrality.

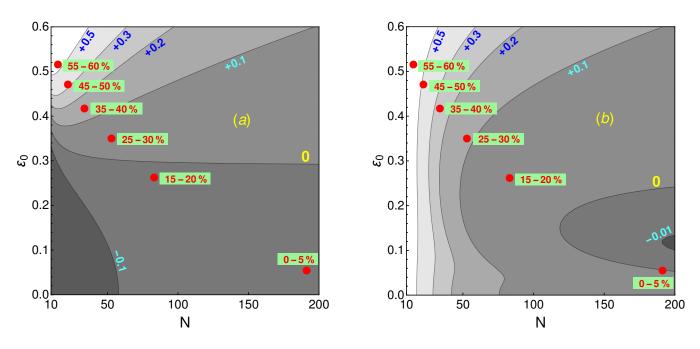


FIG. 2. (Color online) Contour plots of the standardized kurtosis, γ_2 , defined in Eq. (3) [panel (a)], and of the estimated kurtosis, γ_2^{expt} , defined by Eq. (15) [panel (b)], computed using the Elliptic Power distribution. The quantities ε_0 and N are the eccentricity of the sources, and their number, respectively, as in Eq. (5). Symbols correspond to Elliptic Power fits to the T_RENTo calculations (see Fig. 1) in a few centrality windows.

$$+ \frac{k_x^3}{6}\kappa_{30} + \frac{k_xk_y^2}{2}\kappa_{12} \\ + \frac{k_x^4}{24}\kappa_{40} + \frac{k_y^4}{24}\kappa_{04} + \frac{k_x^2k_y^2}{4}\kappa_{22}.$$
 (9)

The four lines in the right-hand side correspond respectively to the mean, variance, skewness and kurtosis of the distribution of (v_x, v_y) . The cumulants κ_{n0} coincide with those defined in Eq. (1).

B. Measured cumulants

We now recall the definition of the measured cumulants $v_2\{n\}$. The only difference with the previously defined cumulants is the coordinate system. We start again from the generating function Eq. (8), where the x axis now denotes a fixed direction in the detector. We first evaluate this generating function in the simple case where v_2 is the same for all events, but the orientation of the reaction plane Φ_R is random. We write $v_x = v_2 \cos 2\Phi_R$, $v_y = v_2 \sin 2\Phi_R$. Averaging over events amounts to averaging over Φ_R if the detector is azimuthally symmetric:

$$\ln\left\langle e^{k_x v_x + k_y v_y} \right\rangle = \ln I_0(kv_2) = \sum_{n=2}^{\infty} c_n k^n (v_2)^n, \quad (10)$$

where I_0 is the modified Bessel function of the first kind, $k \equiv \sqrt{k_x^2 + k_y^2}$, and c_n are rational coefficients which vanish for odd n: $c_2 = 1/4$, $c_4 = -1/64$, $c_6 = 1/576$, $c_8 = -11/49152...$ In the general case where v_2 fluctuates event to event, $\langle e^{k_x v_x + k_y v_y} \rangle$ depends only on k^2 by azimuthal symmetry. One can write, without any loss of generality,

$$\ln \left\langle e^{k_x v_x + k_y v_y} \right\rangle = \ln \langle I_0(k v_2) \rangle = \sum_{n=2}^{\infty} c_n k^n v_2 \{n\}^n, \quad (11)$$

where the coefficients c_n are defined by Eq. (10). Eq. (11) defines $v_2\{n\}^n$ for all even n. Comparison with Eq. (10) shows that if v_2 is the same for all events, then, $v_2\{n\}$ coincides with v_2 for all n. However, despite this (perhaps unfortunate) notation, $v_2\{n\}^n$ defined by Eq. (11) can have positive or negative sign if v_2 fluctuates event to event.

C. From one coordinate system to the other

One can express $v_2\{n\}^n$ as a function of the $\kappa_{n_x n_y}$ in Eq. (8) systematically in the following way:

- Write $k_x = k \cos \theta$, $k_y = k \sin \theta$ in Eq. (8).
- Exponentiate Eq. (8) and average over θ .
- Take the logarithm, expand in powers of k, and match the result order by order to the right-hand side of Eq. (11).

As an illustration, one finds the following exact expressions for the first three cumulants [26]:

$$v_2\{2\}^2 = \kappa_{10}^2$$

$$\begin{aligned} & +\kappa_{20} + \kappa_{02}, \\ v_{2}\{4\}^{4} &= \kappa_{10}^{4} \\ & +2\kappa_{10}^{2}(\kappa_{02} - \kappa_{20}) \\ & -4\kappa_{10}(\kappa_{30} + \kappa_{12}) - (\kappa_{20} - \kappa_{02})^{2} \\ & -(\kappa_{04} + \kappa_{40} + 2\kappa_{22}), \\ v_{2}\{6\}^{6} &= \kappa_{10}^{6} \\ & +3\kappa_{10}^{4}(\kappa_{02} - \kappa_{20}) \\ & -2\kappa_{10}^{3}(2\kappa_{30} + 3\kappa_{12}) \\ & +\frac{3}{2}\kappa_{10}^{2}(\kappa_{40} - \kappa_{04}) - 6\kappa_{10}\kappa_{30}(\kappa_{02} - \kappa_{20}) \\ & +\frac{3}{2}(\kappa_{04} - \kappa_{40})(\kappa_{02} - \kappa_{20}) + \frac{5}{2}\kappa_{30}^{2} + 3\kappa_{30}\kappa_{12} \\ & +\frac{9}{2}\kappa_{12}^{2} + \frac{3}{2}\kappa_{10}(\kappa_{50} + \kappa_{14} + 2\kappa_{32}) \\ & +\frac{3}{4}(\kappa_{24} + \kappa_{42}) + \frac{\kappa_{60}}{4} + \frac{\kappa_{06}}{4}. \end{aligned}$$
(12)

One cannot invert these relations and reconstruct the double sequence $\kappa_{n_xn_y}$ from the single sequence $v_2\{n\}$. Simplifications occur, however, for a large system. Inspired by the Elliptic Power distribution, where $\kappa_{n_xn_y}$ is of order $N^{1-n_x-n_y}$, we expand Eqs. (12) in powers of 1/N. Doing so, one finds that, to leading order in 1/N, the differences between successive cumulants only involve the single sequence κ_{n0} :

$$v_{2}\{2\}^{2} - v_{2}\{4\}^{2} = 2\kappa_{20} + \mathcal{O}\left(\frac{1}{N^{2}}\right),$$

$$v_{2}\{4\}^{3} - v_{2}\{6\}^{3} = -\kappa_{30} + \mathcal{O}\left(\frac{1}{N^{3}}\right),$$

$$v_{2}\{4\}^{4} - 12v_{2}\{6\}^{4} + 11v_{2}\{8\}^{4} = -\frac{8}{3}\kappa_{40} + \mathcal{O}\left(\frac{1}{N^{4}}\right).$$
(13)

Thus the splitting between order 2 and order 4 is due to elliptic flow fluctuations, as has long been known [44]. The splitting between 4 and 6 is the skewness [18]. The new result is the last line of Eq. (13), which shows that the kurtosis κ_{40} can be extracted by combining orders 4, 6, and 8, in a way that eliminates the contribution from the skewness κ_{30} .

The observation that to leading order in 1/N, the splittings between successive cumulants only involve κ_{n0} , that is, fluctuations projected onto the reaction plane, can be understood as follows. Elliptic flow in a given event can be decomposed as $v_2 = \sqrt{(\kappa_{10} + \delta_x)^2 + \delta_y^2}$, where κ_{10} is the mean elliptic flow in the reaction plane and δ_x and δ_y denote the fluctuations in the directions parallel and perpendicular to the reaction plane. Expanding in powers of the fluctuation, one finds to leading order

$$v_2 \simeq \kappa_{10} + \delta_x + \frac{\delta_y^2}{2\kappa_{10}}.$$
(14)

Thus, fluctuations along the y direction enter at higher order than fluctuations along the x direction.

If one keeps only the leading term in the right-hand side of Eqs. (13), one obtains approximate expressions of the standardized skewness and kurtosis of v_x fluctuations, defined by Eqs. (2) and (3), in terms of measured quantities:

$$\gamma_{1} \simeq \gamma_{1}^{\text{expt}} \equiv -2^{3/2} \frac{v_{2}\{4\}^{3} - v_{2}\{6\}^{3}}{\left(v_{2}\{2\}^{2} - v_{2}\{4\}^{2}\right)^{3/2}},$$

$$\gamma_{2} \simeq \gamma_{2}^{\text{expt}} \equiv -\frac{3}{2} \frac{v_{2}\{4\}^{4} - 12v_{2}\{6\}^{4} + 11v_{2}\{8\}^{4}}{\left(v_{2}\{2\}^{2} - v_{2}\{4\}^{2}\right)^{2}}.$$
(15)

The expression of the skewness is essentially equivalent to that proposed in Ref. [18], but slightly simpler. The expression of the kurtosis is new, but similar expressions have been derived by considering fluctuations of the flow magnitude alone [45]. γ_2^{expt} vanishes if

$$v_2\{6\} - v_2\{8\} = \frac{1}{11} (v_2\{4\} - v_2\{6\}),$$
 (16)

where we have linearized Eq. (15) by taking into account the observation that the splittings are very small in practice. The ALICE collaboration [20] has found that data satisfy Eq. (16) within error bars. Measuring a nontrivial value of the kurtosis requires very precise data, as we shall see in Sec. IV.

D. Numerical tests

We now check that Eq. (15) provides a reasonable approximation of the kurtosis for heavy-ion collisions. We first use the TRENTo simulation. The dotted line in Fig. 1 is γ_2^{expt} defined by Eq. (15). Agreement with γ_2 is good for all centralities. γ_2^{expt} is however systematically larger than γ_2 by a few 10^{-2} . In order to confirm that these trends are generic, we use the Elliptic Power distribution. Fig. 2 (b) displays contour plots of γ_2^{expt} in the (N, ε_0) plane. Comparing with panel (a), one immediately notices the differences with γ_2 . These differences are large, even for large N, when ε_0 is small. The reason is that the expansion scheme used to derive Eq. (15)treats fluctuations as small corrections to the mean eccentricity, and this approximation breaks down for small ε_0 . γ_2^{expt} is nevertheless a good approximation of γ_2 as soon as N is large (which is everywhere the case in Fig. 2) and ε_0 is not too small. This is the case in particular in semi-central nucleus-nucleus collisions, where one expects to see a positive kurtosis.

IV. EXPERIMENTAL DATA

Finally, we estimate the kurtosis in Pb+Pb collisions by applying Eq. (15) to CMS data [19]. The result is displayed in Fig. 3. Error bars have been estimated by assuming that the errors on the ratios $v_2\{8\}/v_2\{6\}$ and $v_2\{6\}/v_2\{4\}$ are independent. Despite the large error bars, there is clear evidence that the kurtosis is positive above 30% centrality, in agreement with the theoretical calculations of Figs. 1 and 2. We have also extracted the

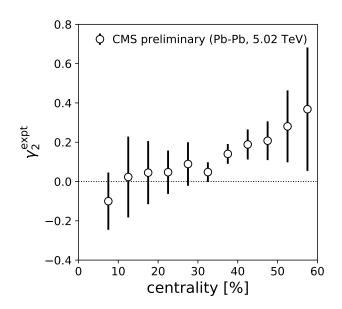


FIG. 3. (Color online) Kurtosis of v_2 fluctuations in Pb+Pb collisions at $\sqrt{s_{\rm NN}} = 5.02$ TeV estimated using Eq. (15) and CMS data [19].

kurtosis from ALICE data [20]. The result (not shown) is compatible with CMS data, but with much larger error bars, and does not show any evidence of a non-zero kurtosis. Note that the results of Abbasi *et al.* [26] for the kurtosis of Monte Carlo Glauber initial conditions are of the same magnitude as the experimental data in Fig. 3. However, this kurtosis is washed out by the hydrodynamic evolution, which implies that the kurtosis of the Glauber model is too small.

We deem that hydrodynamic calculations with extensive statistics, like those of Ref. [26], should be carried out with initial conditions displaying a larger initial kurtosis, such as the T_RENTo p = 0 model shown in Fig. 1.

V. CONCLUSIONS

We have argued that the kurtosis of elliptic flow fluctuations along the direction of impact parameter is positive in non-central Pb+Pb collisions. This is a notable difference with respect to the kurtosis of triangular flow fluctuations, which is observed to be negative in experimental data, as predicted by hydrodynamic calculations.

We have provided a formula to extract the kurtosis of elliptic flow fluctuations from high-precision measurements of $v_2\{2\}$, $v_2\{4\}$, $v_2\{6\}$ and $v_2\{8\}$. This requires very high statistics of events, and, at the present moment, only preliminary Run2 CMS data allow for the extraction of a significant γ_2 , which indeed turns out to be positive above 30% centrality. The outcome of investigating flow fluctuations at such a detailed level is nevertheless evident already after our crude extraction: Combined with the precise calculations of Ref. [26], our result provides clear indication that only models displaying large eccentricity, and these are typically the models inspired by high-energy QCD [46–49], have the potential of yielding a γ_2 of order 0.5 after the hydrodynamic evolution.

The bottom line is that with great precision comes great discriminating power, although at the cost of increasing the statistics of hydrodynamic calculations. With the advent of LHC3 data, we expect future dedicated analyses using robust methods [50] to characterize the details of flow fluctuations, such as the kurtosis studied in this paper, with unprecedented accuracy. This will lead to novel insightful tests of the hydrodynamic picture, in the nontrivial regime where the hydrodynamic response driving anisotropic flow is strongly nonlinear.

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- K. H. Ackermann *et al.* [STAR Collaboration], Phys. Rev. Lett. **86**, 402 (2001) doi:10.1103/PhysRevLett.86.402 [nucl-ex/0009011].
- et[ALICE Aamodt al.Collabora-[2] K. Rev. Lett. tion], Phys. 107, 032301 (2011)doi:10.1103/PhysRevLett.107.032301 [arXiv:1105.3865 [nucl-ex]].
- [3] A. Adare *et al.* [PHENIX Collaboration], Phys. Rev. Lett. **107**, 252301 (2011) doi:10.1103/PhysRevLett.107.252301 [arXiv:1105.3928

[nucl-ex]].

- [4] C. Aidala *et al.* [PHENIX Collaboration], arXiv:1805.02973 [nucl-ex].
- [5] S. Chatrchyan *et al.* [CMS Collaboration], Phys. Lett. B **718**, 795 (2013) doi:10.1016/j.physletb.2012.11.025 [arXiv:1210.5482 [nucl-ex]].
- [6] G. Aad *et al.* [ATLAS Collaboration], Phys. Rev. Lett. **110**, no. 18, 182302 (2013) doi:10.1103/PhysRevLett.110.182302 [arXiv:1212.5198 [hep-ex]].

- J. Y. Ollitrault, Phys. Rev. D 46, 229 (1992). doi:10.1103/PhysRevD.46.229
- [8] B. Alver and G. Roland, Phys. Rev. C 81, 054905 (2010) Erratum: [Phys. Rev. C 82, 039903 (2010)] doi:10.1103/PhysRevC.82.039903, 10.1103/Phys-RevC.81.054905 [arXiv:1003.0194 [nucl-th]].
- [9] B. Alver *et al.* [PHOBOS Collaboration], Phys. Rev. Lett. **98**, 242302 (2007) doi:10.1103/PhysRevLett.98.242302 [nucl-ex/0610037].
- [10] R. S. Bhalerao, M. Luzum and J. Y. Ollitrault, Phys. Rev. C 84, 034910 (2011) doi:10.1103/PhysRevC.84.034910 [arXiv:1104.4740 [nucl-th]].
- [11] E. Retinskaya, M. Luzum and J. Y. Ollitrault, Phys. Rev. C 89, no. 1, 014902 (2014) doi:10.1103/PhysRevC.89.014902 [arXiv:1311.5339 [nucl-th]].
- [12] G. Giacalone, J. Noronha-Hostler and J. Y. Ollitrault, Phys. Rev. C 95, no. 5, 054910 (2017) doi:10.1103/PhysRevC.95.054910 [arXiv:1702.01730 [nucl-th]].
- [13] J. L. Albacete, P. Guerrero-Rodrguez and C. Marquet, arXiv:1808.00795 [hep-ph].
- [14] G. Aad *et al.* [ATLAS Collaboration], JHEP **1311**, 183
 (2013) doi:10.1007/JHEP11(2013)183 [arXiv:1305.2942
 [hep-ex]].
- [15] S. A. Voloshin, A. M. Poskanzer, A. Tang and G. Wang, Phys. Lett. B 659, 537 (2008) doi:10.1016/j.physletb.2007.11.043 [arXiv:0708.0800 [nucl-th]].
- [16] G. Aad *et al.* [ATLAS Collaboration], Phys. Lett.
 B **725**, 60 (2013) doi:10.1016/j.physletb.2013.06.057
 [arXiv:1303.2084 [hep-ex]].
- [17] S. Chatrchyan *et al.* [CMS Collaboration], Phys. Lett. B **724**, 213 (2013) doi:10.1016/j.physletb.2013.06.028
 [arXiv:1305.0609 [nucl-ex]].
- [18] G. Giacalone, L. Yan, J. Noronha-Hostler and J. Y. Ollitrault, Phys. Rev. C 95, no. 1, 014913 (2017) doi:10.1103/PhysRevC.95.014913 [arXiv:1608.01823 [nucl-th]].
- [19] A. M. Sirunyan *et al.* [CMS Collaboration], arXiv:1711.05594 [nucl-ex].
- [20] S. Acharya *et al.* [ALICE Collaboration], JHEP
 1807, 103 (2018) doi:10.1007/JHEP07(2018)103
 [arXiv:1804.02944 [nucl-ex]].
- [21] H. Mehrabpour and S. F. Taghavi, arXiv:1805.04695 [nucl-th].
- [22] B. Alver *et al.* [PHOBOS Collaboration], Phys. Rev. C 77, 014906 (2008) doi:10.1103/PhysRevC.77.014906 [arXiv:0711.3724 [nucl-ex]].
- [23] R. S. Bhalerao, M. Luzum and J. Y. Ollitrault, Phys. Rev. C 84, 054901 (2011) doi:10.1103/PhysRevC.84.054901 [arXiv:1107.5485 [nucl-th]].
- [24] L. Yan and J. Y. Ollitrault, Phys. Rev. Lett. 112, 082301 (2014) doi:10.1103/PhysRevLett.112.082301 [arXiv:1312.6555 [nucl-th]].
- [25] P. A. R. Ade *et al.* [Planck Collaboration], Astron. Astrophys. **571**, A24 (2014) doi:10.1051/0004-6361/201321554 [arXiv:1303.5084 [astro-ph.CO]].
- [26] N. Abbasi, D. Allahbakhshi, A. Davody and S. F. Taghavi, Phys. Rev. C 98, no. 2, 024906 (2018) doi:10.1103/PhysRevC.98.024906 [arXiv:1704.06295 [nucl-th]].

- [27] V. Khachatryan *et al.* [CMS Collaboration], Phys. Rev. Lett. **115**, no. 1, 012301 (2015) doi:10.1103/PhysRevLett.115.012301 [arXiv:1502.05382 [nucl-ex]].
- [28] H. Grönqvist, J. P. Blaizot and J. Y. Ollitrault, Phys. Rev. C 94, no. 3, 034905 (2016) doi:10.1103/PhysRevC.94.034905 [arXiv:1604.07230 [nucl-th]].
- [29] K. Aamodt et al. [ALICE Collaboration], Phys. Rev. Lett. **105**, 252302 (2010) doi:10.1103/PhysRevLett.105.252302 [arXiv:1011.3914 [nucl-ex]].
- [30] N. Borghini, P. M. Dinh and J. Y. Ollitrault, Phys. Rev. C 64, 054901 (2001) doi:10.1103/PhysRevC.64.054901 [nucl-th/0105040].
- [31] G. Aad *et al.* [ATLAS Collaboration], Eur. Phys. J. C 74, no. 11, 3157 (2014) doi:10.1140/epjc/s10052-014-3157-z [arXiv:1408.4342 [hep-ex]].
- [32] M. Luzum, J. Phys. G 38, 124026 (2011) doi:10.1088/0954-3899/38/12/124026 [arXiv:1107.0592 [nucl-th]].
- [33] F. G. Gardim, F. Grassi, M. Luzum and J. Y. Ollitrault, Phys. Rev. C 85, 024908 (2012) doi:10.1103/PhysRevC.85.024908 [arXiv:1111.6538 [nucl-th]].
- [34] H. Niemi, G. S. Denicol, H. Holopainen and P. Huovinen, Phys. Rev. C 87, no. 5, 054901 (2013) doi:10.1103/PhysRevC.87.054901 [arXiv:1212.1008 [nucl-th]].
- [35] M. L. Miller, K. Reygers, S. J. Sanders and P. Steinberg, Ann. Rev. Nucl. Part. Sci. 57, 205 (2007) doi:10.1146/annurev.nucl.57.090506.123020 [nuclex/0701025].
- [36] D. X. Wei, X. G. Huang and L. Yan, arXiv:1807.06299 [nucl-th].
- [37] J. S. Moreland, J. E. Bernhard and S. A. Bass, Phys. Rev. C 92, no. 1, 011901 (2015) doi:10.1103/PhysRevC.92.011901 [arXiv:1412.4708 [nucl-th]].
- [38] G. Giacalone, J. Noronha-Hostler, M. Luzum and J. Y. Ollitrault, arXiv:1807.05557 [nucl-th].
- [39] G. Giacalone, J. Noronha-Hostler, M. Luzum and J. Y. Ollitrault, Phys. Rev. C 97, no. 3, 034904 (2018) doi:10.1103/PhysRevC.97.034904 [arXiv:1711.08499 [nucl-th]].
- [40] L. Yan, J. Y. Ollitrault and A. M. Poskanzer, Phys. Rev. C 90, no. 2, 024903 (2014) doi:10.1103/PhysRevC.90.024903 [arXiv:1405.6595 [nucl-th]].
- [41] L. Yan, J. Y. Ollitrault and A. M. Poskanzer, Phys. Lett.
 B 742, 290 (2015) doi:10.1016/j.physletb.2015.01.039
 [arXiv:1408.0921 [nucl-th]].
- [42] D. Kharzeev, Phys. Lett. B 633, 260 (2006) doi:10.1016/j.physletb.2005.11.075 [hep-ph/0406125].
- [43] S. A. Voloshin, Phys. Rev. C 70, 057901 (2004) doi:10.1103/PhysRevC.70.057901 [hep-ph/0406311].
- [44] R. S. Bhalerao and J. Y. Ollitrault, Phys. Lett. B 641, 260 (2006) doi:10.1016/j.physletb.2006.08.055 [nuclth/0607009].
- [45] J. Jia and S. Radhakrishnan, Phys. Rev. C 92, no. 2, 024911 (2015) doi:10.1103/PhysRevC.92.024911 [arXiv:1412.4759 [nucl-ex]].
- [46] T. Hirano, U. W. Heinz, D. Kharzeev, R. Lacey and Y. Nara, Phys. Lett. B 636, 299 (2006)

doi:10.1016/j.physletb.2006.03.060 [nucl-th/0511046].

- [47] B. Schenke, P. Tribedy and R. Venugopalan, Phys. Rev. Lett. **108**, 252301 (2012) doi:10.1103/PhysRevLett.108.252301 [arXiv:1202.6646 [nucl-th]].
- [48] H. Niemi, K. J. Eskola and R. Paatelainen, Phys. Rev. C **93**, no. 2, 024907 (2016)

doi:10.1103/PhysRevC.93.024907 [arXiv:1505.02677 [hep-ph]].

- [49] J. L. Nagle and W. A. Zajc, arXiv:1808.01276 [nucl-th].
- [50] P. Di Francesco, M. Guilbaud, M. Luzum and J. Y. Ollitrault, Phys. Rev. C 95, no. 4, 044911 (2017) doi:10.1103/PhysRevC.95.044911 [arXiv:1612.05634 [nucl-th]].